

# **Computer Vision - Lecture 18**

#### Camera Calibration & 3D Reconstruction

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#### **Course Outline**

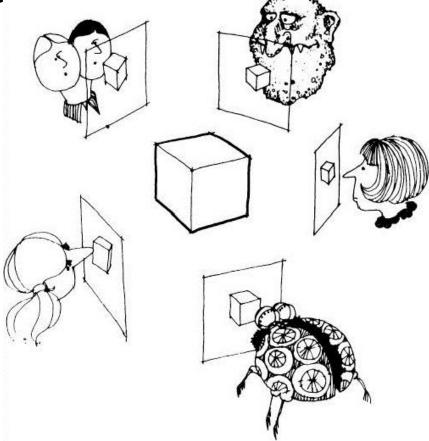
- Image Processing Basics
- Segmentation & Grouping
- Object Recognition
- Local Features & Matching
- Object Categorization
- 3D Reconstruction
  - Epipolar Geometry and Stereo Basics
  - Camera calibration & Uncalibrated Reconstruction
  - Structure-from-Motion
- Motion and Tracking



### Recap: What Is Stereo Vision?

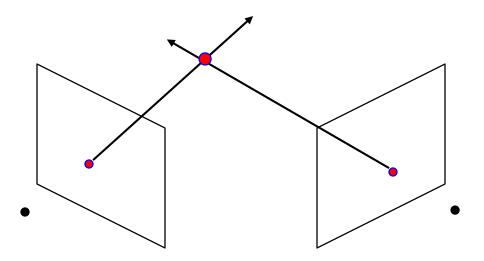
 Generic problem formulation: given several images of the same object or scene, compute a representation of







### Recap: Depth with Stereo - Basic Idea



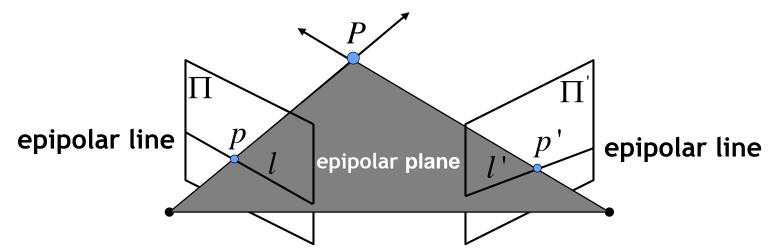
- Basic Principle: Triangulation
  - Gives reconstruction as intersection of two rays
  - Requires
    - Camera pose (calibration)
    - Point correspondence

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### Recap: Epipolar Geometry

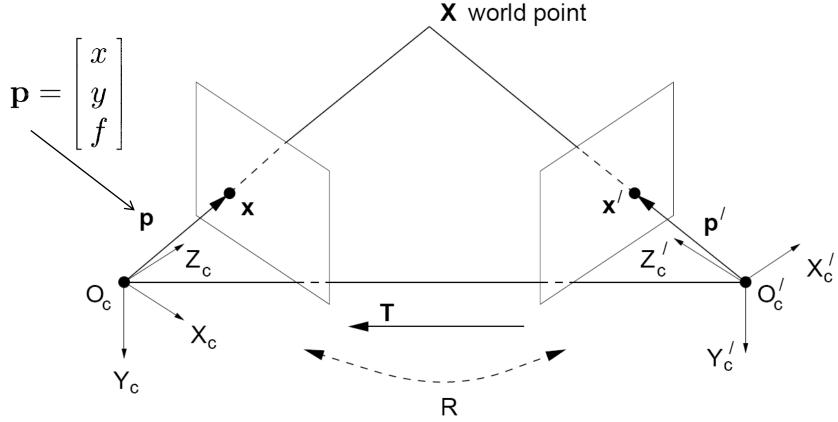
 Geometry of two views allows us to constrain where the corresponding pixel for some image point in the first view must occur in the second view.



- Epipolar constraint:
  - > Correspondence for point p in  $\Pi$  must lie on the epipolar line l in  $\Pi$  (and vice versa).
  - Reduces correspondence problem to 1D search along conjugate epipolar lines.



### Recap: Stereo Geometry With Calibrated Cameras



 Camera-centered coordinate systems are related by known rotation R and translation T:

$$\mathbf{X}' = \mathbf{R}\mathbf{X} + \mathbf{T}$$





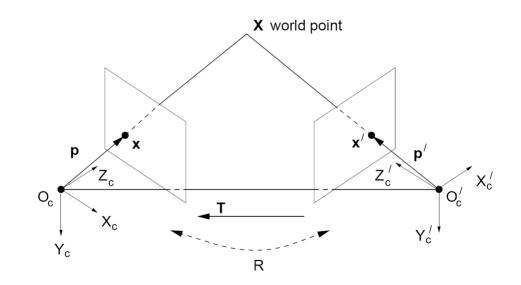
### Recap: Essential Matrix

$$\mathbf{X}' \cdot (\mathbf{T} \times \mathbf{R} \mathbf{X}) = 0$$

$$\mathbf{X}' \cdot \left(\mathbf{T}_x \ \mathbf{R}\mathbf{X}\right) = 0$$

Let 
$$\mathbf{E} = \mathbf{T}_x \mathbf{R}$$

$$\mathbf{X}'^T \mathbf{E} \mathbf{X} = \mathbf{0}$$



 This holds for the rays p and p' that are parallel to the camera-centered position vectors X and X', so we have:

$$\mathbf{p'}^T \mathbf{E} \mathbf{p} = 0$$

 E is called the essential matrix, which relates corresponding image points [Longuet-Higgins 1981]

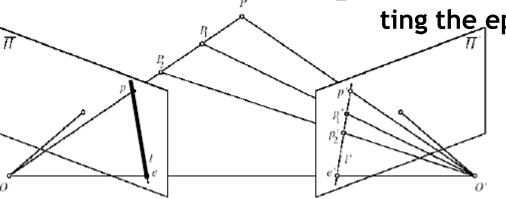


### **Essential Matrix and Epipolar Lines**

$$\mathbf{p'}^{\mathrm{T}}\mathbf{E}\mathbf{p} = 0$$

Epipolar constraint: if we observe point p in one image, then its position p in second image must satisfy this equation.

 $m{l}' = m{Ep}$  is the coordinate vector representing the epipolar line for point p

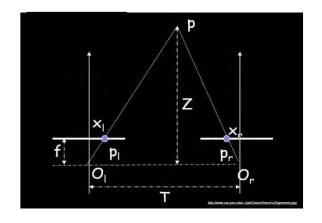


(i.e., the line is given by:  $l'^{\top}\mathbf{x} = 0$ )

 $oldsymbol{l} oldsymbol{l} = oldsymbol{E}^T oldsymbol{p}'$  is the coordinate vector representing the epipolar line for point p'



### Essential Matrix Example: Parallel Cameras



$$\mathbf{R} =$$

$$T =$$

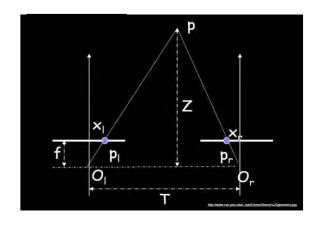
$$\mathbf{E} = [\mathbf{T}_{\mathbf{x}}]\mathbf{R} =$$

$$\mathbf{p'}^{\mathrm{T}}\mathbf{E}\mathbf{p} = 0$$

For the parallel cameras, image of any point must lie on same horizontal line in each image plane.

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### Essential Matrix Example: Parallel Cameras



$$\mathbf{R} = \mathbf{I}$$

$$\mathbf{T} = [-d, 0, 0]^{\mathrm{T}}$$

$$\mathbf{E} = [\mathbf{T}_{\mathbf{x}}]\mathbf{R} = \begin{bmatrix} \mathbf{0} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \mathbf{d} \\ \mathbf{0} & \mathbf{d} & \mathbf{0} \end{bmatrix}$$

$$\mathbf{p'}^{\mathsf{T}}\mathbf{E}\mathbf{p} = \mathbf{0} \qquad \begin{bmatrix} x' & y' & f \end{bmatrix} \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{bmatrix} \begin{bmatrix} x \\ y \\ f \end{bmatrix} = 0$$

For the parallel cameras, image of any point must lie on same horizontal line in each image plane.

$$\Leftrightarrow \begin{bmatrix} x' \ y' \ f \end{bmatrix} \begin{bmatrix} 0 \\ df \\ -dy \end{bmatrix} = 0$$
$$\Leftrightarrow y = y'$$



#### **More General Case**

Image I(x,y)

Disparity map D(x,y)

Image I'(x',y')







$$(x',y') = (x+D(x,y),y)$$

What about when cameras' optical axes are not parallel?

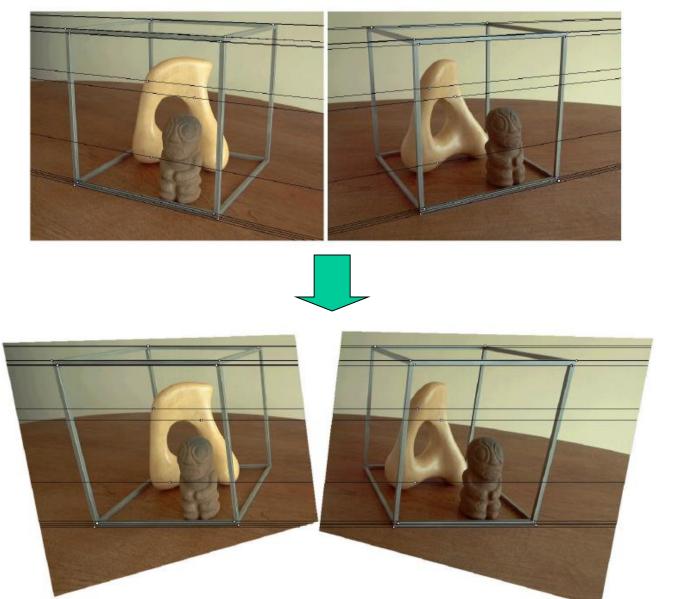
**Stereo Image Rectification** 

 In practice, it is convenient if image scanlines are the epipolar lines.



- Reproject image planes onto a common plane parallel to the line between optical centers
- Pixel motion is horizontal after this transformation
- > Two homographies  $(3\times3)$  transforms, one for each input image reprojection

### Stereo Image Rectification: Example



Source: Alyosha Efros



### Recap: Stereo Reconstruction

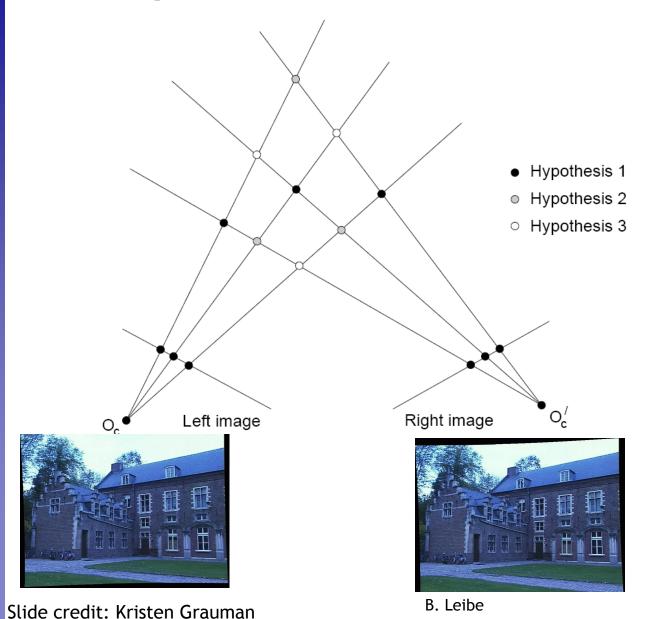
- Main Steps
  - Calibrate cameras
  - Rectify images
  - Compute disparity
  - Estimate depth







### **Correspondence Problem**



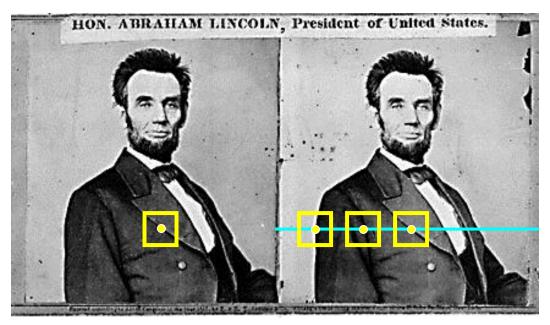
Multiple match hypotheses satisfy the epipolar constraint, but which is the correct one?

⇒ This ambiguity is what makes stereo estimation difficult!

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### Dense Correspondence Search

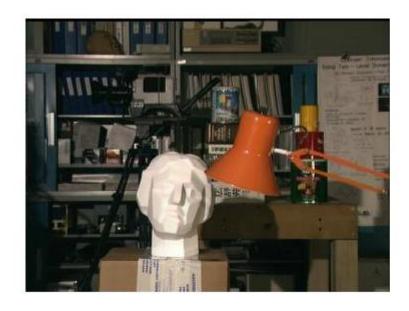


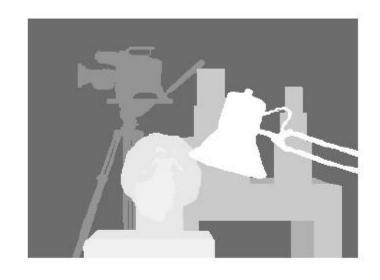
- For each pixel in the first image
  - Find corresponding epipolar line in the right image
  - Examine all pixels on the epipolar line and pick the best match (e.g. SSD, correlation)
  - Triangulate the matches to get depth information
- This is easiest when epipolar lines are scanlines
  - ⇒ Rectify images first



## **Example: Window Search**

Data from University of Tsukuba





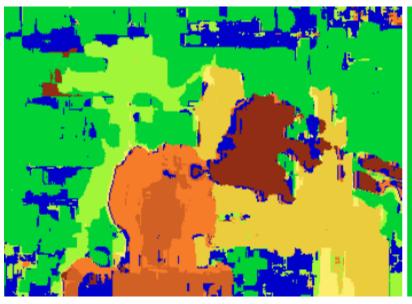
Scene

Ground truth



### **Example: Window Search**

Data from University of Tsukuba





Window-based matching (best window size)

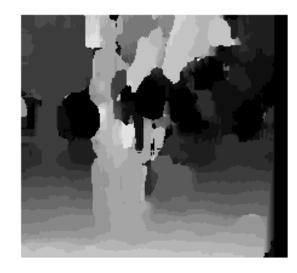
Ground truth



#### **Effect of Window Size**





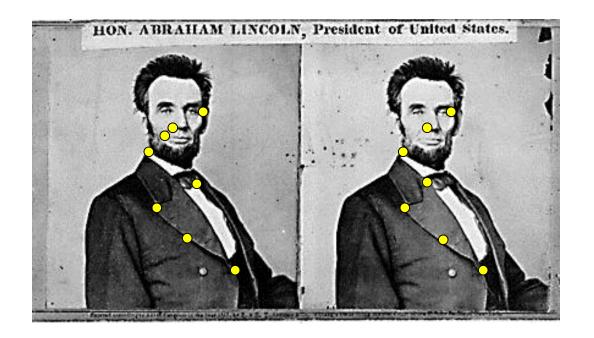


$$W=3$$

W = 20

Want window large enough to have sufficient intensity variation, yet small enough to contain only pixels with about the same disparity.

## Alternative: Sparse Correspondence Search



#### Idea:

- Restrict search to sparse set of detected features
- Rather than pixel values (or lists of pixel values) use feature descriptor and an associated feature distance
- Still narrow search further by epipolar geometry

What would make good features?

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### Dense vs. Sparse

#### Sparse

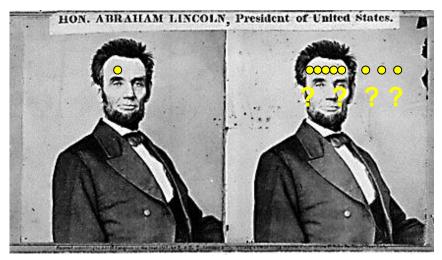
- Efficiency
- Can have more reliable feature matches, less sensitive to illumination than raw pixels
- But...
  - Have to know enough to pick good features
  - Sparse information

#### Dense

- Simple process
- More depth estimates, can be useful for surface reconstruction
- > **But...** 
  - Breaks down in textureless regions anyway
  - Raw pixel distances can be brittle
  - Not good with very different viewpoints



### **Difficulties in Similarity Constraint**



**Untextured surfaces** 



**Occlusions** 

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Slide credit: Kristen Grauman



## **Summary: Stereo Reconstruction**

#### Main Steps

- Calibrate cameras
- Rectify images
- Compute disparity
- Estimate depth



Left



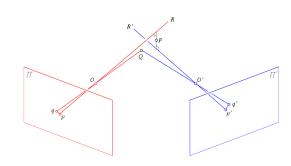




Left

Right

- So far, we have only considered calibrated cameras...
- Today
  - Uncalibrated cameras
  - Camera parameters
  - Revisiting epipolar geometry
  - Robust fitting





### Recap: A General Point

Equations of the form

$$Ax = 0$$

- How do we solve them? (always!)
  - Apply SVD

$$\mathbf{SVD} \downarrow \mathbf{A} = \mathbf{UDV}^T = \mathbf{U} \begin{bmatrix} d_{11} & & & \\ & \ddots & & \\ & & d_{NN} \end{bmatrix} \begin{bmatrix} v_{11} & \cdots & v_{1N} \\ \vdots & \ddots & \vdots \\ v_{N1} & \cdots & v_{NN} \end{bmatrix}^T$$

Singular values Singular vectors

- $\rightarrow$  Singular values of A = square roots of the eigenvalues of A<sup>T</sup>A.
- The solution of Ax=0 is the nullspace vector of A.
- This corresponds to the smallest singular vector of A.

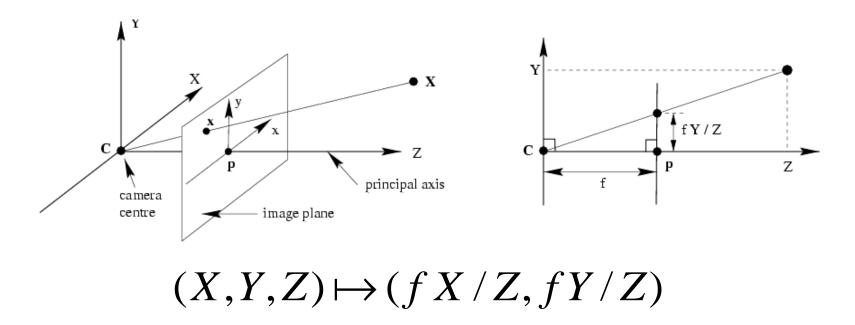


### **Topics of This Lecture**

- Camera Calibration
  - Camera parameters
  - Calibration procedure
- Revisiting Epipolar Geometry
  - Triangulation
  - > Calibrated case: Essential matrix
  - Uncalibrated case: Fundamental matrix
  - Weak calibration
  - > Epipolar Transfer
- Active Stereo
  - Laser scanning
  - Kinect sensor



### Recall: Pinhole Camera Model

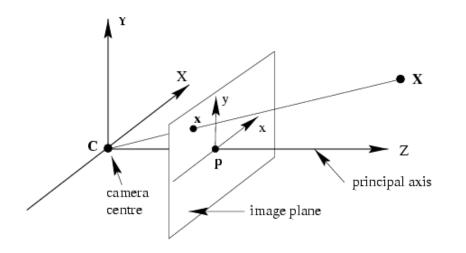


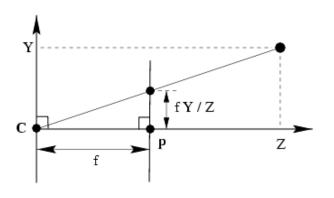
$$\begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix} \mapsto \begin{pmatrix} fX \\ fY \\ Z \end{pmatrix} = \begin{bmatrix} f & & & 0 \\ & f & & 0 \\ & & 1 & 0 \end{bmatrix} \begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix}$$

x = PX



#### Pinhole Camera Model



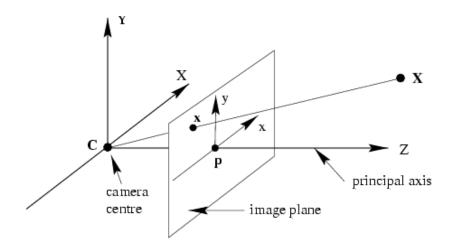


$$\begin{pmatrix} fX \\ fY \\ Z \end{pmatrix} = \begin{bmatrix} f \\ f \\ 1 \end{bmatrix} \begin{bmatrix} 1 \\ 1 \\ 0 \\ 1 \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \\ 1 \end{bmatrix}$$

$$x = PX$$
  $P = diag(f, f, 1)[I \mid 0]$ 



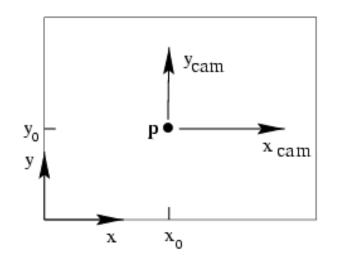
### **Camera Coordinate System**



- Principal axis:
  - Line from the camera center perpendicular to the image plane
- Normalized (camera) coordinate system:
  - Camera center is at the origin and the principal axis is the z-axis
- Principal point (p):
  - Point where principal axis intersects the image plane (origin of normalized coordinate system)



### **Principal Point Offset**

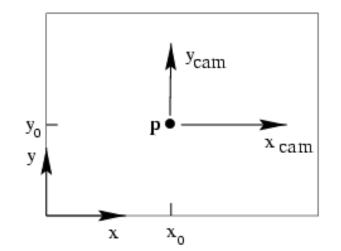


principal point: $(p_x, p_y)$ 

- Camera coordinate system: origin at the principal point
- Image coordinate system: origin is in the corner



### **Principal Point Offset**



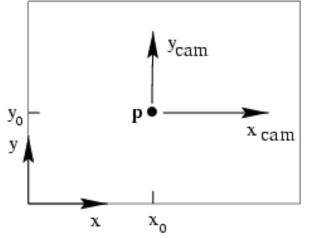
 $\rightarrow_{\mathbf{x}_{cam}}$  principal point: $(p_x, p_y)$ 

$$(X,Y,Z) \mapsto (fX/Z + p_x, fY/Z + p_y)$$

$$\begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix} \mapsto \begin{pmatrix} fX + Zp_x \\ fY + Zp_x \\ Z \end{pmatrix} = \begin{bmatrix} f & p_x & 0 \\ f & p_y & 0 \\ 1 & 0 \end{bmatrix} \begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix}$$



### Principal Point Offset



 $\rightarrow$  principal point: $(p_x, p_y)$ 

$$\begin{pmatrix} fX + Zp_x \\ fY + Zp_x \\ Z \end{pmatrix} = \begin{bmatrix} f & p_x \\ f & p_y \\ 1 \end{bmatrix} \begin{bmatrix} 1 & 0 \\ 1 & 0 \\ 1 \end{bmatrix} \begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix}$$

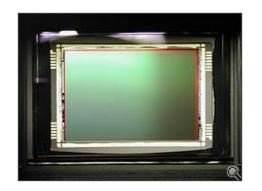
$$K = \begin{bmatrix} f & p_x \\ f & p_y \\ 1 \end{bmatrix}$$
 calibration matrix  $P = K[I | 0]$ 

$$P = K[I \mid 0]$$



### Pixel Coordinates: Non-Square Pixels





Pixel size: 
$$\frac{1}{m_x} \times \frac{1}{m_y}$$

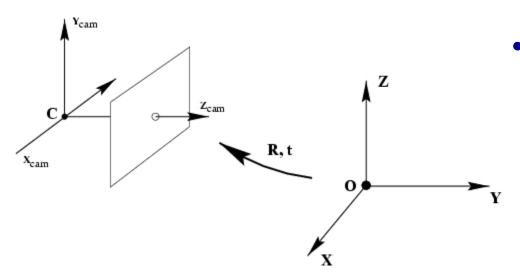
 $m_{x}$  pixels per meter in horizontal direction,  $m_{\rm v}$  pixels per meter in vertical direction

$$K = \begin{bmatrix} m_x & & \\ & m_y & \\ & & 1 \end{bmatrix} \begin{bmatrix} f & p_x \\ & f & p_y \\ & & 1 \end{bmatrix} = \begin{bmatrix} \alpha_x & x_0 \\ & \alpha_y & y_0 \\ & & 1 \end{bmatrix}$$

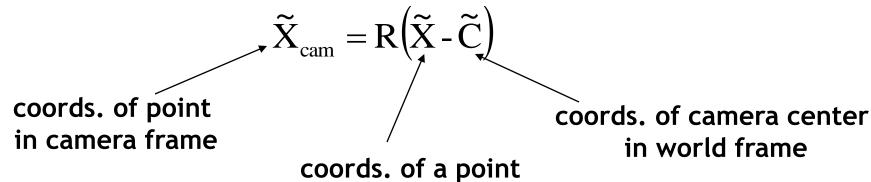
[pixels/m] [m] [pixels]



### **Camera Rotation and Translation**



In general, the camera coordinate frame will be related to the world coordinate frame by a rotation and a translation

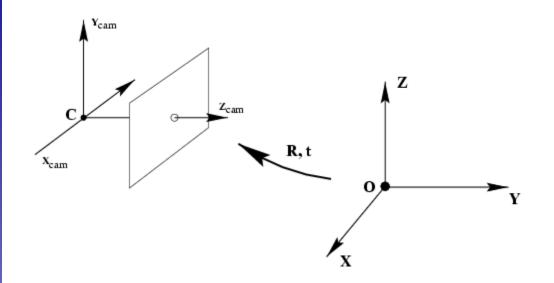


in world frame (nonhomogeneous)





#### Camera Rotation and Translation



In non-homogeneous coordinates:

$$\widetilde{X}_{cam} = R(\widetilde{X} - \widetilde{C})$$

$$X_{cam} = \begin{bmatrix} R & -R\tilde{C} \\ 0 & 1 \end{bmatrix} \begin{bmatrix} \tilde{X} \\ 1 \end{bmatrix} = \begin{bmatrix} R & -R\tilde{C} \\ 0 & 1 \end{bmatrix} X$$

$$x = K[I|0]X_{cam} = K[R|-R\tilde{C}]X$$
  $P = K[R|t],$   $t = -R\tilde{C}$ 

Note: C is the null space of the camera projection matrix (PC=0)

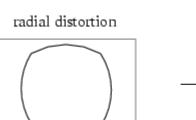


### **Summary: Camera Parameters**

#### Intrinsic parameters

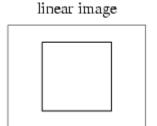
- Principal point coordinates
- Focal length
- Pixel magnification factors
- Skew (non-rectangular pixels)
- Radial distortion







 $K = \begin{bmatrix} m_x & & & \\ & m_y & & \\ & & 1 \end{bmatrix} \begin{bmatrix} f & \mathbf{S} & p_x \\ & f & p_y \\ & & 1 \end{bmatrix} = \begin{bmatrix} \alpha_x & \mathbf{S} & x_0 \\ & \alpha_y & y_0 \\ & & 1 \end{bmatrix}$ 



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correction



## **Summary: Camera Parameters**

#### Intrinsic parameters

- Principal point coordinates
- Focal length
- Pixel magnification factors
- Skew (non-rectangular pixels)
- Radial distortion

#### Extrinsic parameters

- ightharpoonup Rotation  ${f R}$
- Translation t

 $K = \begin{bmatrix} m_x & & & & \\ & m_y & & & \\ & & 1 \end{bmatrix} \begin{bmatrix} f & \mathbf{S} & p_x \\ & f & p_y \\ & & 1 \end{bmatrix} = \begin{bmatrix} \alpha_x & \mathbf{S} & x_0 \\ & \alpha_y & y_0 \\ & & 1 \end{bmatrix}$ 

How many degrees of freedom does P have?

## Camera Parameters: Degrees of Freedom

Intrinsic parameters

DoF

Principal point coordinates

- 2
- Focal length
- 1

 $K = \begin{vmatrix} g t_x & s & p_x t_0 \\ f \alpha_y & p_y t_0 \\ 1 & 1 \end{vmatrix}$ 

- Pixel magnification factors
- 4
- Skew (non-rectangular pixels)
- Radial distortion
- Extrinsic parameters
  - Rotation R

- 3
- Translation t
   (both relative to world coordinate system)
- Camera projection matrix

 $P = K[R \mid t]$ 

⇒ General pinhole camera:

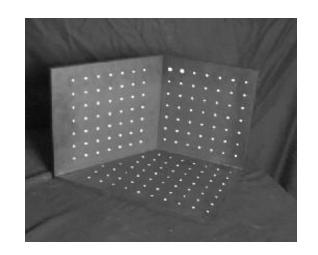
- 9 DoF
- ⇒ CCD Camera with square pixels: 10 DoF
- ⇒ General camera:

11 DoF



### Calibrating a Camera

 Compute intrinsic and extrinsic parameters using observed camera data.



#### Main idea

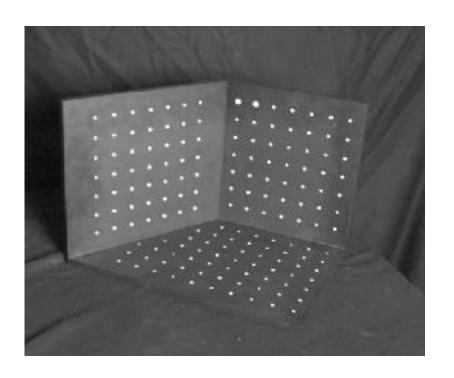
- Place "calibration object" with known geometry in the scene
- Get correspondences
- Solve for mapping from scene to image: estimate P=P<sub>int</sub>P<sub>ext</sub>

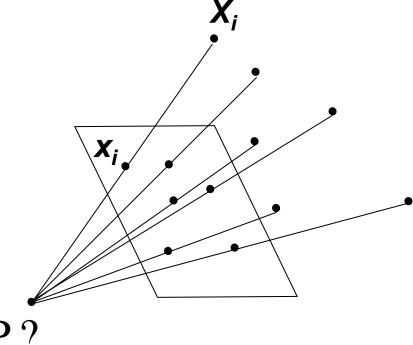




#### **Camera Calibration**

• Given n points with known 3D coordinates  $X_i$  and known image projections  $x_i$ , estimate the camera parameters





# Camera Calibration: Obtaining the Points

- For best results, it is important that the calibration points are measured with subpixel accuracy.
- How this can be done depends on the exact pattern.
- Algorithm for checkerboard pattern
  - 1. Perform Canny edge detection.
  - 2. Fit straight lines to detected linked edges.
  - 3. Intersect lines to obtain corners.
  - If sufficient care is taken, the points can then be obtained with localization accuracy < 1/10 pixel.</p>

#### Rule of thumb

- Number of constraints should exceed number of unknowns by a factor of five.
- $\Rightarrow$  For 11 parameters of P, at least 28 points should be used.



### Camera Calibration: DLT Algorithm

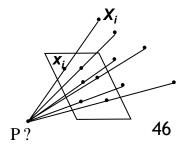
(DLT = "Direct Linear Transform")

$$\lambda \mathbf{x}_{i} = \mathbf{P} \mathbf{X}_{i} \qquad \lambda \begin{bmatrix} x_{i} \\ y_{i} \\ 1 \end{bmatrix} = \begin{bmatrix} P_{11} & P_{12} & P_{13} & P_{14} \\ P_{21} & P_{22} & P_{23} & P_{24} \\ P_{31} & P_{32} & P_{33} & P_{34} \end{bmatrix} \begin{bmatrix} \mathbf{X}_{i,1} \\ \mathbf{X}_{i,2} \\ \mathbf{X}_{i,3} \\ 1 \end{bmatrix} = \begin{bmatrix} \mathbf{P}_{1}^{T} \\ \mathbf{P}_{2}^{T} \\ \mathbf{P}_{3}^{T} \end{bmatrix} \mathbf{X}_{i}$$

$$\mathbf{x}_i \times \mathbf{P} \mathbf{X}_i = 0$$

$$\begin{bmatrix} 0 & -X_i^T & y_i X_i^T \\ X_i^T & 0 & -x_i X_i^T \\ -y_i X_i^T & x_i X_i^T & 0 \end{bmatrix} \begin{pmatrix} P_1 \\ P_2 \\ P_3 \end{pmatrix} = 0$$

Only two linearly independent equations



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### Camera Calibration: DLT Algorithm

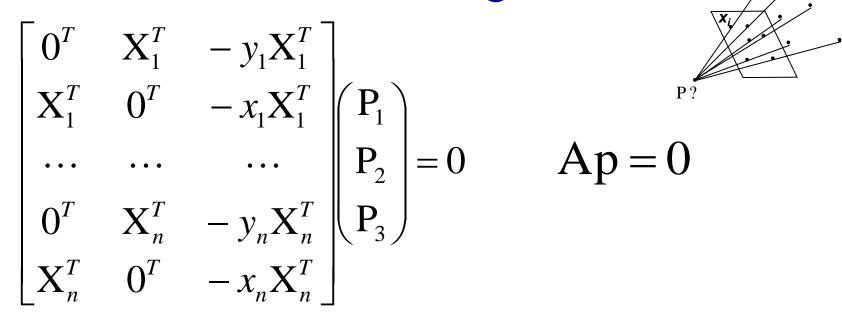
$$\begin{bmatrix} 0^T & X_1^T & -y_1 X_1^T \\ X_1^T & 0^T & -x_1 X_1^T \\ \cdots & \cdots & \cdots \\ 0^T & X_n^T & -y_n X_n^T \\ X_n^T & 0^T & -x_n X_n^T \end{bmatrix} P_1$$

$$\begin{bmatrix} P_1 \\ P_2 \\ P_3 \end{bmatrix} = 0 \quad Ap = 0$$
Solve using... SVD!

#### Notes

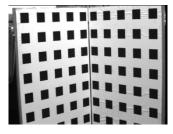
- P has 11 degrees of freedom (12 parameters, but scale is arbitrary).
- One 2D/3D correspondence gives us two linearly independent equations.
- Homogeneous least squares (similar to homography est.)
- $\Rightarrow$  5 ½ correspondences needed for a minimal solution.

### Camera Calibration: DLT Algorithm



#### Notes

- For coplanar points that satisfy  $\Pi^TX=0$ , we will get degenerate solutions  $(\Pi,0,0)$ ,  $(0,\Pi,0)$ , or  $(0,0,\Pi)$ .
- ⇒ We need calibration points in more than one plane!





#### **Camera Calibration**

- Once we've recovered the numerical form of the camera matrix, we still have to figure out the intrinsic and extrinsic parameters
- This is a matrix decomposition problem, not an estimation problem (see F&P sec. 3.2, 3.3)

### Camera Calibration: Some Practical Tips

- For numerical reasons, it is important to carry out some data normalization.
  - > Translate the image points  $x_i$  to the (image) origin and scale them such that their RMS distance to the origin is  $\sqrt{2}$ .
  - > Translate the 3D points  $X_i$  to the (world) origin and scale them such that their RMS distance to the origin is  $\sqrt{3}$ .
  - (This is valid for compact point distributions on calibration objects).
- The DLT algorithm presented here is easy to implement, but there are some more accurate algorithms available (see H&Z sec. 7.2).
- For practical applications, it is also often needed to correct for radial distortion. Algorithms for this can be found in H&Z sec. 7.4, or F&P sec. 3.3.

B Leib



### **Topics of This Lecture**

- Camera Calibration
  - > Camera parameters
  - Calibration procedure
- Revisiting Epipolar Geometry
  - Triangulation
  - Calibrated case: Essential matrix
  - Uncalibrated case: Fundamental matrix
  - Weak calibration
  - Epipolar Transfer
- Active Stereo
  - Laser scanning
  - Kinect sensor



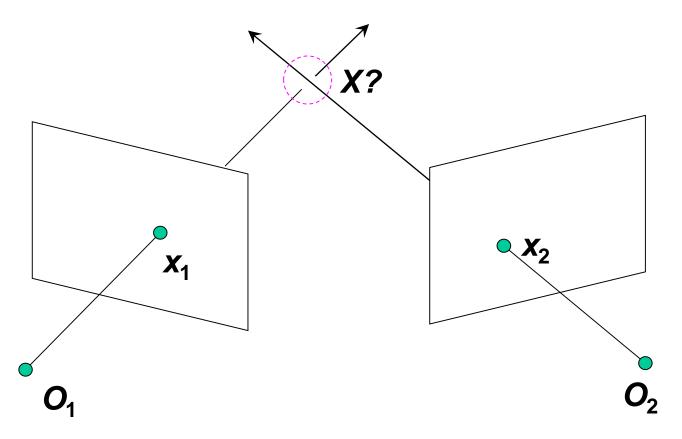
### **Two-View Geometry**

- Scene geometry (structure):
  - Given corresponding points in two or more images, where is the pre-image of these points in 3D?
- Correspondence (stereo matching):
  - Given a point in just one image, how does it constrain the position of the corresponding point x' in another image?
- Camera geometry (motion):
  - Given a set of corresponding points in two images, what are the cameras for the two views?



### **Revisiting Triangulation**

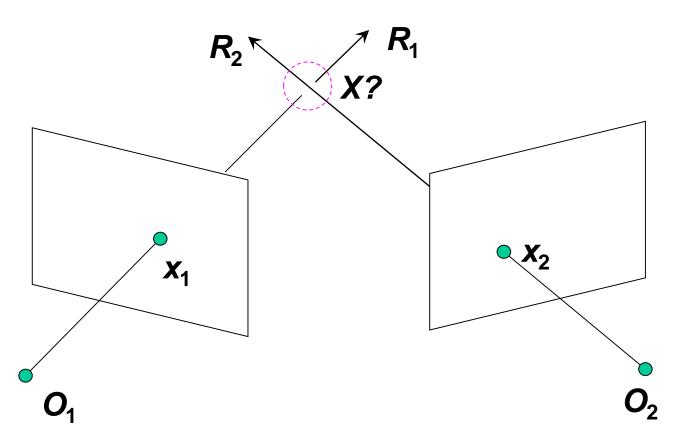
 Given projections of a 3D point in two or more images (with known camera matrices), find the coordinates of the point





### **Revisiting Triangulation**

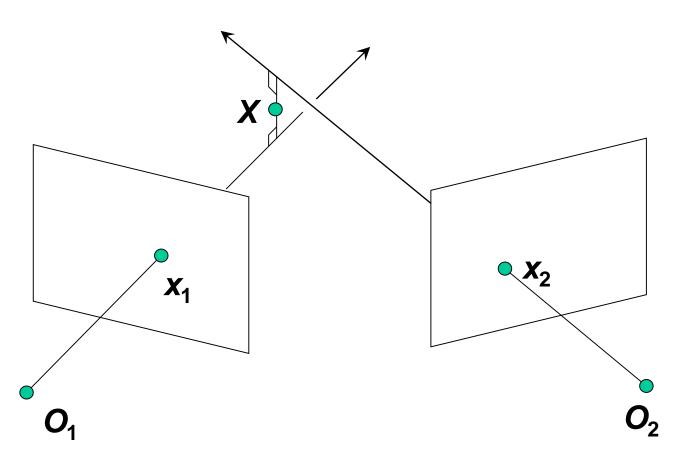
• We want to intersect the two visual rays corresponding to  $x_1$  and  $x_2$ , but because of noise and numerical errors, they will never meet exactly. How can this be done?





### Triangulation: 1) Geometric Approach

 Find shortest segment connecting the two viewing rays and let X be the midpoint of that segment.



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### Triangulation: 2 )Linear Algebraic Approach

$$\lambda_1 X_1 = P_1 X$$

$$X_1 \times P_1 X = 0$$

$$[\mathbf{x}_{1\times}]\mathbf{P}_1\mathbf{X} = 0$$

$$\lambda_2 \mathbf{x}_2 = \mathbf{P}_2 \mathbf{X}$$

$$\mathbf{x}_2 \times \mathbf{P}_2 \mathbf{X} = \mathbf{0}$$

$$\lambda_1 x_1 = P_1 X$$
  $x_1 \times P_1 X = 0$   $[x_{1x}]P_1 X = 0$   
 $\lambda_2 x_2 = P_2 X$   $x_2 \times P_2 X = 0$   $[x_{2x}]P_2 X = 0$ 

#### Cross product as matrix multiplication:

$$\mathbf{a} \times \mathbf{b} = \begin{bmatrix} 0 & -a_z & a_y \\ a_z & 0 & -a_x \\ -a_y & a_x & 0 \end{bmatrix} \begin{bmatrix} b_x \\ b_y \\ b_z \end{bmatrix} = [\mathbf{a}_{\times}]\mathbf{b}$$

### Triangulation: 2) Linear Algebraic Approach

$$\lambda_1 x_1 = P_1 X$$
  $x_1 \times P_1 X = 0$   $[x_{1x}]P_1 X = 0$   
 $\lambda_2 x_2 = P_2 X$   $x_2 \times P_2 X = 0$   $[x_{2x}]P_2 X = 0$ 

Two independent equations each in terms of three unknown entries of X

⇒ Stack them and solve using SVD!

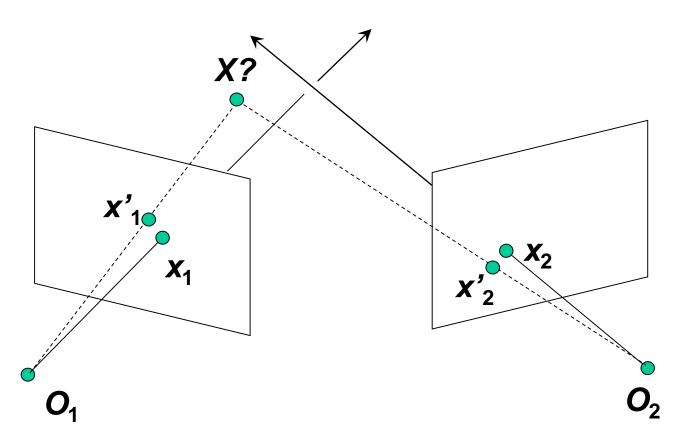
 This approach is often preferable to the geometric approach, since it nicely generalizes to multiple cameras.



### Triangulation: 3) Nonlinear Approach

Find X that minimizes

$$d^{2}(x_{1}, P_{1}X) + d^{2}(x_{2}, P_{2}X)$$





### Triangulation: 3) Nonlinear Approach

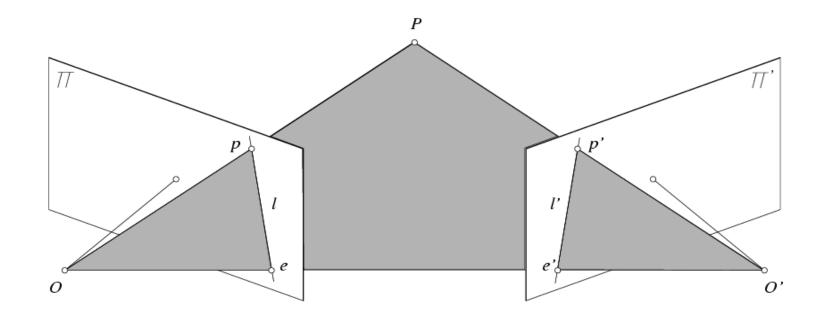
Find X that minimizes

$$d^{2}(x_{1}, P_{1}X) + d^{2}(x_{2}, P_{2}X)$$

- This approach is the most accurate, but unlike the other two methods, it doesn't have a closed-form solution.
- Iterative algorithm
  - Initialize with linear estimate.
  - Optimize with Gauss-Newton or Levenberg-Marquardt (see F&P sec. 3.1.2 or H&Z Appendix 6).



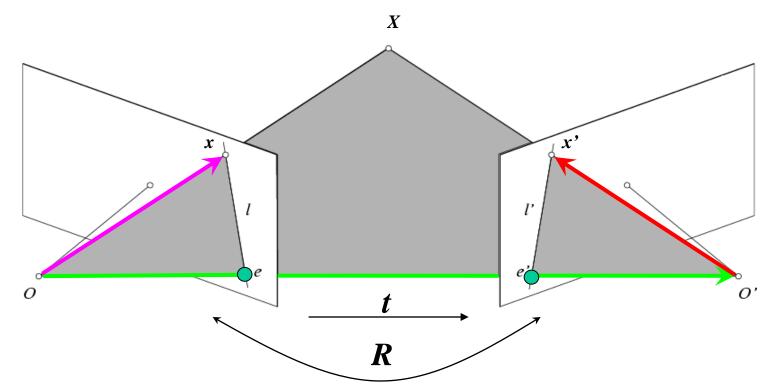
### **Revisiting Epipolar Geometry**



- Let's look again at the epipolar constraint
  - For the calibrated case (but in homogenous coordinates)
  - For the uncalibrated case



### **Epipolar Geometry: Calibrated Case**



#### Camera matrix: [I|0]

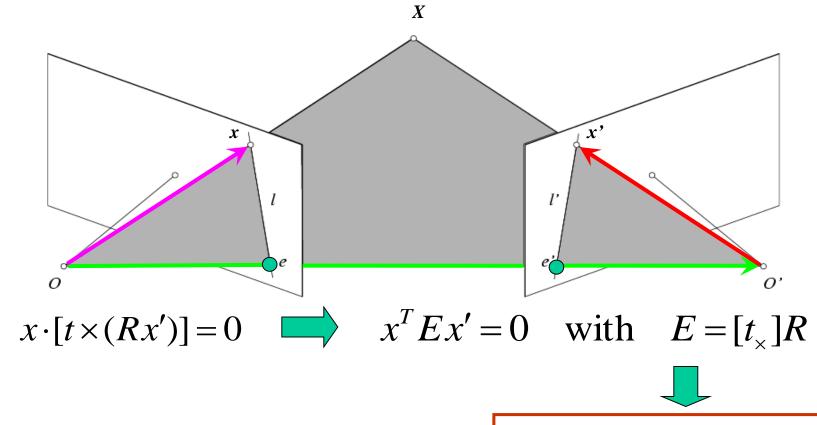
$$X = (u, v, w, 1)^T$$
$$x = (u, v, w)^T$$

Camera matrix:  $[R^T | -R^T t]$ Vector x' in second coord. system has coordinates Rx' in the first one.

The vectors x, t, and Rx are coplanar

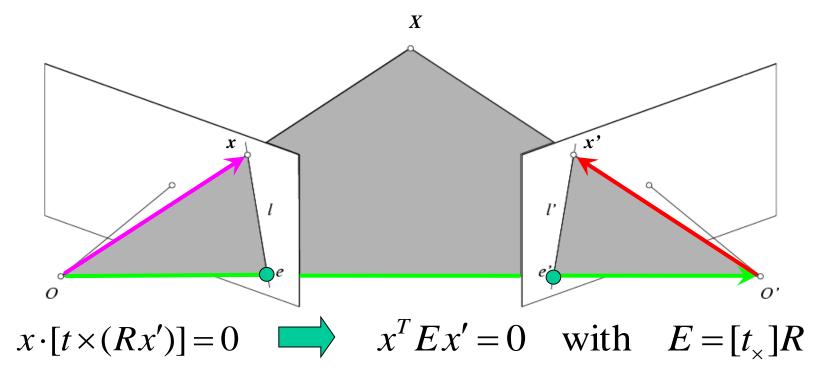
61

### **Epipolar Geometry: Calibrated Case**



Essential Matrix (Longuet-Higgins, 1981)

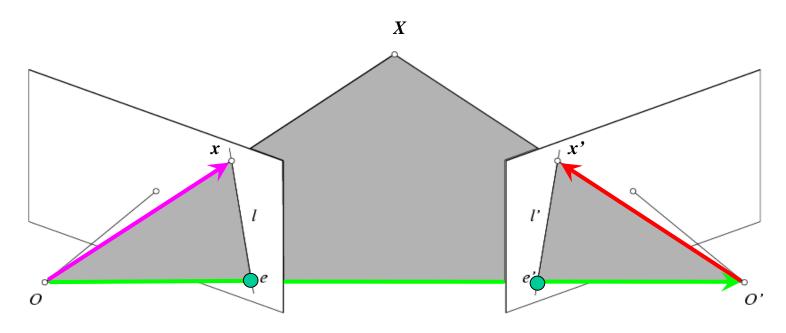
### **Epipolar Geometry: Calibrated Case**



- E x' is the epipolar line associated with x' (l = E x')
- $E^Tx$  is the epipolar line associated with x ( $l' = E^Tx$ )
- E e' = 0 and  $E^{T}e = 0$
- E is singular (rank two)
- E has five degrees of freedom (up to scale)



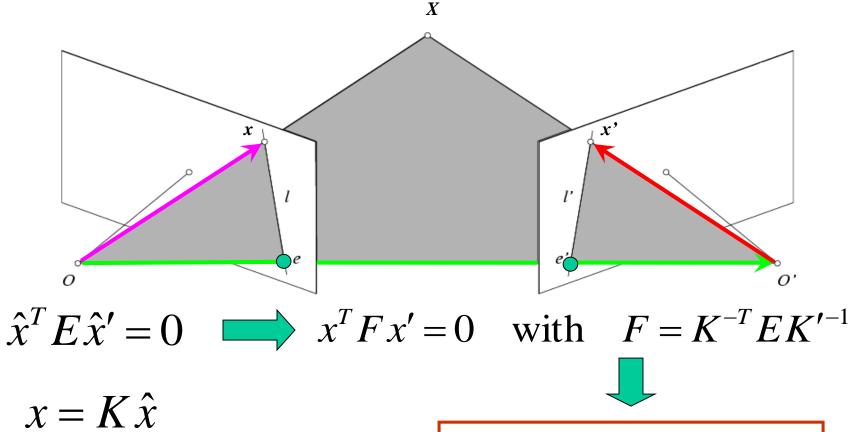
### **Epipolar Geometry: Uncalibrated Case**



- The calibration matrices K and K' of the two cameras are unknown
- We can write the epipolar constraint in terms of unknown normalized coordinates:

$$\hat{x}^T E \hat{x}' = 0 \qquad x = K \hat{x}, \quad x' = K' \hat{x}'$$

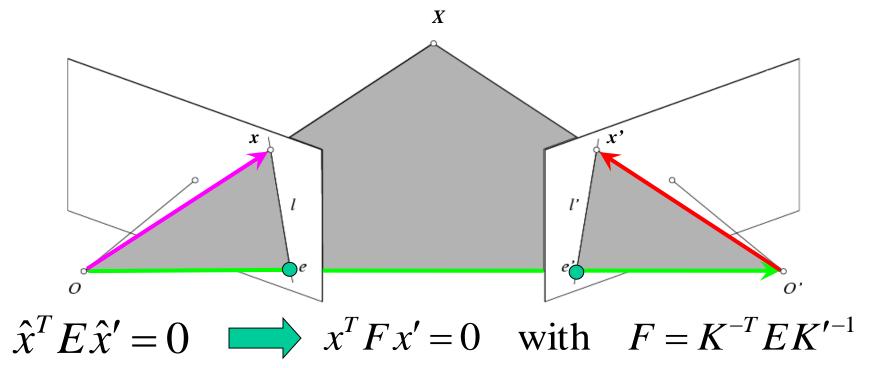
### Epipolar Geometry: Uncalibrated Case



$$x' = K'\hat{x}'$$

Fundamental Matrix (Faugeras and Luong, 1992)

### **Epipolar Geometry: Uncalibrated Case**



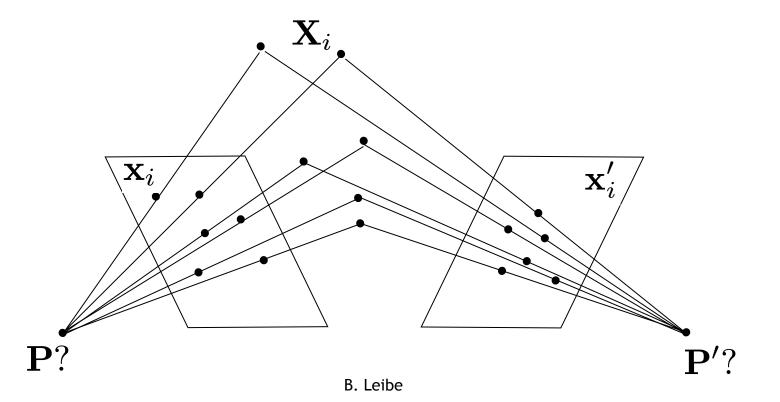
- F x' is the epipolar line associated with x'(l = F x')
- $F^Tx$  is the epipolar line associated with x  $(l' = F^Tx)$
- Fe' = 0 and  $F^{T}e = 0$
- F is singular (rank two)
- F has seven degrees of freedom

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### **Estimating the Fundamental Matrix**

- The Fundamental matrix defines the epipolar geometry between two uncalibrated cameras.
- How can we estimate F from an image pair?
  - We need correspondences...







 $F_{11}$ 

 $F_{12}$ 

 $F_{13}$ 

 $F_{31}$ 

### The Eight-Point Algorithm

$$x = (u, v, 1)^T, x' = (u', v', 1)^T$$





#### Taking 8 correspondences:

$\begin{vmatrix} u_2' u_2 \\ u_3' u_3 \\ u_4' u_4 \\ u_5' u_5 \\ u_6' u_6 \end{vmatrix}$	$u_{3}^{7}v_{3} \ u_{4}^{7}v_{4} \ u_{5}^{7}v_{5} \ u_{6}^{7}v_{6}$	$u'_{2} \\ u'_{3} \\ u'_{4} \\ u'_{5} \\ u'_{6}$	$u_{2}v_{2}'$ $u_{3}v_{3}'$ $u_{4}v_{4}'$ $u_{5}v_{5}'$ $u_{6}v_{6}'$	$v_2v_2' \ v_3v_3' \ v_4v_4' \ v_5v_5' \ v_6v_6'$	$v_{2}^{'} \\ v_{3}^{'} \\ v_{4}^{'} \\ v_{5}^{'} \\ v_{6}^{'}$	$u_2 \\ u_3 \\ u_4 \\ u_5 \\ u_6$	$egin{array}{c} v_2 \ v_3 \ v_4 \ v_5 \ v_6 \end{array}$	1 1 1 1 1 1 1
$u_7'u_7$	$u_{6}'v_{6} \ u_{7}'v_{7} \ u_{8}'v_{8}$	$u_7^{'}$	$u_7v_7^{\prime}$	$v_7v_7^{''}$	$v_7^{\prime}$	$u_7$	$v_7$	1 1 1

#### Solve using... SVD!

Af = 0

This minimizes:

Slide adapted from Svetlana Lazebnik

B. Leibe



### **Excursion: Properties of SVD**

- Frobenius norm
  - Generalization of the Euclidean norm to matrices

$$||A||_F = \sqrt{\sum_{i=1}^m \sum_{j=1}^n |a_{ij}|^2} = \sqrt{\sum_{i=1}^{\min(m,n)} \sigma_i^2}$$

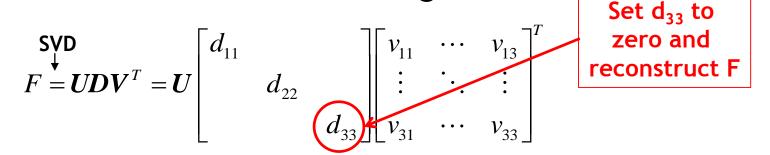
- Partial reconstruction property of SVD
  - Let  $\sigma_i$  i=1,...,N be the singular values of A.
  - Let  $A_p = U_p D_p V_p^T$  be the reconstruction of A when we set  $\sigma_{p+1},...,\ \sigma_N$  to zero.
  - Then  $A_p = U_p D_p V_p^T$  is the best rank-p approximation of A in the sense of the Frobenius norm (i.e. the best least-squares approximation).

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### The Eight-Point Algorithm

- Problem with noisy data
  - The solution will usually not fulfill the constraint that F only has rank 2.
  - ⇒ There will be no epipoles through which all epipolar lines pass!
- Enforce the rank-2 constraint using SVD



 As we have just seen, this provides the best leastsquares approximation to the rank-2 solution.



### Problem with the Eight-Point Algorithm

• In practice, this often looks as follows:

$egin{array}{c} u_2'u_2 \ u_3'u_3 \ u_4'u_4 \ u_5'u_5 \ u_6'u_6 \ \end{array}$	$u_{2}^{'}v_{2} \ u_{3}^{'}v_{3} \ u_{4}^{'}v_{4} \ u_{5}^{'}v_{5} \ u_{6}^{'}v_{6}$	$u_{2}^{'}$ $u_{3}^{'}$ $u_{4}^{'}$ $u_{5}^{'}$ $u_{6}^{'}$	$u_1v_1' \\ u_2v_2' \\ u_3v_3' \\ u_4v_4' \\ u_5v_5' \\ u_6v_6'$	$v_2v_2' \ v_3v_3' \ v_4v_4' \ v_5v_5' \ v_6v_6'$	$v_{2}^{'} \\ v_{3}^{'} \\ v_{4}^{'} \\ v_{5}^{'} \\ v_{6}^{'}$	$u_2 \\ u_3 \\ u_4 \\ u_5 \\ u_6$	$egin{array}{c} v_3 \ v_4 \ v_5 \ v_6 \end{array}$	1 1 1 1 1 1 1	$egin{array}{c} F_{11} \ F_{12} \ F_{13} \ F_{21} \ F_{22} \ F_{23} \ F_{31} \ \end{array}$	=	
$\begin{vmatrix} u_6'u_6 \\ u_7'u_7 \end{vmatrix}$	$u_6'v_6 \\ u_7'v_7$	$u_6^\prime \ u_7^\prime$		$egin{array}{c} v_6v_6' \ v_7v_7' \end{array}$	$v_6^\prime \ v_7^\prime$	$u_6\\u_7$	$egin{array}{c} v_6 \ v_7 \end{array}$	1 1 1 1			

### Problem with the Eight-Point Algorithm

In practice, this often looks as follows:

250906.36 2692.28	183269.57 131633.03		200931.10	146766.13 302975.59		272.19 15.27	198.81 1 746.79 1	$\left egin{array}{c} F_{11} \ F_{12} \ \end{array} ight $
416374.23 191183.60	871684.30 171759.40	935.47	408110.89		916.90 893.65	445.10	931.81 1 418.65 1	$\left egin{array}{c} F_{13} \ F_{21} \ \end{array} ight $
48988.86	30401.76	57.89	298604.57	185309.58	352.87	846.22	525.15 1	$\left  \begin{array}{c} F_{22} \\ F_{23} \end{array} \right $
164786.04 116407.01	546559.67 2727.75		1998.37 169941.27	6628.15 3982.21	9.86 202.77	202.65 838.12	19.64	$\left egin{array}{c} F_{31} \ F_{32} \end{array} ight $
135384.58	75411.13	198.72	411350.03	229127.78	603.79	681.28	379.48 1	$\mid \mid_{F_{33}}^{32} \mid$

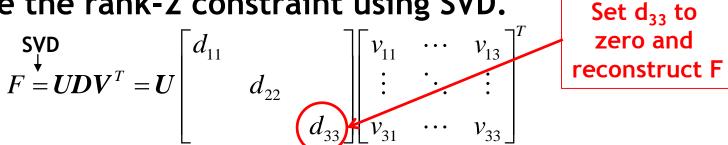
	$\lceil F_{11} \rceil$		
			$\begin{bmatrix} 0 \end{bmatrix}$
	$ F_{12} $		0
	$ F_{13} $		0
	$ F_{21} $		_
			0
	$ F_{22} $	=	0
	$ F_{23} $		
			0
	$F_{31}$		0
	$ F_{32} $		0
]	$F_{22}$		$LU_{L}$

- ⇒ Poor numerical conditioning
- ⇒ Can be fixed by rescaling the data



### The Normalized Eight-Point Algorithm

- 1. Center the image data at the origin, and scale it so the mean squared distance between the origin and the data points is 2 pixels.
- 2. Use the eight-point algorithm to compute *F* from the normalized points.
- 3. Enforce the rank-2 constraint using SVD.



4. Transform fundamental matrix back to original units: if T and T' are the normalizing transformations in the two images, than the fundamental matrix in original coordinates is  $T^T F T'$ .

73 [Hartley, 1995]



### The Eight-Point Algorithm

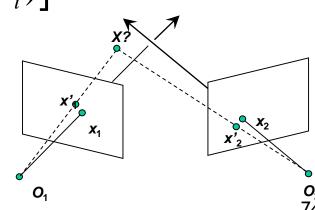
• Meaning of error  $\sum_{i=1}^{N} (x_i^T F x_i')^2$ :

Sum of Euclidean distances between points  $x_i$  and epipolar lines  $Fx'_i$  (or points  $x'_i$  and epipolar lines  $F^Tx_i$ ), multiplied by a scale factor

Nonlinear approach: minimize

$$\sum_{i=1}^{N} \left[ d^{2}(x_{i}, F x_{i}') + d^{2}(x_{i}', F^{T} x_{i}) \right]$$

- Similar to nonlinear minimization approach for triangulation.
- Iterative approach (Gauss-Newton, Levenberg-Marquardt,...)



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## **Comparison of Estimation Algorithms**









	8-point	Normalized 8-point	Nonlinear least squares
Av. Dist. 1	2.33 pixels	0.92 pixel	0.86 pixel
Av. Dist. 2	2.18 pixels	0.85 pixel	0.80 pixel

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Slide credit: Svetlana Lazebnik

### 3D Reconstruction with Weak Calibration

- Want to estimate world geometry without requiring calibrated cameras.
- Many applications:
  - Archival videos
  - Photos from multiple unrelated users
  - Dynamic camera system
- Main idea:
  - Estimate epipolar geometry from a (redundant) set of point correspondences between two uncalibrated cameras.



### Stereo Pipeline with Weak Calibration

- So, where to start with uncalibrated cameras?
  - Need to find fundamental matrix F and the correspondences (pairs of points  $(u',v') \leftrightarrow (u,v)$ ).





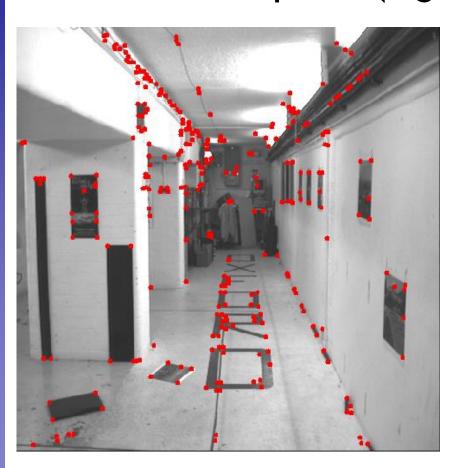
#### Procedure

- 1. Find interest points in both images
- 2. Compute correspondences
- 3. Compute epipolar geometry
- 4. Refine

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# Stereo Pipeline with Weak Calibration

#### 1. Find interest points (e.g. Harris corners)

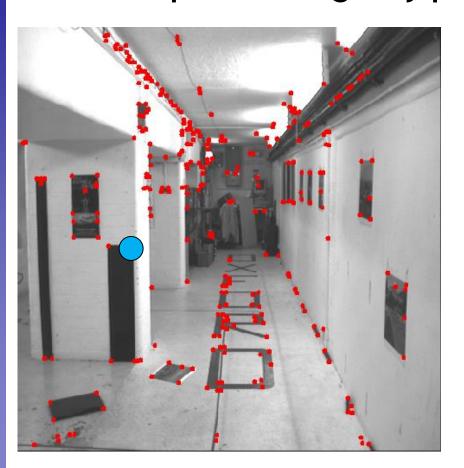


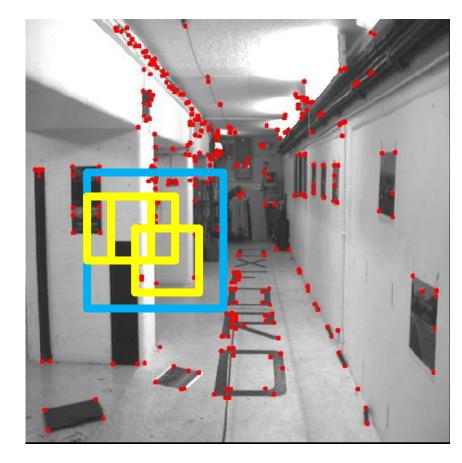


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# Stereo Pipeline with Weak Calibration

#### 2. Match points using only proximity







#### Putative Matches based on Correlation Search





Many wrong matches (10-50%), but enough to compute F



### **RANSAC** for Robust Estimation of F

- Select random sample of correspondences
- Compute F using them
  - This determines epipolar constraint
- Evaluate amount of support number of inliers within threshold distance of epipolar line

Choose F with most support (#inliers)





#### Putative Matches based on Correlation Search





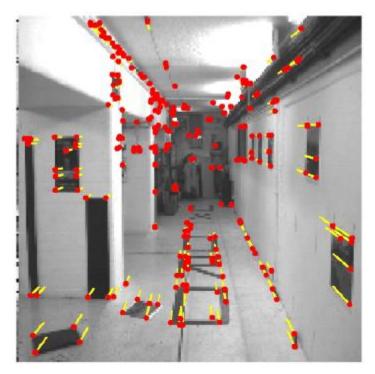
Many wrong matches (10-50%), but enough to compute F



#### **Pruned Matches**

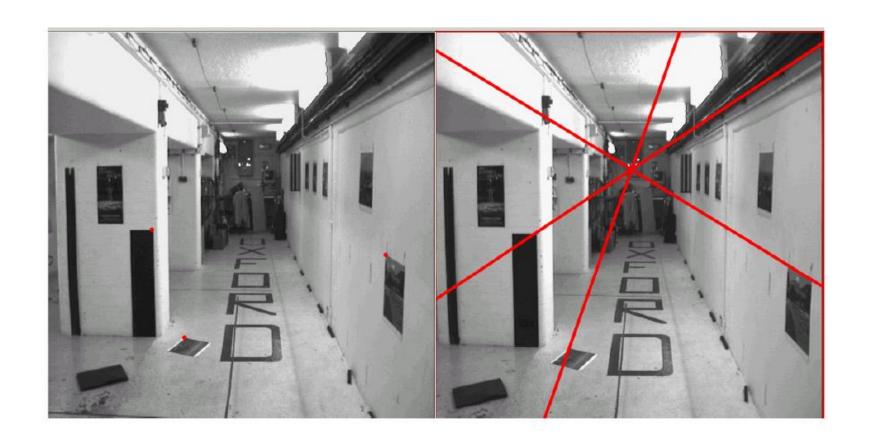
Correspondences consistent with epipolar geometry







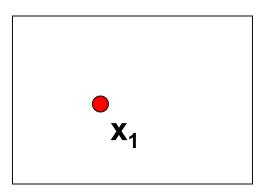
# **Resulting Epipolar Geometry**

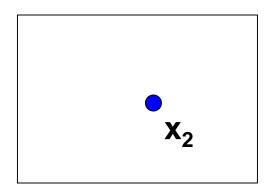


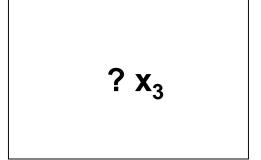


# **Epipolar Transfer**

- Assume the epipolar geometry is known
- Given projections of the same point in two images, how can we compute the projection of that point in a third image?



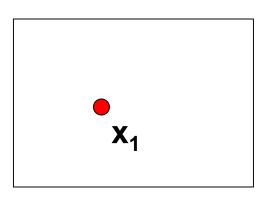


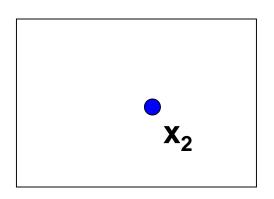


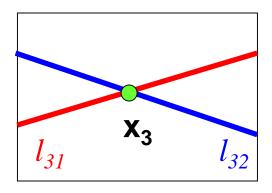


# Extension: Epipolar Transfer

- Assume the epipolar geometry is known
- Given projections of the same point in two images, how can we compute the projection of that point in a third image?







$$l_{31} = F^T_{13} x_1$$

 $l_{32} = F^{T}_{23} x_2$ 

When does epipolar transfer fail?



# **Topics of This Lecture**

- Camera Calibration
  - > Camera parameters
  - Calibration procedure
- Revisiting Epipolar Geometry
  - > Triangulation
  - Calibrated case: Essential matrix
  - Uncalibrated case: Fundamental matrix
  - Weak calibration
  - Epipolar Transfer
- Active Stereo
  - Laser scanning
  - Kinect sensor

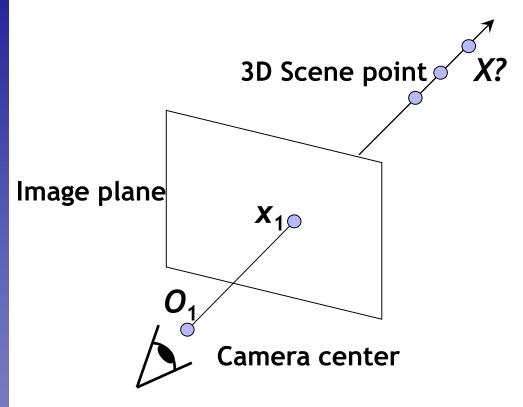
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### Microsoft Kinect - How Does It Work?



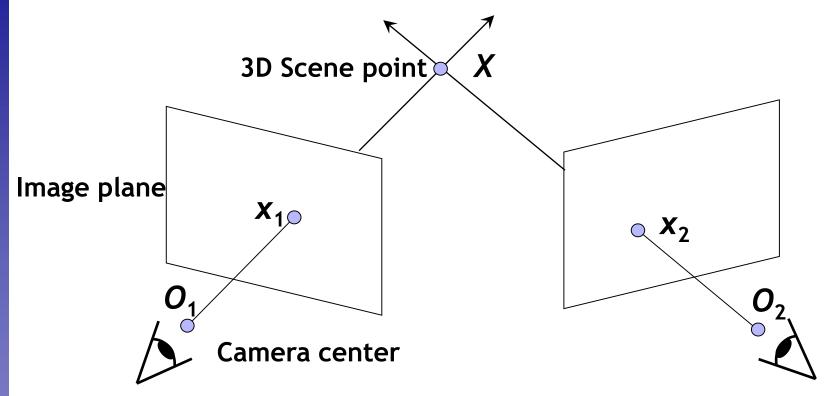


# Recall: Optical Triangulation





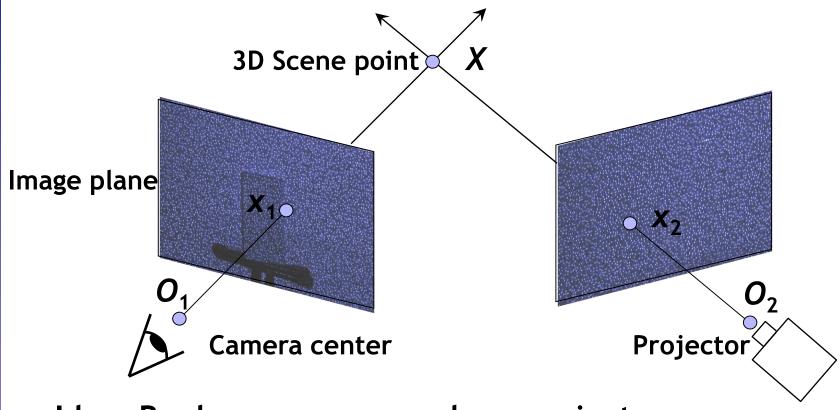
# Recall: Optical Triangulation



- Principle: 3D point given by intersection of two rays.
  - Crucial information: point correspondence
  - Most expensive and error-prone step in the pipeline...



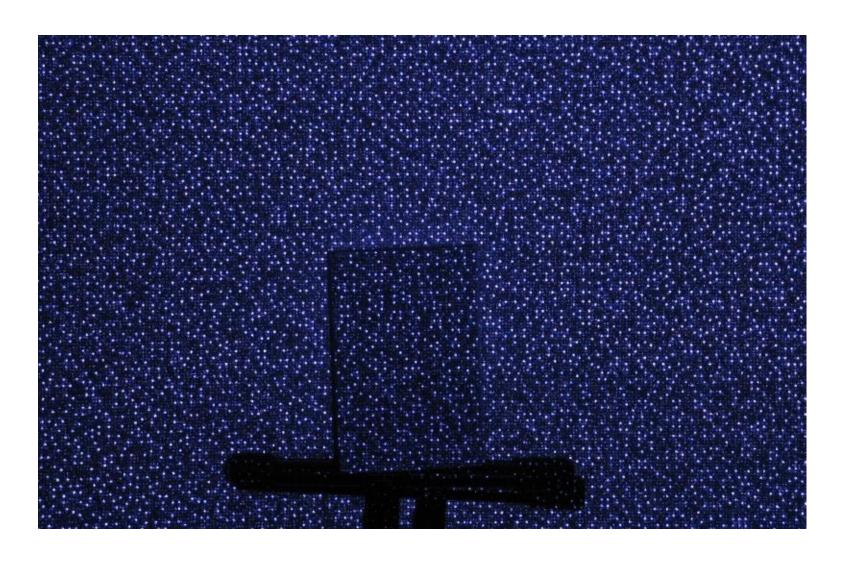
## **Active Stereo with Structured Light**



- Idea: Replace one camera by a projector.
  - Project "structured" light patterns onto the object
  - Simplifies the correspondence problem



### What the Kinect Sees...





#### 3D Reconstruction with the Kinect



#### SIGGRAPH Talks 2011

# **KinectFusion:**

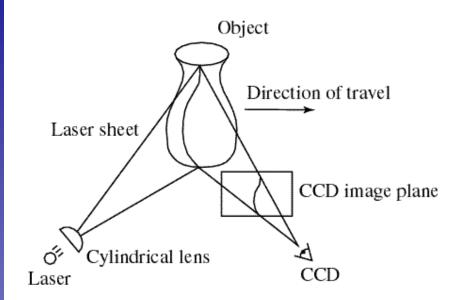
Real-Time Dynamic 3D Surface Reconstruction and Interaction

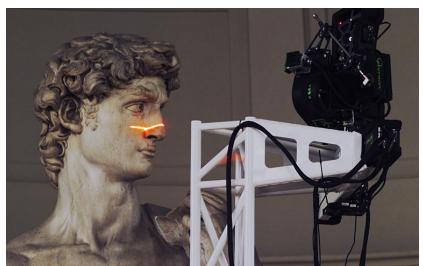
Shahram Izadi 1, Richard Newcombe 2, David Kim 1,3, Otmar Hilliges 1,
David Molyneaux 1,4, Pushmeet Kohli 1, Jamie Shotton 1,
Steve Hodges 1, Dustin Freeman 5, Andrew Davison 2, Andrew Fitzgibbon 1

1 Microsoft Research Cambridge 2 Imperial College London 3 Newcastle University 4 Lancaster University 5 University of Toronto



## **Laser Scanning**





Digital Michelangelo Project <a href="http://graphics.stanford.edu/projects/mich/">http://graphics.stanford.edu/projects/mich/</a>

#### Optical triangulation

- Project a single stripe of laser light
- Scan it across the surface of the object
- This is a very precise version of structured light scanning





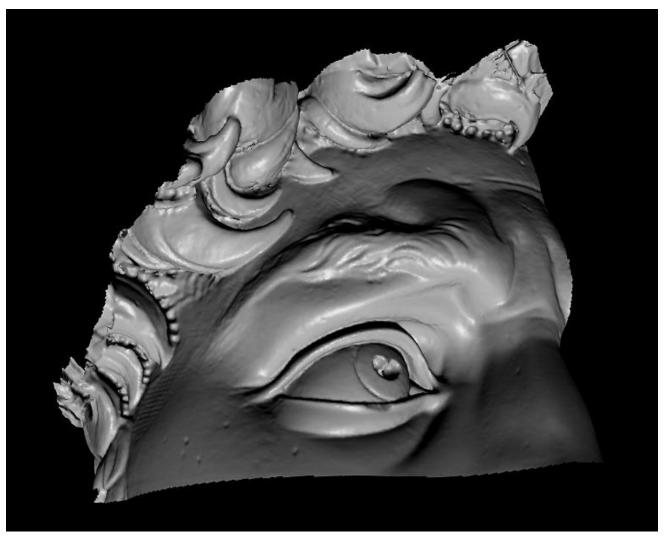
The Digital Michelangelo Project, Levoy et al.





The Digital Michelangelo Project, Levoy et al.

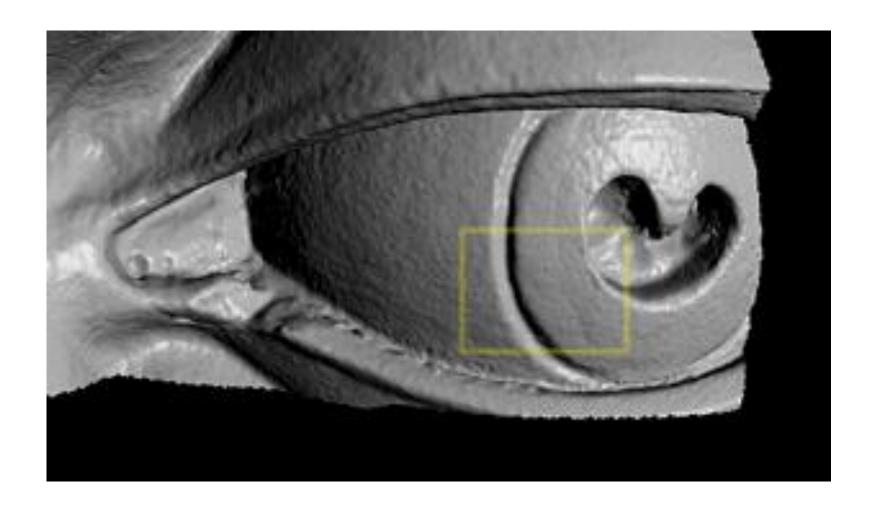




The Digital Michelangelo Project, Levoy et al.



#### **Laser Scanned Models**

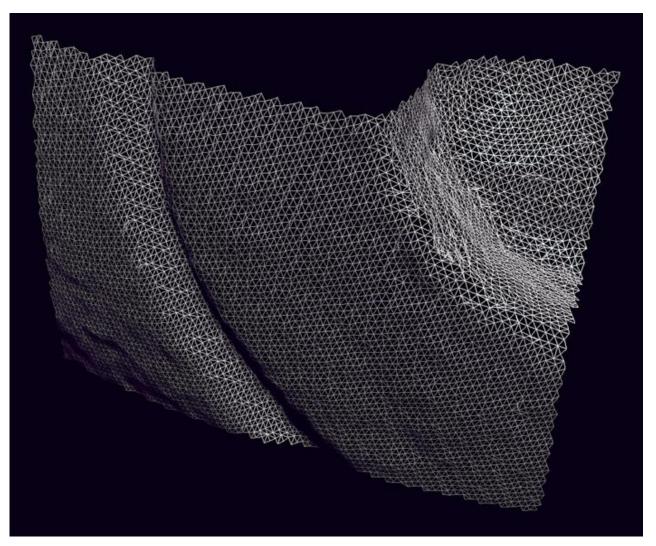


The Digital Michelangelo Project, Levoy et al.

Slide credit: Steve Seitz

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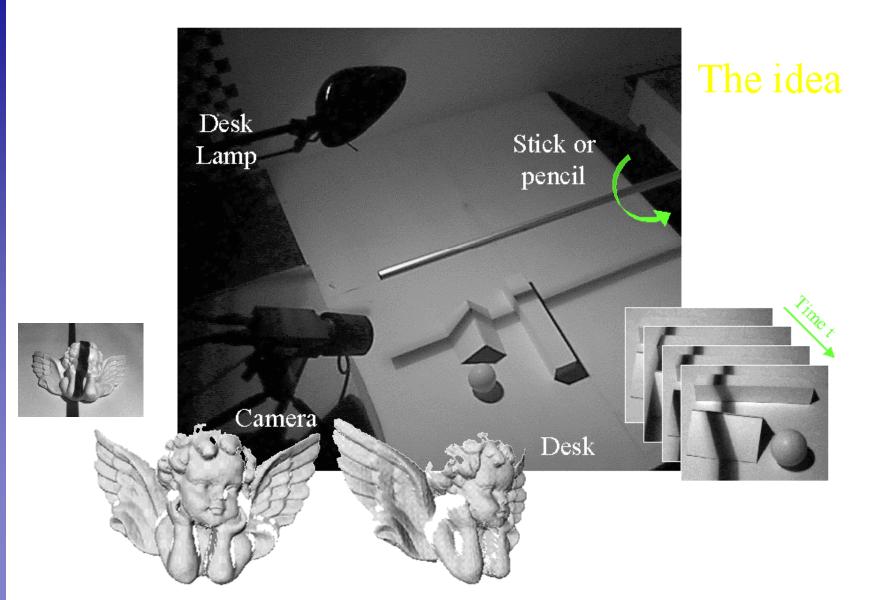




The Digital Michelangelo Project, Levoy et al.

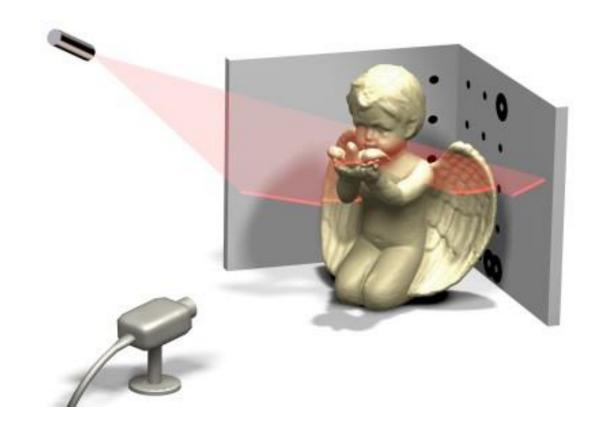


## Poor Man's Scanner



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# Slightly More Elaborate (But Still Cheap)



Software freely available from Robotics Institute TU Braunschweig <a href="http://www.david-laserscanner.com/">http://www.david-laserscanner.com/</a>



# References and Further Reading

 Background information on camera models and calibration algorithms can be found in Chapters 6 and 7 of

> R. Hartley, A. Zisserman Multiple View Geometry in Computer Vision 2nd Ed., Cambridge Univ. Press, 2004

• Also recommended: Chapter 9 of the same book on Epipolar geometry and the Fundamental Matrix and Chapter 11.1-11.6 on automatic computation of F.