RWTH AACHEN UNIVERSITY

Computer Vision – Lecture 18

Camera Calibration & 3D Reconstruction

21.01.2016

Bastian Leibe
RWTH Aachen
http://www.vision.rwth-aachen.de

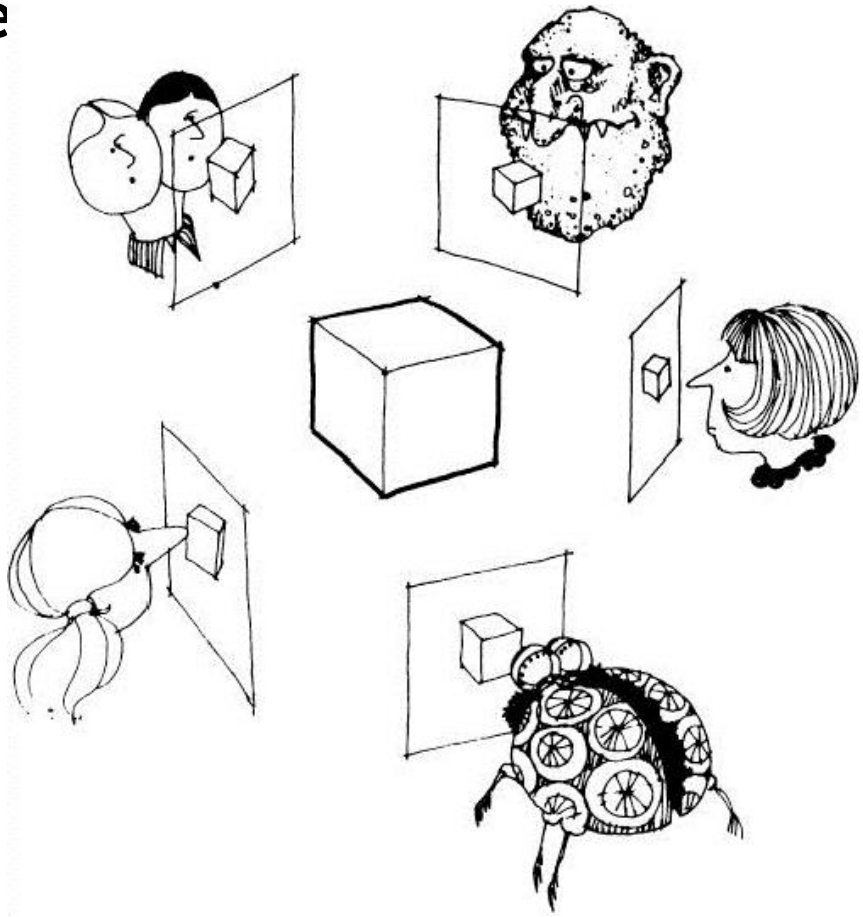
leibe@vision.rwth-aachen.de

Course Outline

- Image Processing Basics
- Segmentation & Grouping
- Object Recognition
- Local Features & Matching
- Object Categorization
- 3D Reconstruction
  - Epipolar Geometry and Stereo Basics
  - Camera calibration & Uncalibrated Reconstruction
  - Structure-from-Motion
- Motion and Tracking

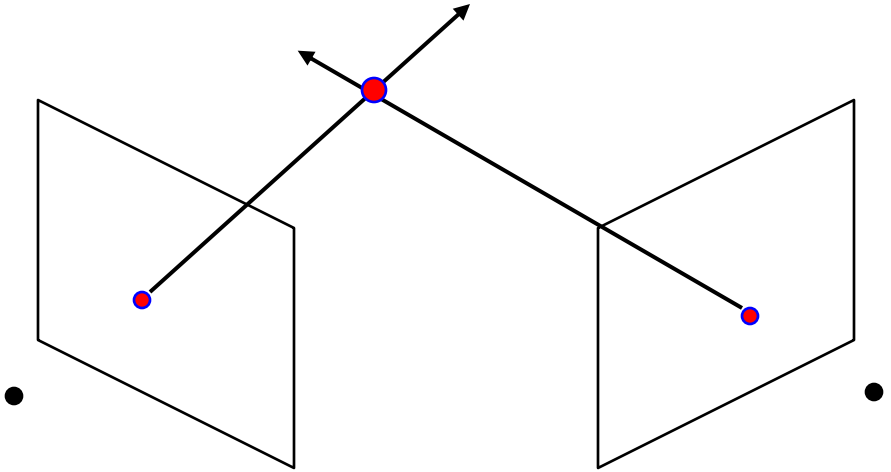
Recap: What Is Stereo Vision?

- Generic problem formulation: given several images of the same object or scene, compute a representation of its 3D shape



Slide credit: Svetlana Lazebnik, Steve Seitz

Recap: Depth with Stereo – Basic Idea

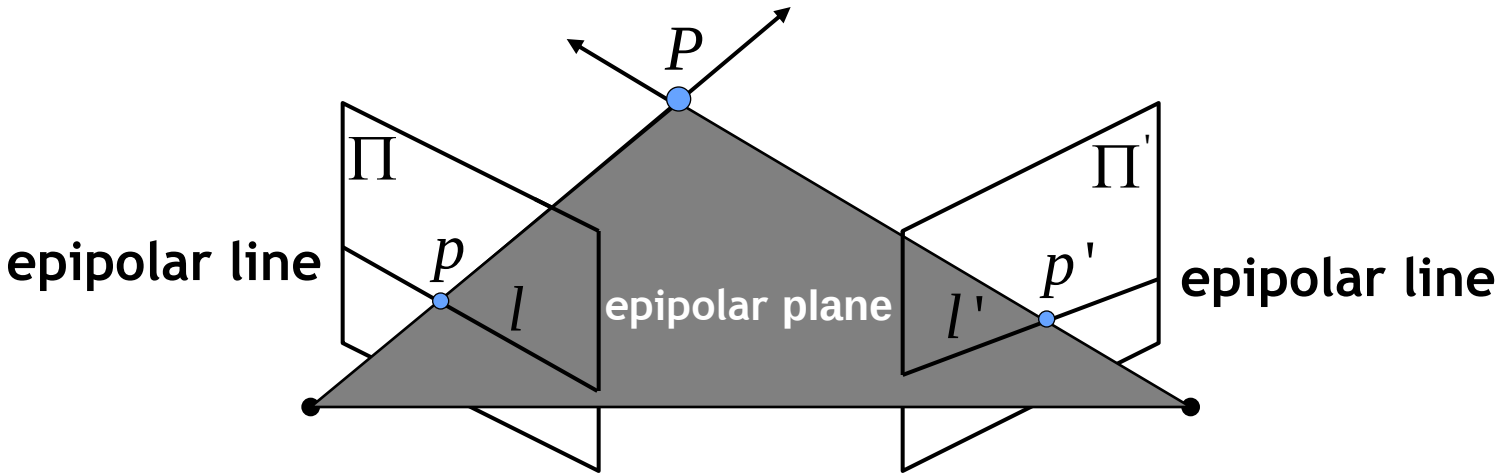


- **Basic Principle: Triangulation**
  - Gives reconstruction as intersection of two rays
  - Requires
    - Camera pose (calibration)
    - Point correspondence

Slide credit: Steve Seitz

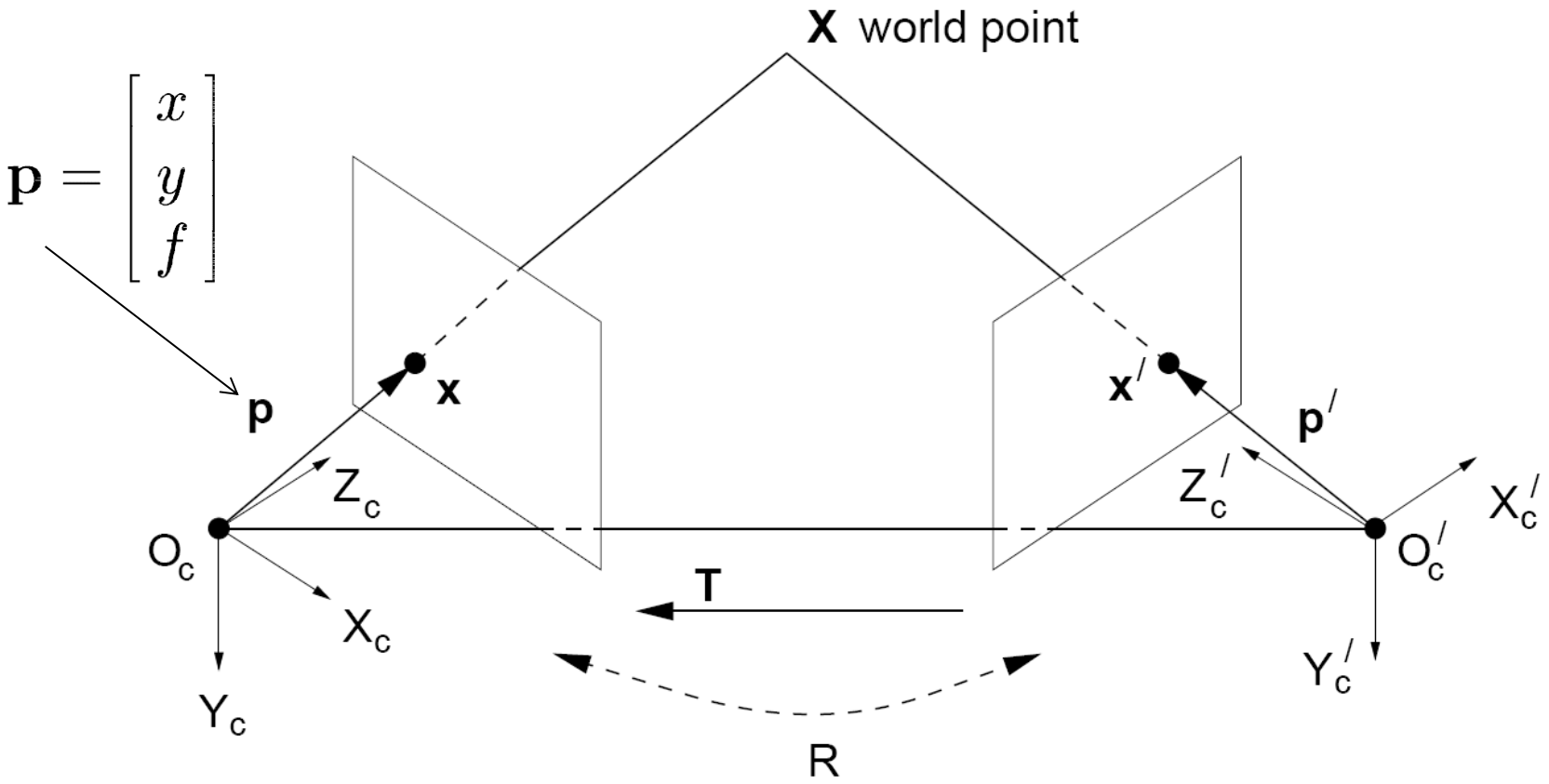
Recap: Epipolar Geometry

- Geometry of two views allows us to constrain where the corresponding pixel for some image point in the first view must occur in the second view.



- Epipolar constraint:
  - Correspondence for point p in Π must lie on the epipolar line l' in Π' (and vice versa).
  - Reduces correspondence problem to 1D search along conjugate epipolar lines.

Recap: Stereo Geometry With Calibrated Cameras



- **Camera-centered coordinate systems are related by known rotation R and translation T:**

$$\mathbf{X}' = \mathbf{R}\mathbf{X} + \mathbf{T}$$

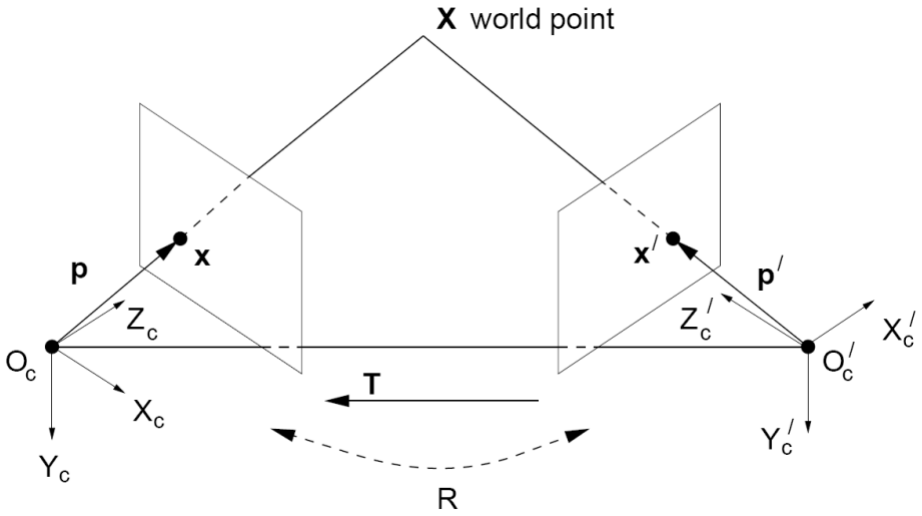
Recap: Essential Matrix

$$\mathbf{X}' \cdot \left(\mathbf{T} \times \mathbf{R}\mathbf{X}\right) = 0$$

$$\mathbf{X}' \cdot \left(\mathbf{T}_x \ \mathbf{R}\mathbf{X}\right) = 0$$

Let $\mathbf{E} = \mathbf{T}_x\mathbf{R}$

$$\mathbf{X}'^T \mathbf{E}\mathbf{X} = 0$$



- This holds for the rays p and p' that are parallel to the camera-centered position vectors X and X', so we have:

$$\mathbf{p}'^T \mathbf{E}\mathbf{p} = 0$$

- $\mathbf{E}$ is called the essential matrix, which relates corresponding image points [Longuet-Higgins 1981]

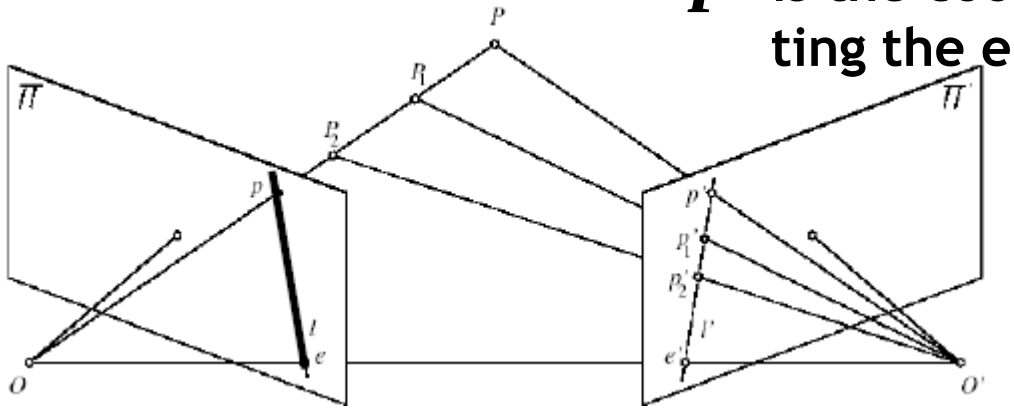
Recap: Essential Matrix and Epipolar Lines

$$\mathbf{p'}^{\mathrm{T}}\mathbf{E}\mathbf{p} = 0$$

Epipolar constraint: if we observe point p in one image, then its position p' in second image must satisfy this equation.

$l' = Ep$ is the coordinate vector representing the epipolar line for point p

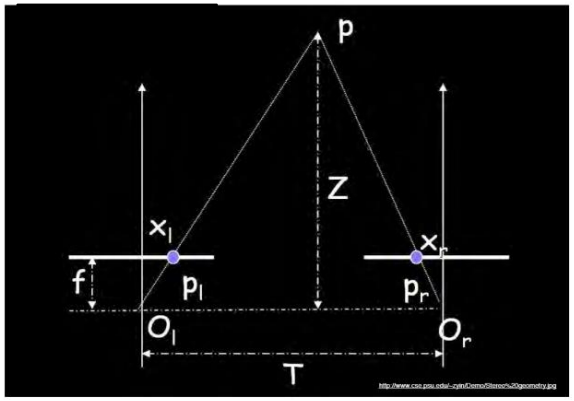
(i.e., the line is given by: $l'^{\top}\mathbf{x} = 0$)



$l = E^T p'$ is the coordinate vector representing the epipolar line for point p'

Slide credit: Kristen Grauman

Essential Matrix Example: Parallel Cameras



$$\mathbf{R} =$$

$$\mathbf{T} =$$

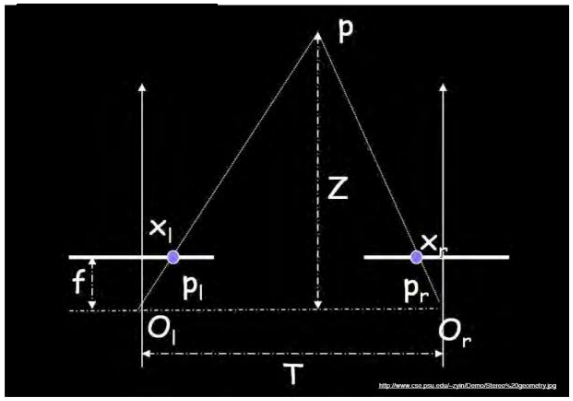
$$\mathbf{E} = [\mathbf{T}_x]\mathbf{R} =$$

$$\mathbf{p}'^{\mathrm{T}}\mathbf{E}\mathbf{p} = 0$$

For the parallel cameras, image of any point must lie on same horizontal line in each image plane.

Slide credit: Kristen Grauman

Essential Matrix Example: Parallel Cameras



$$\mathbf{R} = \mathbf{I}$$

$$\mathbf{T} = [-d, 0, 0]^{\mathrm{T}}$$

$$\mathbf{E} = [\mathbf{T}_x]\mathbf{R} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{pmatrix}$$

$$\mathbf{p}'^{\mathrm{T}}\mathbf{E}\mathbf{p} = 0$$

$$\begin{bmatrix} x' & y' & f \end{bmatrix} \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{bmatrix} \begin{bmatrix} x \\ y \\ f \end{bmatrix} = 0$$

$$\Leftrightarrow \begin{bmatrix} x' & y' & f \end{bmatrix} \begin{bmatrix} 0 \\ df \\ -dy \end{bmatrix} = 0$$

$$\Leftrightarrow y = y'$$

For the parallel cameras, image of any point must lie on same horizontal line in each image plane.

Slide credit: Kristen Grauman

More General Case

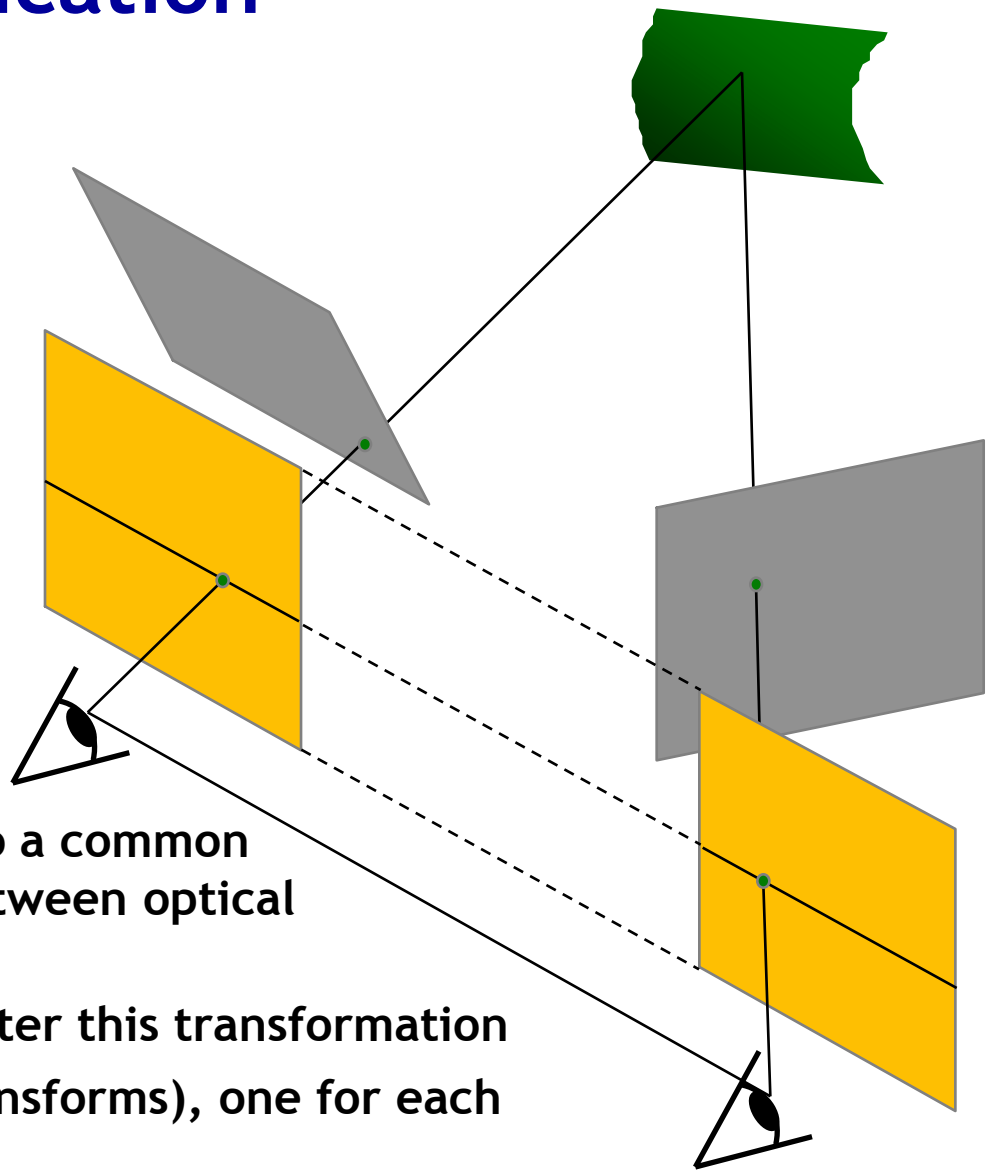
Image $I(x,y)$ **Disparity map** $D(x,y)$ **Image** $I'(x',y')$

$$(x', y') = (x + D(x, y), y)$$

What about when cameras' optical axes are not parallel?

Stereo Image Rectification

- In practice, it is convenient if image scanlines are the epipolar lines.

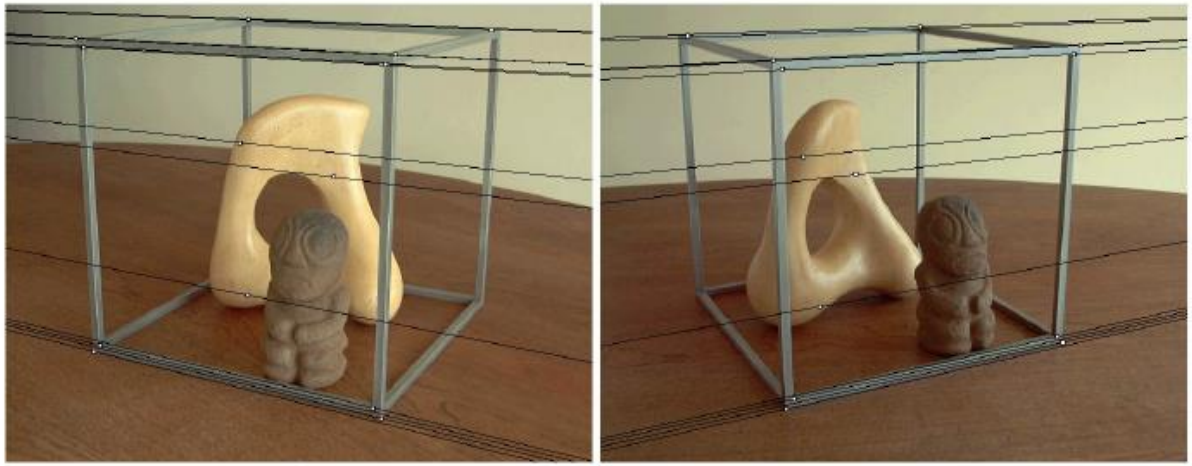


- Algorithm
  - Reproject image planes onto a common plane parallel to the line between optical centers
  - Pixel motion is horizontal after this transformation
  - Two homographies (3×3 transforms), one for each input image reprojection

Slide adapted from Li Zhang

C. Loop & Z. Zhang, Computing Rectifying Homographies for Stereo Vision. CVPR'99

Stereo Image Rectification: Example



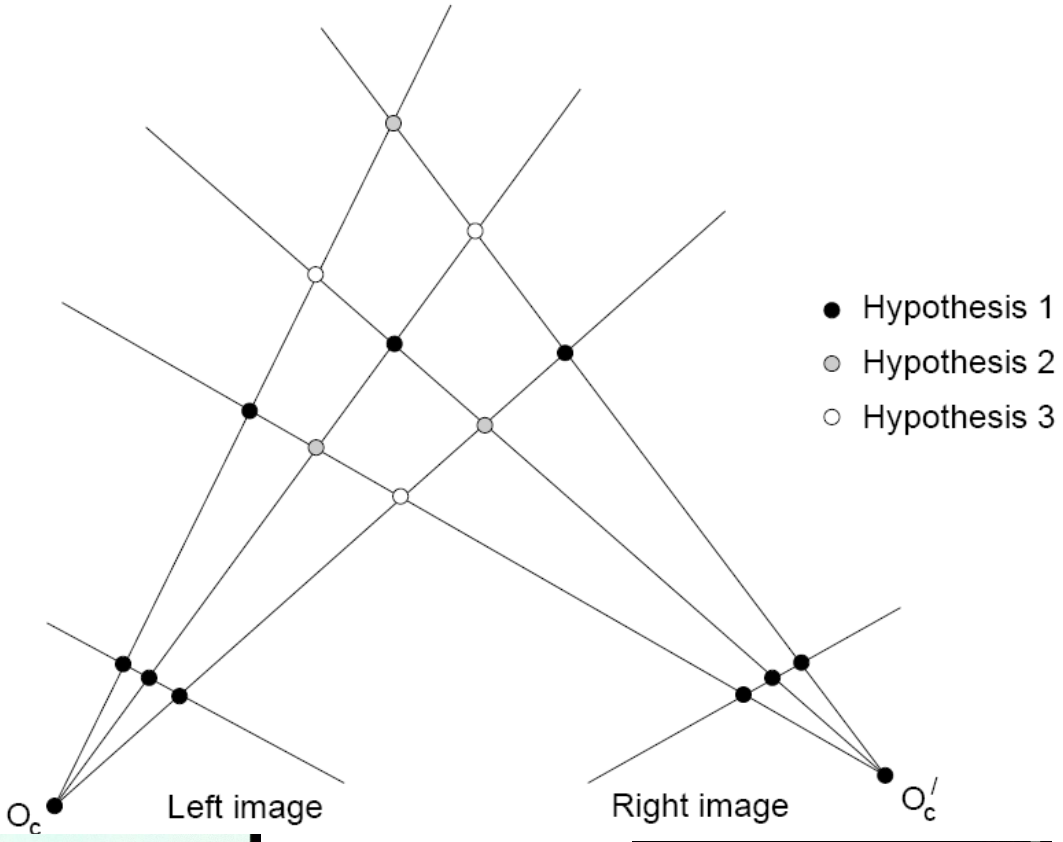
Source: Alyosha Efros

Recap: Stereo Reconstruction

- Main Steps
  - Calibrate cameras
  - Rectify images
  - Compute disparity
  - Estimate depth

Slide credit: Kristen Grauman

Correspondence Problem



Multiple match hypotheses satisfy the epipolar constraint, but which is the correct one?

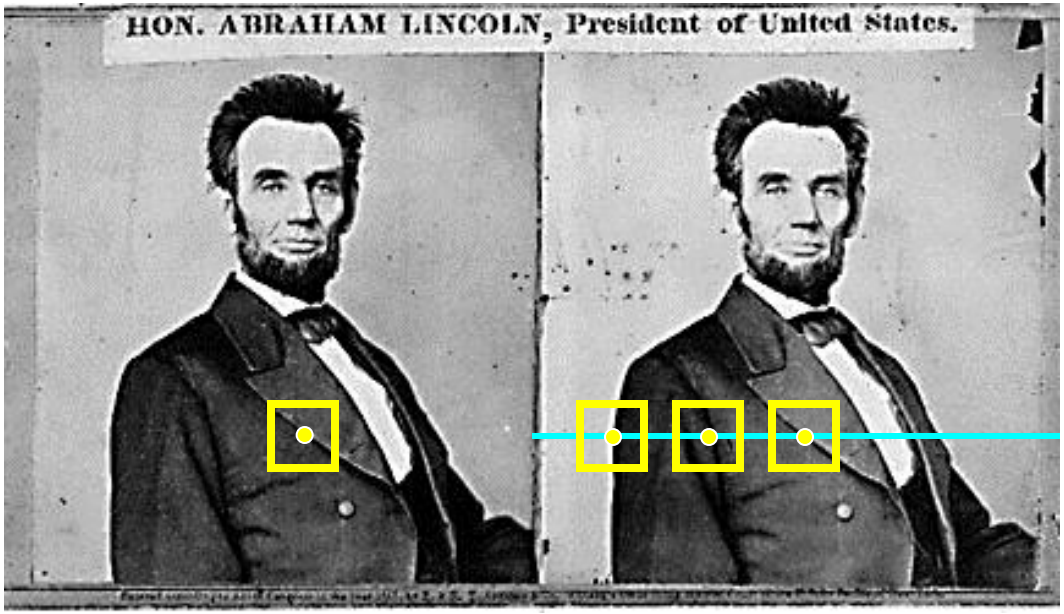
⇒ ***This ambiguity is what makes stereo estimation difficult!***

Slide credit: Kristen Grauman

B. Leibe

Figure from Gee & Cipolla 1999

Dense Correspondence Search



- For each pixel in the first image
  - Find corresponding epipolar line in the right image
  - Examine all pixels on the epipolar line and pick the best match (e.g. SSD, correlation)
  - Triangulate the matches to get depth information
- This is easiest when epipolar lines are scanlines

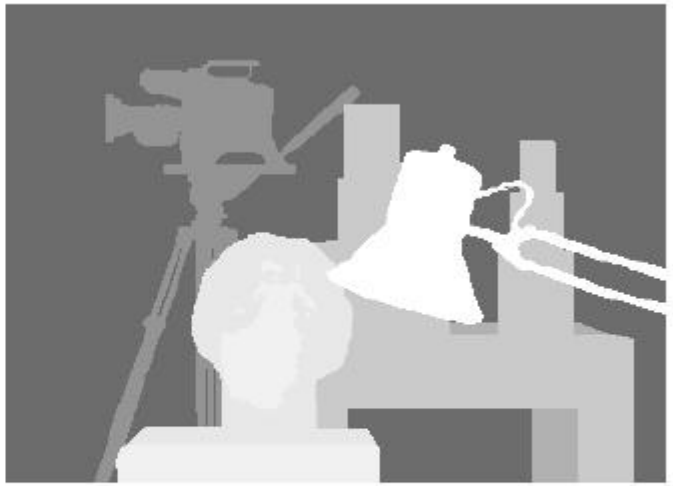
  $\Rightarrow$ Rectify images first

adapted from Svetlana Lazebnik, Li Zhang

Example: Window Search

- **Data from University of Tsukuba**

Scene

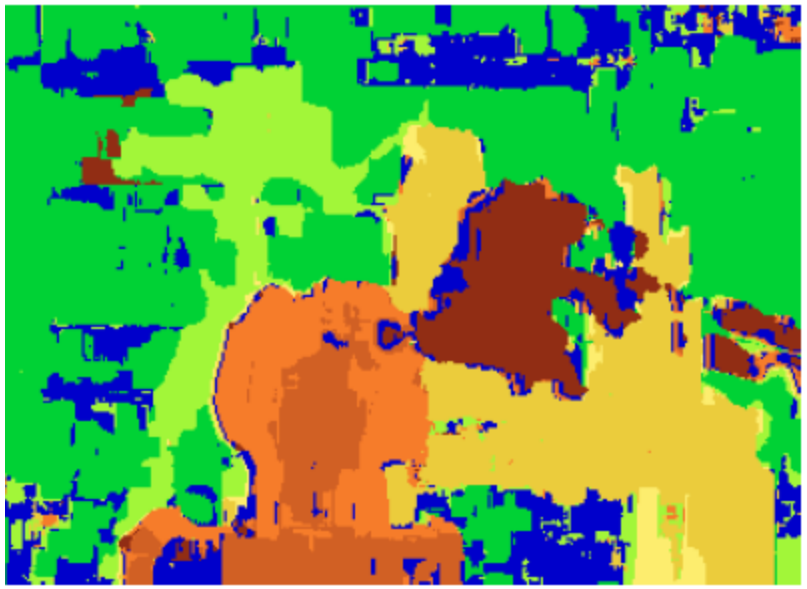


Ground truth

Slide credit: Kristen Grauman

Example: Window Search

- **Data from University of Tsukuba**



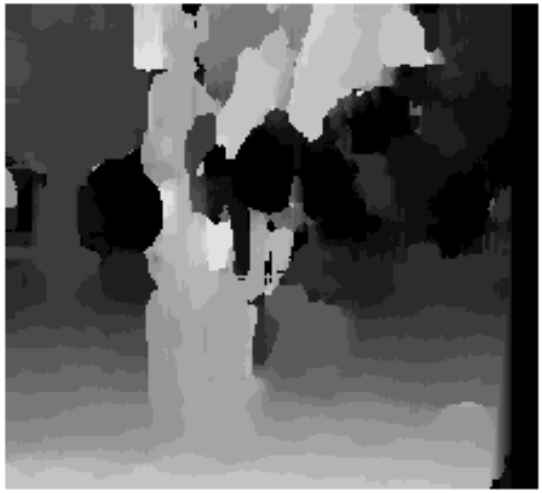
Window-based matching
(best window size)

Ground truth

Slide credit: Kristen Grauman

Effect of Window Size

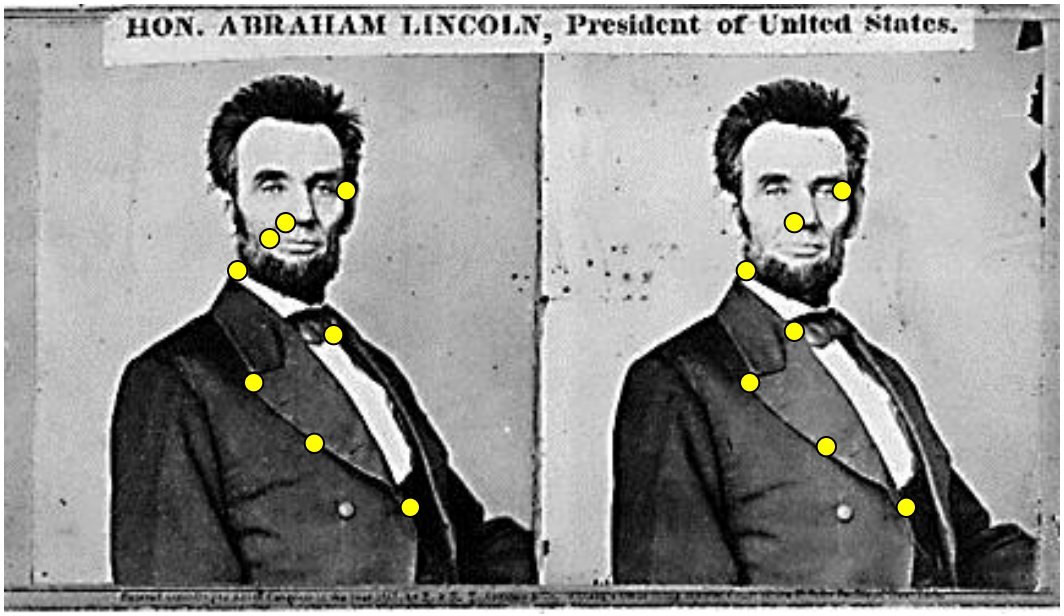
$W = 3$



$W = 20$

Want window large enough to have sufficient intensity variation, yet small enough to contain only pixels with about the same disparity.

Alternative: Sparse Correspondence Search



- Idea:
  - Restrict search to sparse set of detected features
  - Rather than pixel values (or lists of pixel values) use *feature descriptor* and an associated *feature distance*
  - Still narrow search further by epipolar geometry

What would make good features?

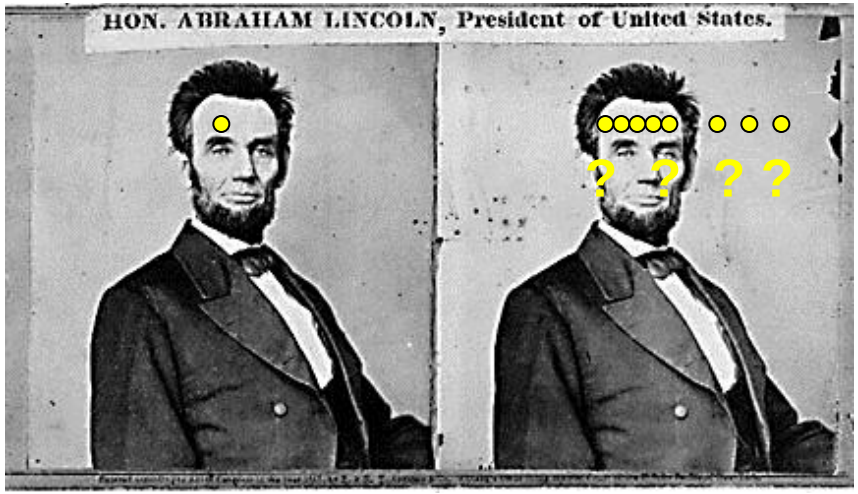
Slide credit: Kristen Grauman

Dense vs. Sparse

- Sparse
  - Efficiency
  - Can have more reliable feature matches, less sensitive to illumination than raw pixels
  - But…
    - Have to know enough to pick good features
    - Sparse information
- Dense
  - Simple process
  - More depth estimates, can be useful for surface reconstruction
  - But…
    - Breaks down in textureless regions anyway
    - Raw pixel distances can be brittle
    - Not good with very different viewpoints

Slide credit: Kristen Grauman

Difficulties in Similarity Constraint



Untextured surfaces

Occlusions

Slide credit: Kristen Grauman

Possible Sources of Error?

- Low-contrast / textureless image regions
- Occlusions
- Camera calibration errors
- Violations of *brightness constancy* (e.g., specular reflections)
- Large motions

Slide credit: Kristen Grauman

Summary: Stereo Reconstruction

- **Main Steps**
  - Calibrate cameras
  - Rectify images
  - Compute disparity
  - Estimate depth

Left

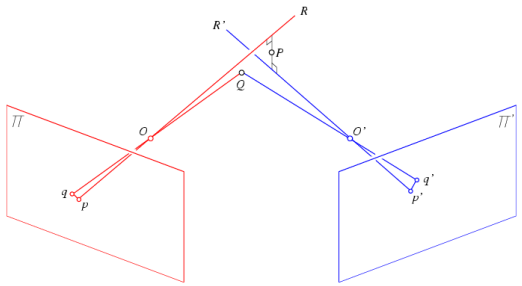
Right

- **So far, we have only considered calibrated cameras…**

Left

Right

- **Today**
  - Uncalibrated cameras
  - Camera parameters
  - Revisiting epipolar geometry
  - Robust fitting



Slide credit: Kristen Grauman

Recap: A General Point

- Equations of the form

$$Ax = 0$$

- How do we solve them? (always!)
  - Apply SVD

$$\underset{\text{SVD}}{\downarrow}$$

$$\boldsymbol{A} = \boldsymbol{U}\boldsymbol{D}\boldsymbol{V}^T = \boldsymbol{U}\begin{bmatrix} d_{11} & & \\ & \ddots & \\ & & d_{NN} \end{bmatrix}\begin{bmatrix} v_{11} & \cdots & v_{1N} \\ \vdots & \ddots & \vdots \\ v_{N1} & \cdots & v_{NN} \end{bmatrix}^T$$

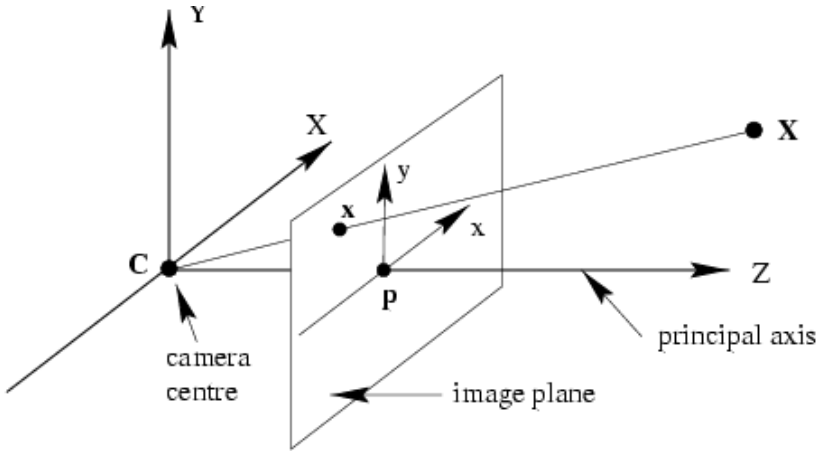
Singular values Singular vectors

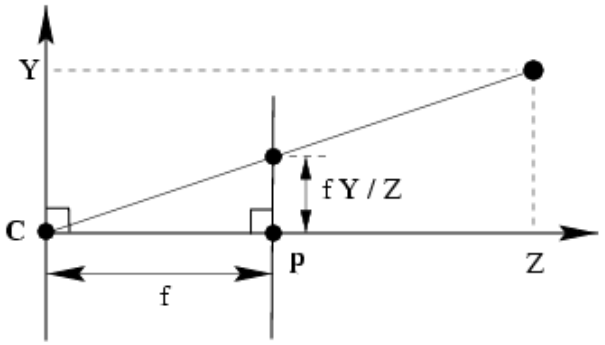
  - Singular values of A = square roots of the eigenvalues of A^TA.
  - The solution of Ax=0 is the *nullspace* vector of A.
  - This corresponds to the *smallest singular vector* of A.

Topics of This Lecture

- **Camera Calibration**
  - Camera parameters
  - Calibration procedure
- **Revisiting Epipolar Geometry**
  - Triangulation
  - Calibrated case: Essential matrix
  - Uncalibrated case: Fundamental matrix
  - Weak calibration
  - Epipolar Transfer
- **Active Stereo**
  - Laser scanning
  - Kinect sensor

Recall: Pinhole Camera Model



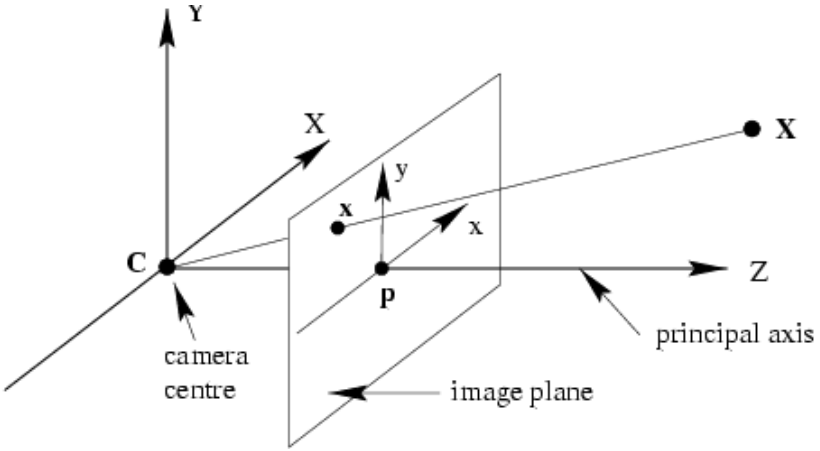


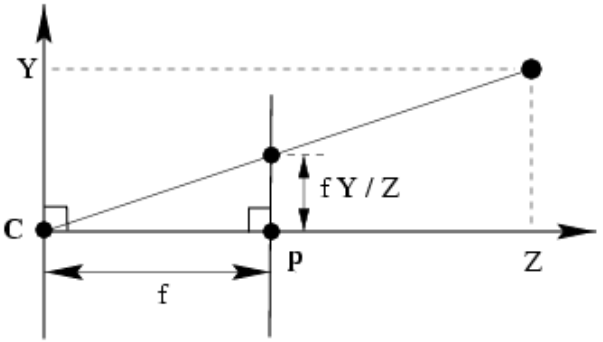
$$(X, Y, Z) \mapsto (f\,X/Z,\ f\,Y/Z)$$

$$\begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix} \mapsto \begin{pmatrix} f\,X \\ f\,Y \\ Z \end{pmatrix} = \begin{bmatrix} f & & & 0 \\ & f & & 0 \\ & & 1 & 0 \end{bmatrix} \begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix}$$

$$\mathbf{x} = \mathbf{P}\mathbf{X}$$

Pinhole Camera Model

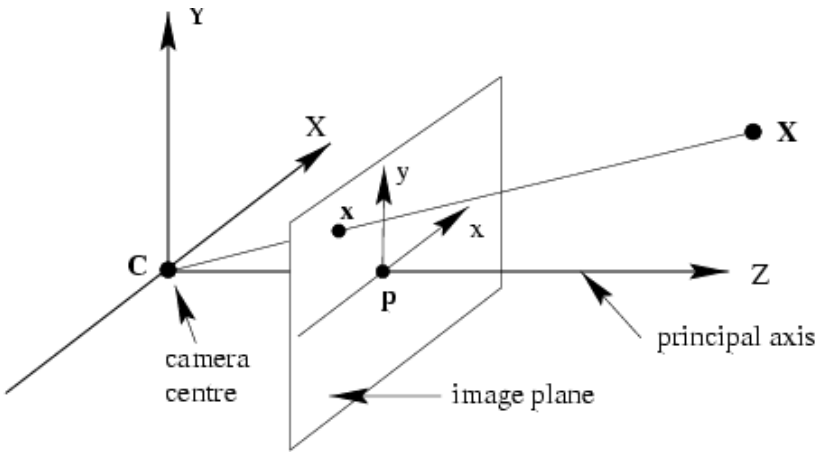




$$\begin{pmatrix} fX \\ fY \\ Z \end{pmatrix} = \begin{bmatrix} f & & & \\ & f & & \\ & & 1 & \end{bmatrix} \begin{bmatrix} 1 & & & 0 \\ & 1 & & 0 \\ & & 1 & 0 \end{bmatrix} \begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix}$$

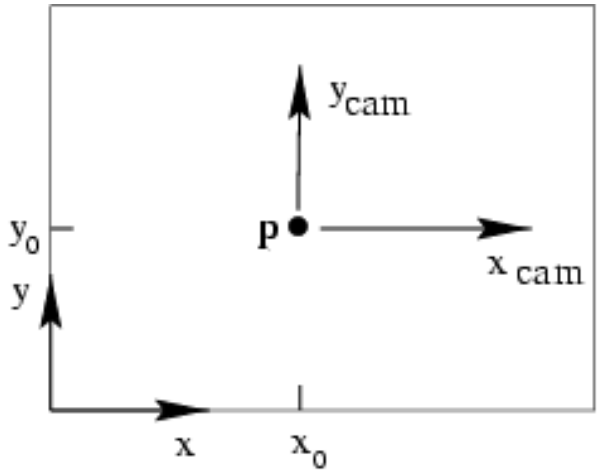
$$\mathrm{x} = \mathrm{PX} \qquad \mathrm{P} = \mathrm{diag}(f, f, 1)[\mathrm{I} \mid 0]$$

Camera Coordinate System



- *Principal axis*:
  - Line from the camera center perpendicular to the image plane
- *Normalized (camera) coordinate system*:
  - Camera center is at the origin and the principal axis is the z-axis
- *Principal point* (p):
  - Point where principal axis intersects the image plane (origin of normalized coordinate system)

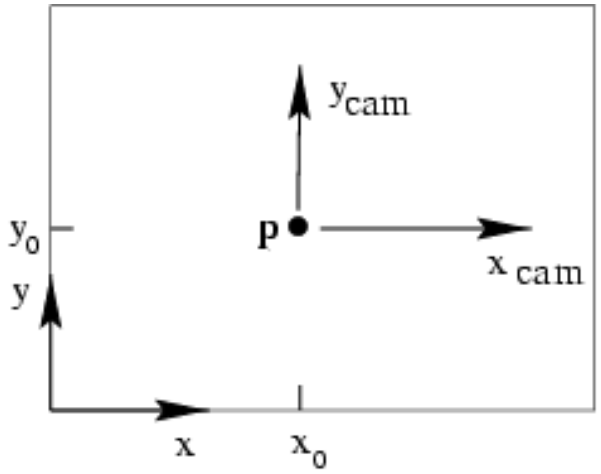
Principal Point Offset



principal point: (p_x, p_y)

- Camera coordinate system: origin at the principal point
- Image coordinate system: origin is in the corner

Principal Point Offset



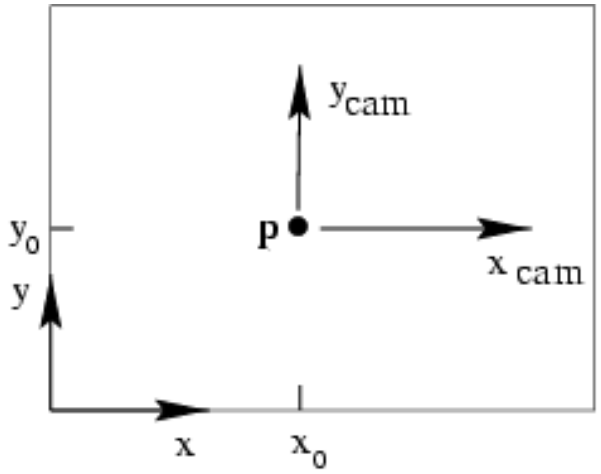
principal point: (p_x, p_y)

$$(X, Y, Z) \mapsto (f\,X/Z + p_x,\ f\,Y/Z + p_y)$$

$$\begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix} \mapsto \begin{pmatrix} f\,X + Z\,p_x \\ f\,Y + Z\,p_x \\ Z \end{pmatrix} = \begin{bmatrix} f & & p_x & 0 \\ & f & p_y & 0 \\ & & 1 & 0 \end{bmatrix} \begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix}$$

Slide credit: Svetlana Lazebnik
Image from Hartley & Zisserman

Principal Point Offset

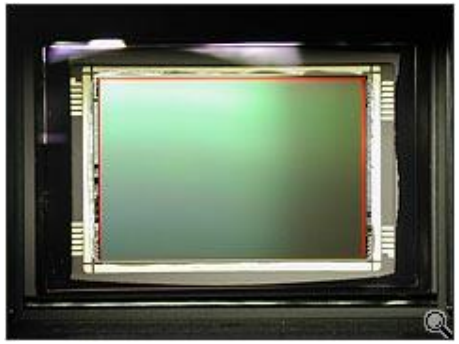


principal point: (p_x, p_y)

$$\begin{pmatrix} fX + Zp_x \\ fY + Zp_x \\ Z \end{pmatrix} = \begin{bmatrix} f & & p_x \\ & f & p_y \\ & & 1 \end{bmatrix} \begin{bmatrix} 1 & & & 0 \\ & 1 & & 0 \\ & & 1 & 0 \end{bmatrix} \begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix}$$

$$K = \begin{bmatrix} f & & p_x \\ & f & p_y \\ & & 1 \end{bmatrix}$$ **calibration matrix** $$\mathrm{P} = \mathrm{K}[\mathrm{I} \mid 0]$$

Pixel Coordinates: Non-Square Pixels



Pixel size: $\frac{1}{m_x} \times \frac{1}{m_y}$

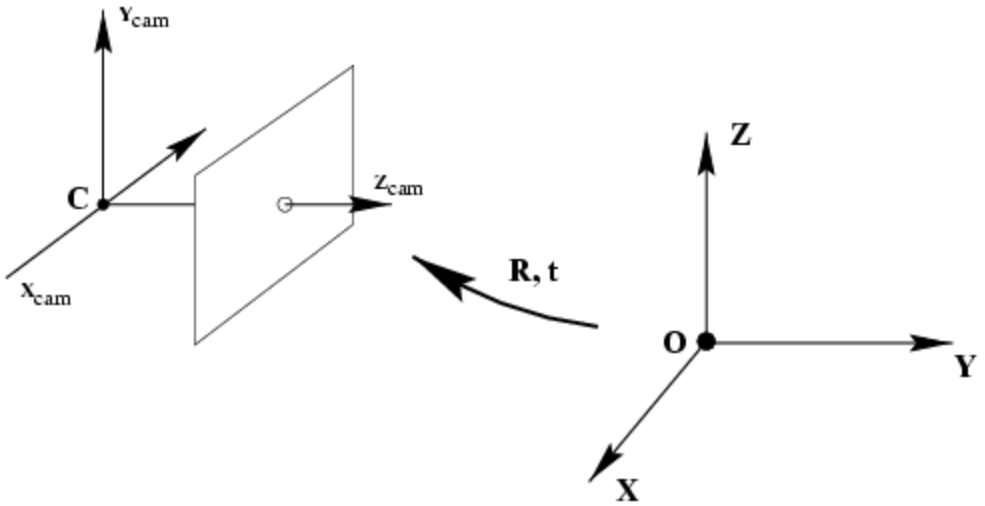
m_x pixels per meter in horizontal direction,
m_y pixels per meter in vertical direction

$$K = \begin{bmatrix} m_x & & \\ & m_y & \\ & & 1 \end{bmatrix} \begin{bmatrix} f & & p_x \\ & f & p_y \\ & & 1 \end{bmatrix} = \begin{bmatrix} \alpha_x & & x_0 \\ & \alpha_y & y_0 \\ & & 1 \end{bmatrix}$$

[pixels/m] **[m]** **[pixels]**

Slide credit: Svetlana Lazebnik

Camera Rotation and Translation



- In general, the camera coordinate frame will be related to the world coordinate frame by a rotation and a translation

$$\tilde{X}_{cam} = R\left(\tilde{X} - \tilde{C}\right)$$

coords. of point in camera frame

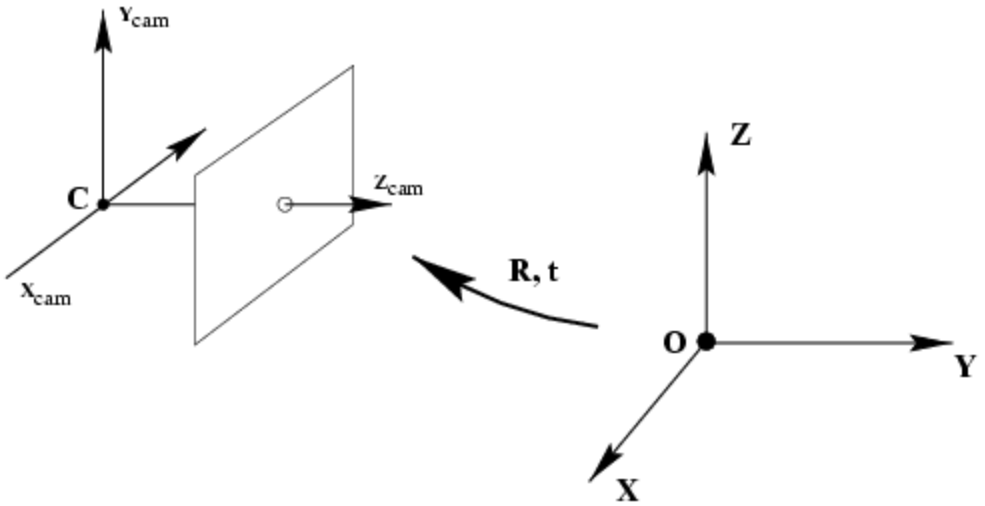
coords. of a point in world frame (nonhomogeneous)

coords. of camera center in world frame

Slide credit: Svetlana Lazebnik

Image from Hartley & Zisserman

Camera Rotation and Translation



In non-homogeneous coordinates:

$$\tilde{X}_{cam} = R\left(\tilde{X} - \tilde{C}\right)$$

$$X_{cam} = \begin{bmatrix} R & -R\tilde{C} \\ 0 & 1 \end{bmatrix} \begin{pmatrix} \tilde{X} \\ 1 \end{pmatrix} = \begin{bmatrix} R & -R\tilde{C} \\ 0 & 1 \end{bmatrix} X$$

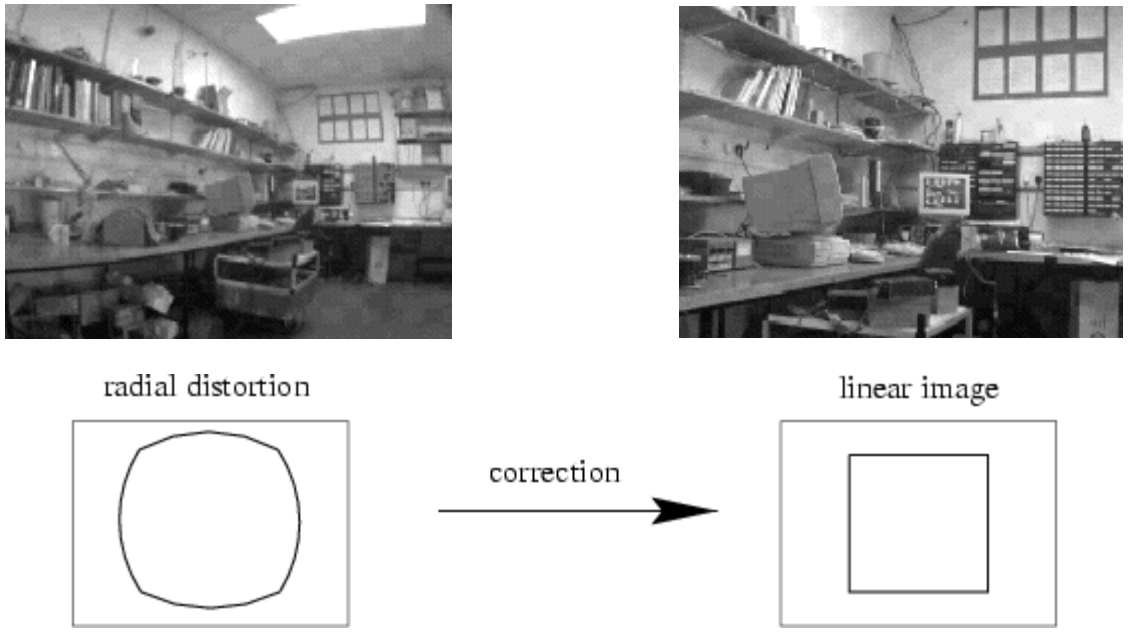
$$x = K\left[I \mid 0\right] X_{cam} = K\left[R \mid -R\tilde{C}\right] X \qquad P = K\left[R \mid t\right], \quad t = -R\tilde{C}$$

Note: $\mathbf{C}$ is the null space of the camera projection matrix ($\mathbf{PC}=0$)

Summary: Camera Parameters

- **Intrinsic parameters**
  - Principal point coordinates
  - Focal length
  - Pixel magnification factors
  - *Skew (non-rectangular pixels)*
  - *Radial distortion*

$$K = \begin{bmatrix} m_x & & \\ & m_y & \\ & & 1 \end{bmatrix} \begin{bmatrix} f & s & p_x \\ & f & p_y \\ & & 1 \end{bmatrix} = \begin{bmatrix} \alpha_x & s & x_0 \\ & \alpha_y & y_0 \\ & & 1 \end{bmatrix}$$



Slide credit: Svetlana Lazebnik

Summary: Camera Parameters

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$$K = \begin{bmatrix} m_x & & \\ & m_y & \\ & & 1 \end{bmatrix} \begin{bmatrix} f & s & p_x \\ & f & p_y \\ & & 1 \end{bmatrix} = \begin{bmatrix} \alpha_x & s & x_0 \\ & \alpha_y & y_0 \\ & & 1 \end{bmatrix}$$

- **Extrinsic parameters**
  - Rotation $\mathbf{R}$
  - Translation $\mathbf{t}$
    (both relative to world coordinate system)

- **Camera projection matrix**

$$\mathrm{P} = \mathrm{K}\left[\mathrm{R} \mid \mathrm{t}\right] = \begin{bmatrix} P_{11} & P_{12} & P_{13} & P_{14} \\ P_{21} & P_{22} & P_{23} & P_{24} \\ P_{31} & P_{32} & P_{33} & P_{34} \end{bmatrix}$$

How many degrees of freedom does P have?

Slide adapted from Svetlana Lazebnik

Camera Parameters: Degrees of Freedom

- **Intrinsic parameters** — **DoF**
  - Principal point coordinates — 2
  - Focal length — 1
  - Pixel magnification factors — 1
  - *Skew (non-rectangular pixels)* — 1
  - *Radial distortion*

$$K = \begin{bmatrix} f\alpha_x & s & px_0 \\ & f\alpha_y & py_0 \\ & & 1 \end{bmatrix}$$

- **Extrinsic parameters**
  - Rotation R — 3
  - Translation t — 3
    (both relative to world coordinate system)

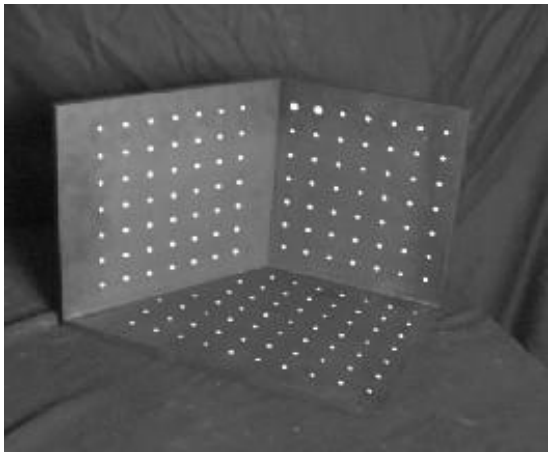
- **Camera projection matrix**

$$P = K[R \mid t]$$

  - ⇒ General pinhole camera: 9 DoF
  - ⇒ CCD Camera with square pixels: 10 DoF
  - ⇒ General camera: 11 DoF

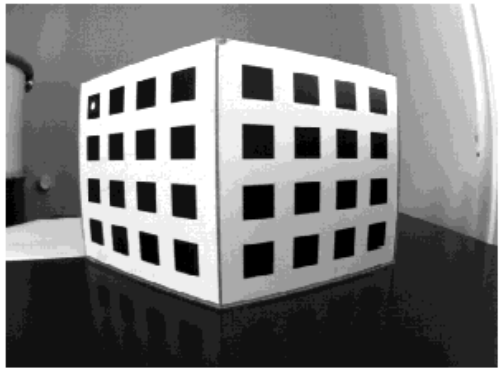
Calibrating a Camera

- Compute intrinsic and extrinsic parameters using observed camera data.



Main idea

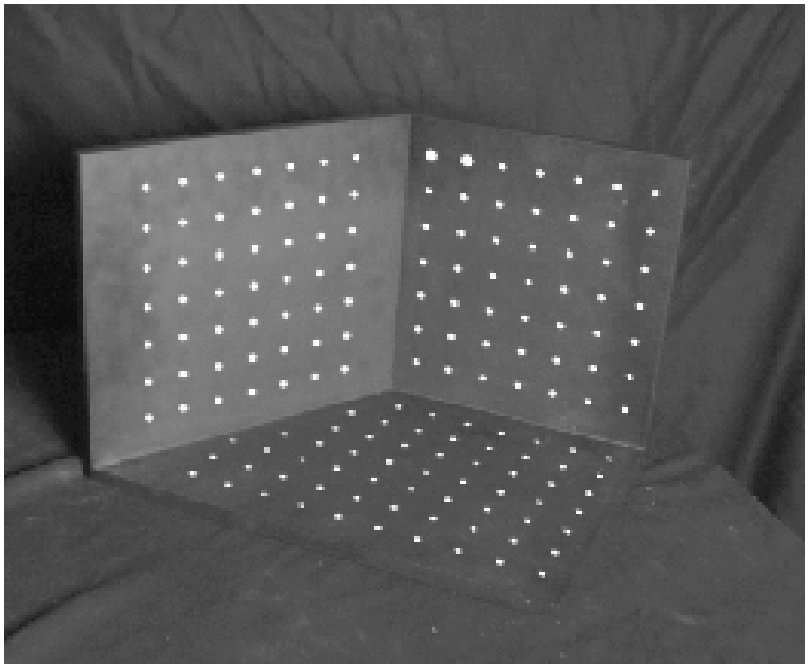
- Place "calibration object" with known geometry in the scene
- Get correspondences
- Solve for mapping from scene to image: estimate $P=P_{int}P_{ext}$

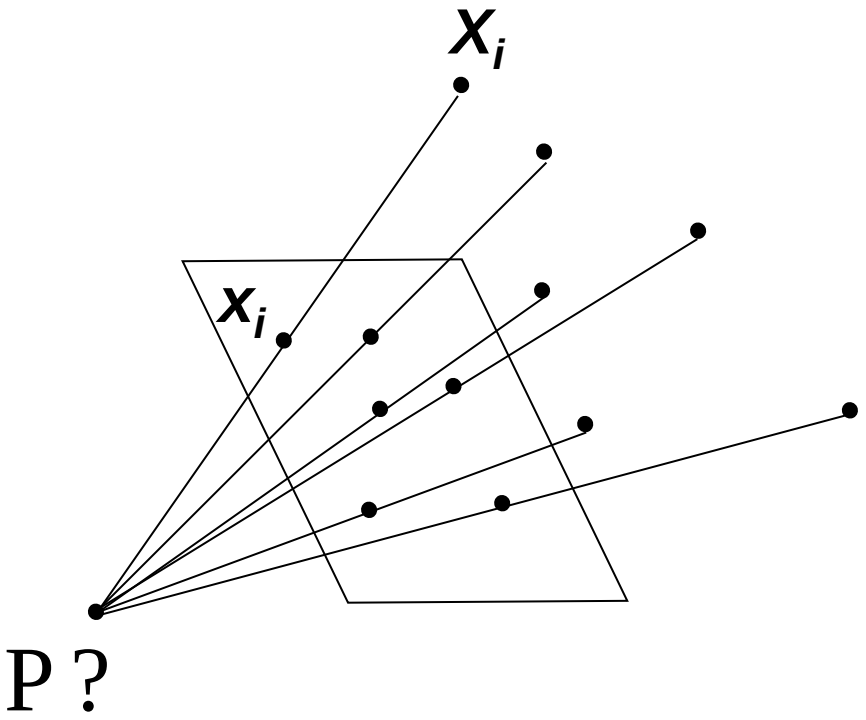


Slide credit: Kristen Grauman

Camera Calibration

- Given n points with known 3D coordinates X_i and known image projections x_i, estimate the camera parameters



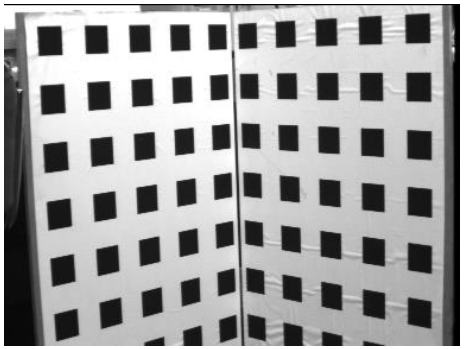


Slide credit: Svetlana Lazebnik

Camera Calibration: Obtaining the Points

- For best results, it is important that the calibration points are measured with *subpixel accuracy*.
- How this can be done depends on the exact pattern.

- Algorithm for checkerboard pattern
  1. Perform Canny edge detection.
  2. Fit straight lines to detected linked edges.
  3. Intersect lines to obtain corners.
  - If sufficient care is taken, the points can then be obtained with localization accuracy < 1/10 pixel.



- Rule of thumb
  - Number of constraints should exceed number of unknowns by a factor of five.

  ⇒ For 11 parameters of P, at least 28 points should be used.

Camera Calibration: DLT Algorithm

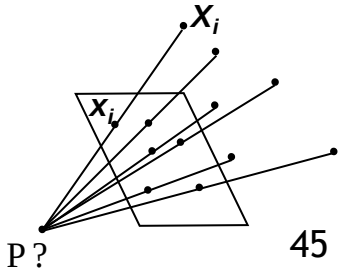
(DLT = "Direct Linear Transform")

$$\lambda \mathrm{x}_i = \mathrm{P}\mathrm{X}_i \qquad \lambda \begin{bmatrix} x_i \\ y_i \\ 1 \end{bmatrix} = \begin{bmatrix} P_{11} & P_{12} & P_{13} & P_{14} \\ P_{21} & P_{22} & P_{23} & P_{24} \\ P_{31} & P_{32} & P_{33} & P_{34} \end{bmatrix} \begin{bmatrix} \mathrm{X}_{i,1} \\ \mathrm{X}_{i,2} \\ \mathrm{X}_{i,3} \\ 1 \end{bmatrix} = \begin{bmatrix} \mathrm{P}_1^T \\ \mathrm{P}_2^T \\ \mathrm{P}_3^T \end{bmatrix} \mathrm{X}_i$$

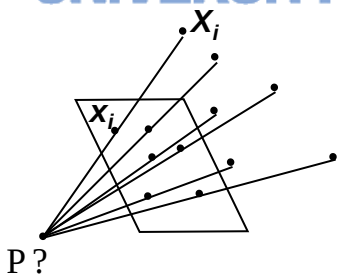
$$\mathrm{x}_i \times \mathrm{P}\mathrm{X}_i = 0$$

$$\begin{bmatrix} 0 & -\mathrm{X}_i^T & y_i \mathrm{X}_i^T \\ \mathrm{X}_i^T & 0 & -x_i \mathrm{X}_i^T \\ -y_i \mathrm{X}_i^T & x_i \mathrm{X}_i^T & 0 \end{bmatrix} \begin{pmatrix} \mathrm{P}_1 \\ \mathrm{P}_2 \\ \mathrm{P}_3 \end{pmatrix} = 0$$

Only two linearly independent equations



Camera Calibration: DLT Algorithm



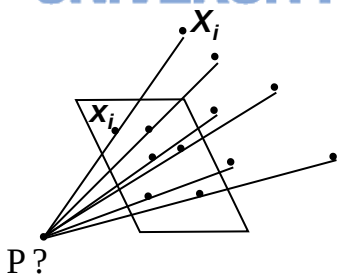
$$\begin{bmatrix} 0^T & \mathrm{X}_1^T & -y_1\mathrm{X}_1^T \\ \mathrm{X}_1^T & 0^T & -x_1\mathrm{X}_1^T \\ \cdots & \cdots & \cdots \\ 0^T & \mathrm{X}_n^T & -y_n\mathrm{X}_n^T \\ \mathrm{X}_n^T & 0^T & -x_n\mathrm{X}_n^T \end{bmatrix} \begin{pmatrix} \mathrm{P}_1 \\ \mathrm{P}_2 \\ \mathrm{P}_3 \end{pmatrix} = 0 \qquad \mathrm{Ap} = 0$$

Solve using… SVD!

- **Notes**
  - **P has 11 degrees of freedom (12 parameters, but scale is arbitrary).**
  - **One 2D/3D correspondence gives us two linearly independent equations.**
  - **Homogeneous least squares (similar to homography est.)**

⇒ 5 ½ correspondences needed for a minimal solution.

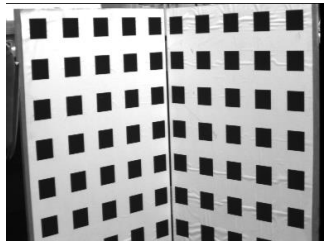
Camera Calibration: DLT Algorithm



$$\begin{bmatrix} 0^T & \mathrm{X}_1^T & -y_1\mathrm{X}_1^T \\ \mathrm{X}_1^T & 0^T & -x_1\mathrm{X}_1^T \\ \cdots & \cdots & \cdots \\ 0^T & \mathrm{X}_n^T & -y_n\mathrm{X}_n^T \\ \mathrm{X}_n^T & 0^T & -x_n\mathrm{X}_n^T \end{bmatrix} \begin{pmatrix} \mathrm{P}_1 \\ \mathrm{P}_2 \\ \mathrm{P}_3 \end{pmatrix} = 0 \qquad \mathrm{Ap} = 0$$

- **Notes**
  - **For coplanar points that satisfy $\Pi^T\mathrm{X}=0$, we will get degenerate solutions $(\Pi,0,0)$, $(0,\Pi,0)$, or $(0,0,\Pi)$.**

$\Rightarrow$ We need calibration points in more than one plane!



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Camera Calibration

- Once we've recovered the numerical form of the camera matrix, we still have to figure out the intrinsic and extrinsic parameters
- This is a matrix decomposition problem, not an estimation problem (see F&P sec. 3.2, 3.3)

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Camera Calibration: Some Practical Tips

- For numerical reasons, it is important to carry out some data normalization.
  - Translate the image points x_i to the (image) origin and scale them such that their RMS distance to the origin is $\sqrt{2}$.
  - Translate the 3D points X_i to the (world) origin and scale them such that their RMS distance to the origin is $\sqrt{3}$.
  - (This is valid for compact point distributions on calibration objects).
- The DLT algorithm presented here is easy to implement, but there are some more accurate algorithms available (see H&Z sec. 7.2).
- For practical applications, it is also often needed to correct for radial distortion. Algorithms for this can be found in H&Z sec. 7.4, or F&P sec. 3.3.

Topics of This Lecture

- Camera Calibration
  - Camera parameters
  - Calibration procedure
- **Revisiting Epipolar Geometry**
  - **Triangulation**
  - **Calibrated case: Essential matrix**
  - **Uncalibrated case: Fundamental matrix**
  - **Weak calibration**
  - **Epipolar Transfer**
- Active Stereo
  - Laser scanning
  - Kinect sensor

Two-View Geometry

- **Scene geometry (structure):**
  - Given corresponding points in two or more images, where is the pre-image of these points in 3D?
- **Correspondence (stereo matching):**
  - Given a point in just one image, how does it constrain the position of the corresponding point x' in another image?
- **Camera geometry (motion):**
  - Given a set of corresponding points in two images, what are the cameras for the two views?

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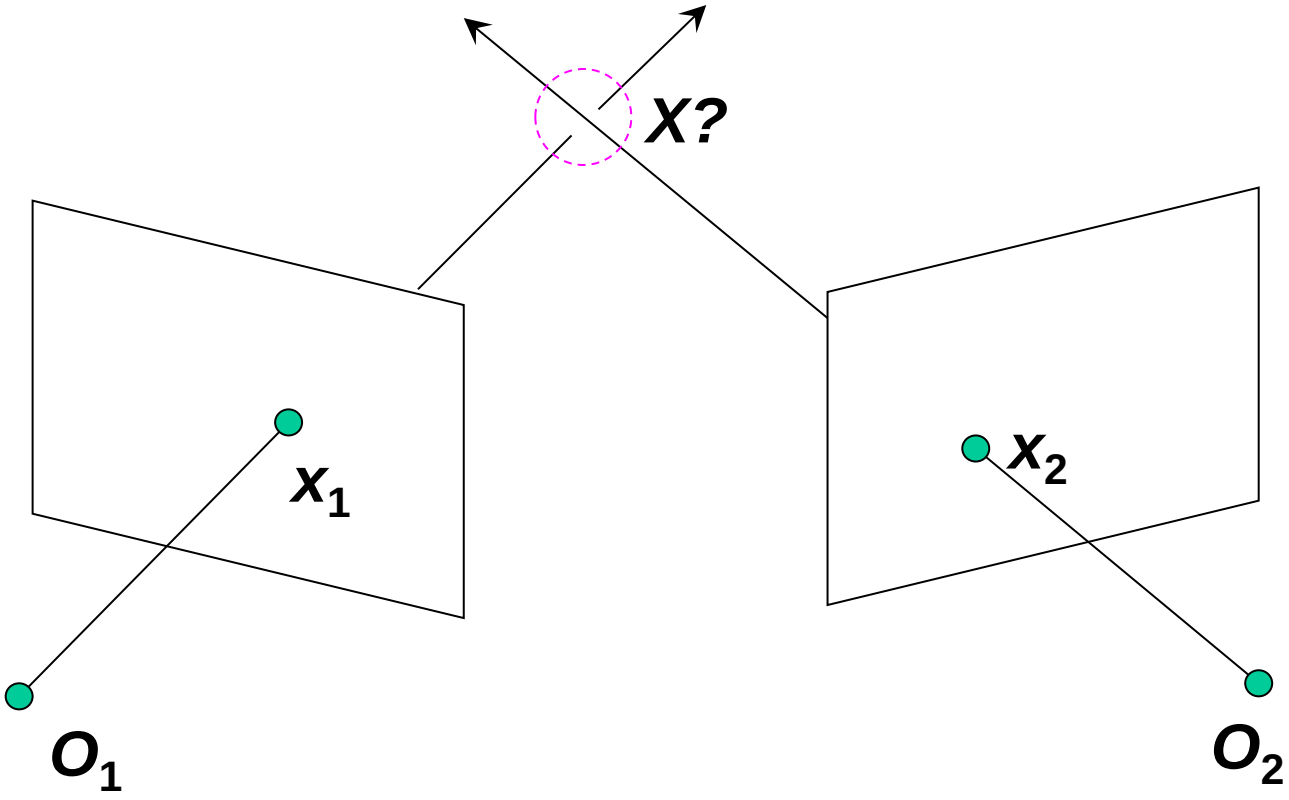
Revisiting Triangulation

- Given projections of a 3D point in two or more images (with known camera matrices), find the coordinates of the point



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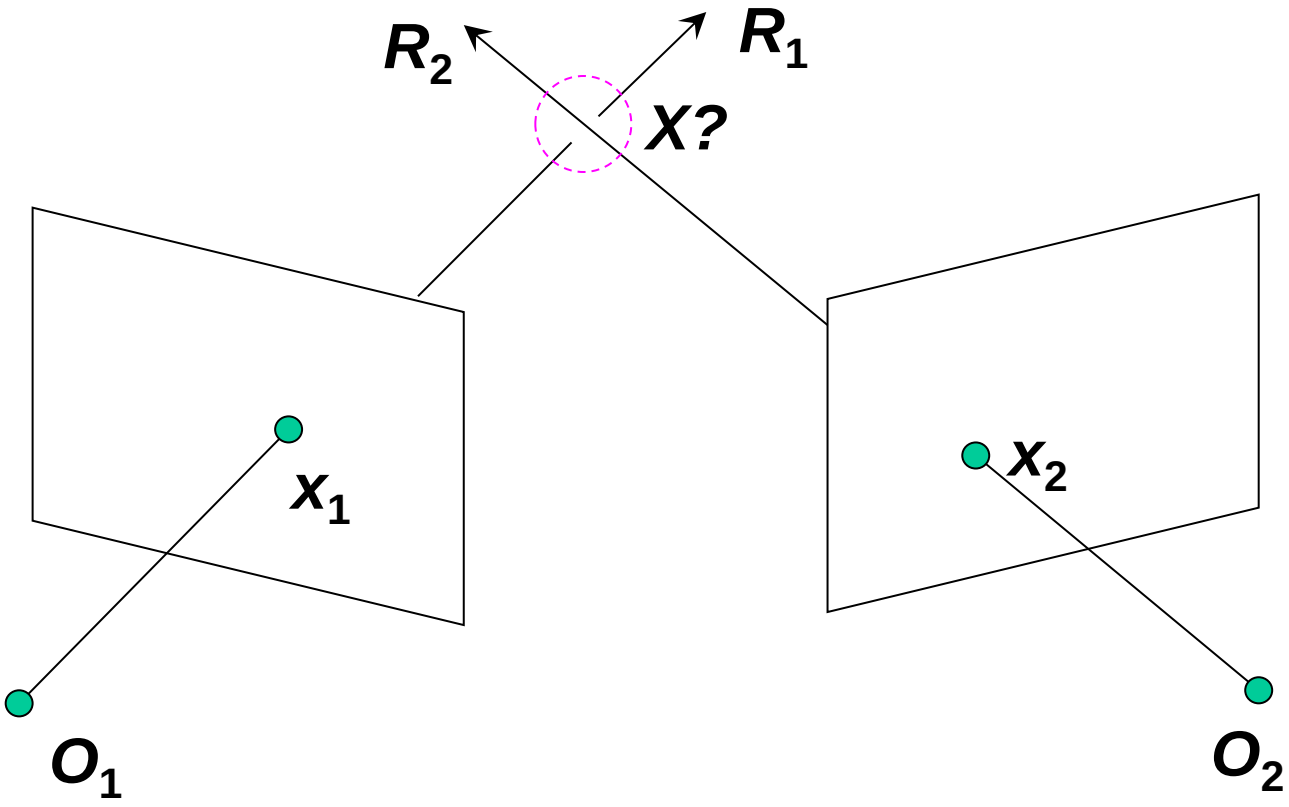
Revisiting Triangulation

- We want to intersect the two visual rays corresponding to x_1 and x_2, but because of noise and numerical errors, they will never meet exactly. How can this be done?



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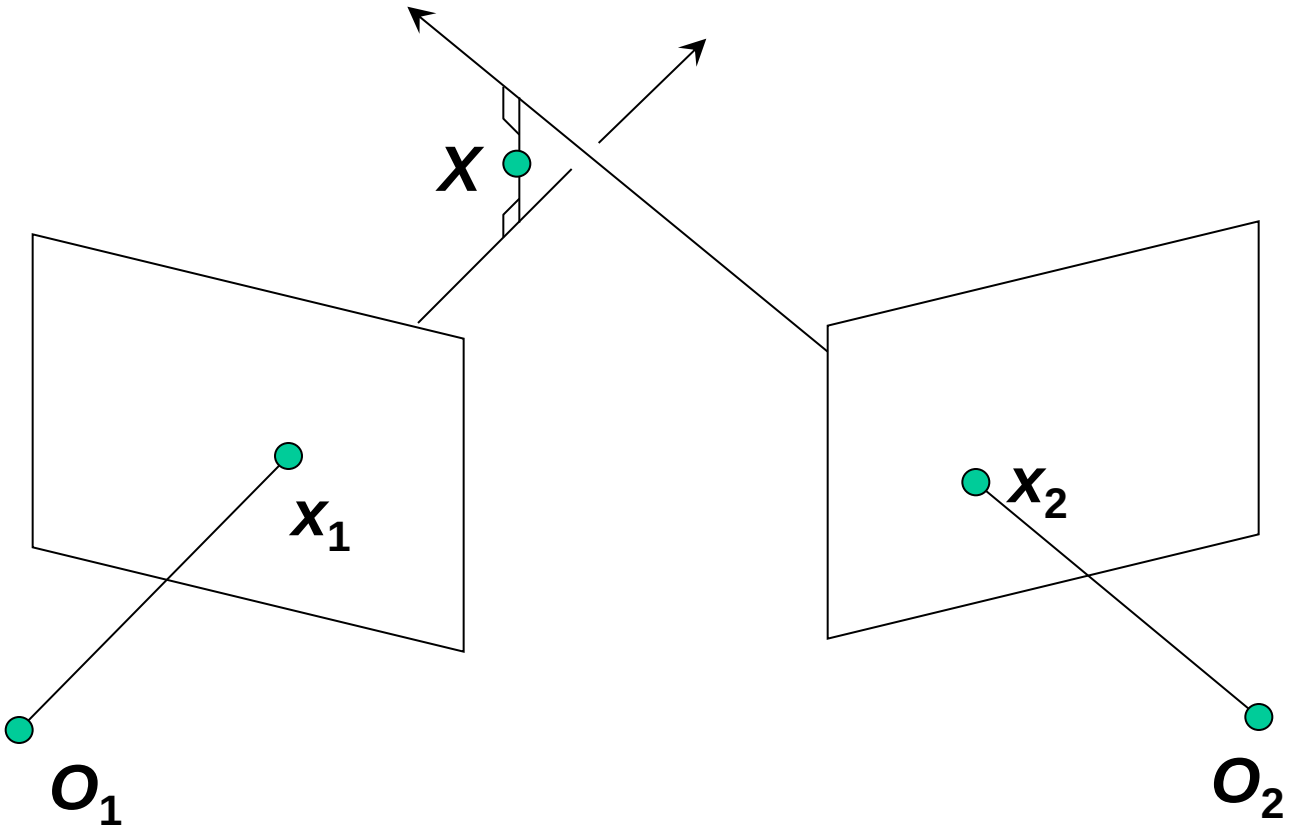
Triangulation: 1) Geometric Approach

- Find shortest segment connecting the two viewing rays and let X be the midpoint of that segment.



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Triangulation: 2)Linear Algebraic Approach

$$\lambda_1 \mathbf{x}_1 = \mathbf{P}_1 \mathbf{X} \qquad \mathbf{x}_1 \times \mathbf{P}_1 \mathbf{X} = 0 \qquad [\mathbf{x}_{1\times}]\mathbf{P}_1 \mathbf{X} = 0$$

$$\lambda_2 \mathbf{x}_2 = \mathbf{P}_2 \mathbf{X} \qquad \mathbf{x}_2 \times \mathbf{P}_2 \mathbf{X} = 0 \qquad [\mathbf{x}_{2\times}]\mathbf{P}_2 \mathbf{X} = 0$$

Cross product as matrix multiplication:

$$\mathbf{a} \times \mathbf{b} = \begin{bmatrix} 0 & -a_z & a_y \\ a_z & 0 & -a_x \\ -a_y & a_x & 0 \end{bmatrix} \begin{bmatrix} b_x \\ b_y \\ b_z \end{bmatrix} = [\mathbf{a}_\times]\mathbf{b}$$

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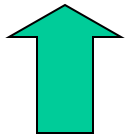
Triangulation: 2) Linear Algebraic Approach

$$\lambda_1 x_1 = P_1 X \qquad x_1 \times P_1 X = 0 \qquad [x_{1\times}] P_1 X = 0$$

$$\lambda_2 x_2 = P_2 X \qquad x_2 \times P_2 X = 0 \qquad [x_{2\times}] P_2 X = 0$$



Two independent equations each in terms of three unknown entries of X

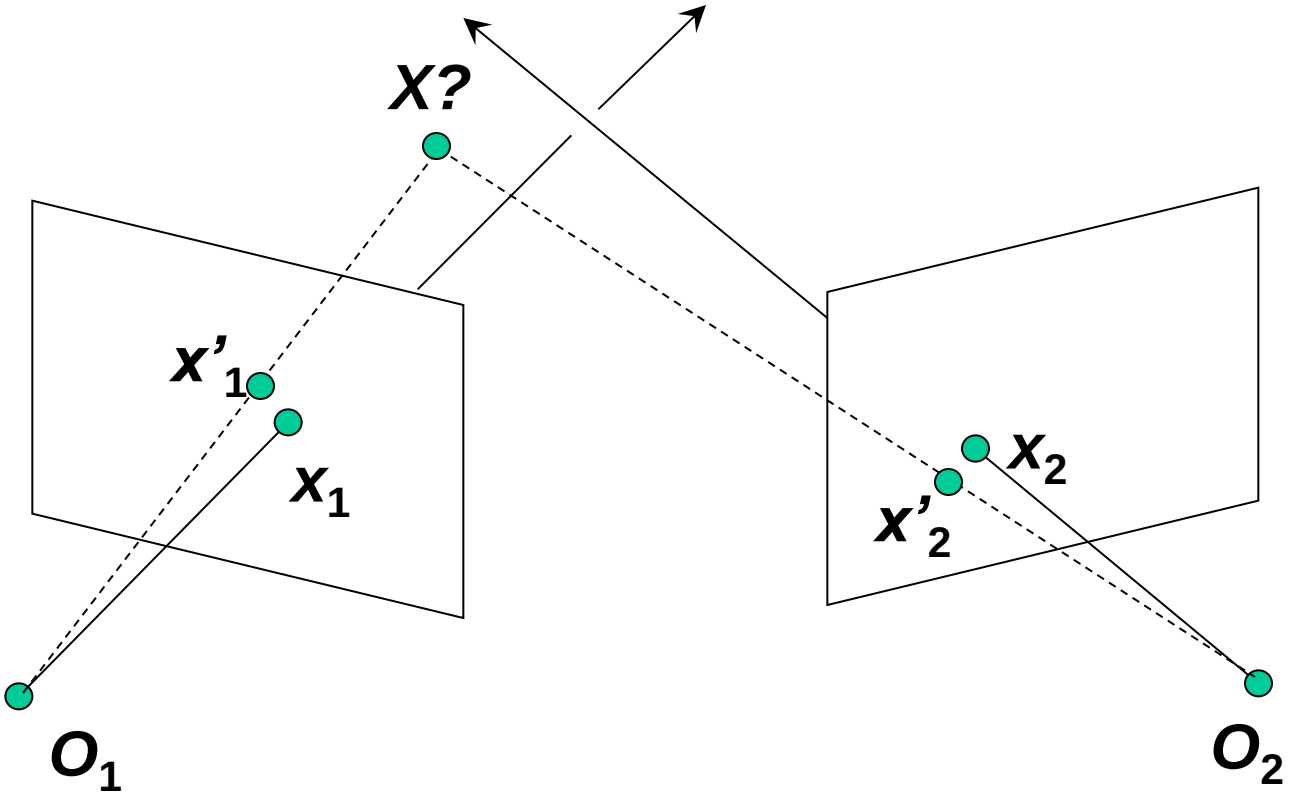
$\Rightarrow$ Stack them and solve using SVD!

- **This approach is often preferable to the geometric approach, since it nicely generalizes to multiple cameras.**

Triangulation: 3) Nonlinear Approach

- **Find X that minimizes**

$$d^2(x_1, P_1X) + d^2(x_2, P_2X)$$



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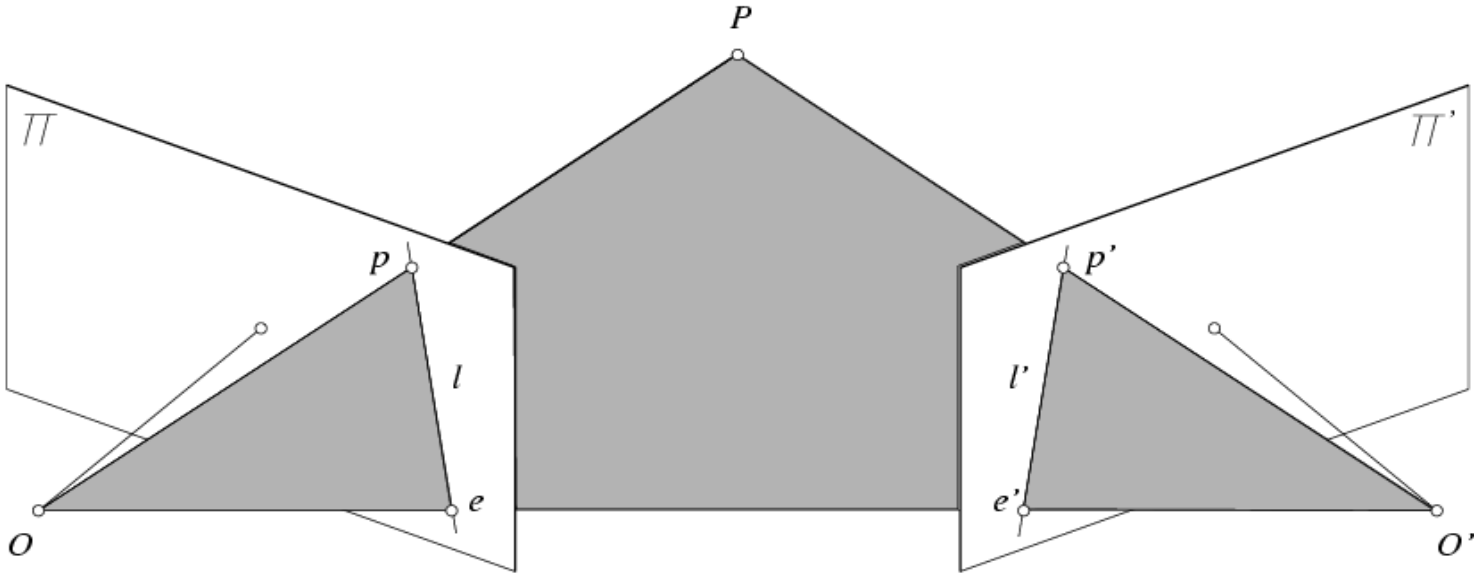
Triangulation: 3) Nonlinear Approach

- Find X that minimizes

$$d^2(x_1, P_1X) + d^2(x_2, P_2X)$$

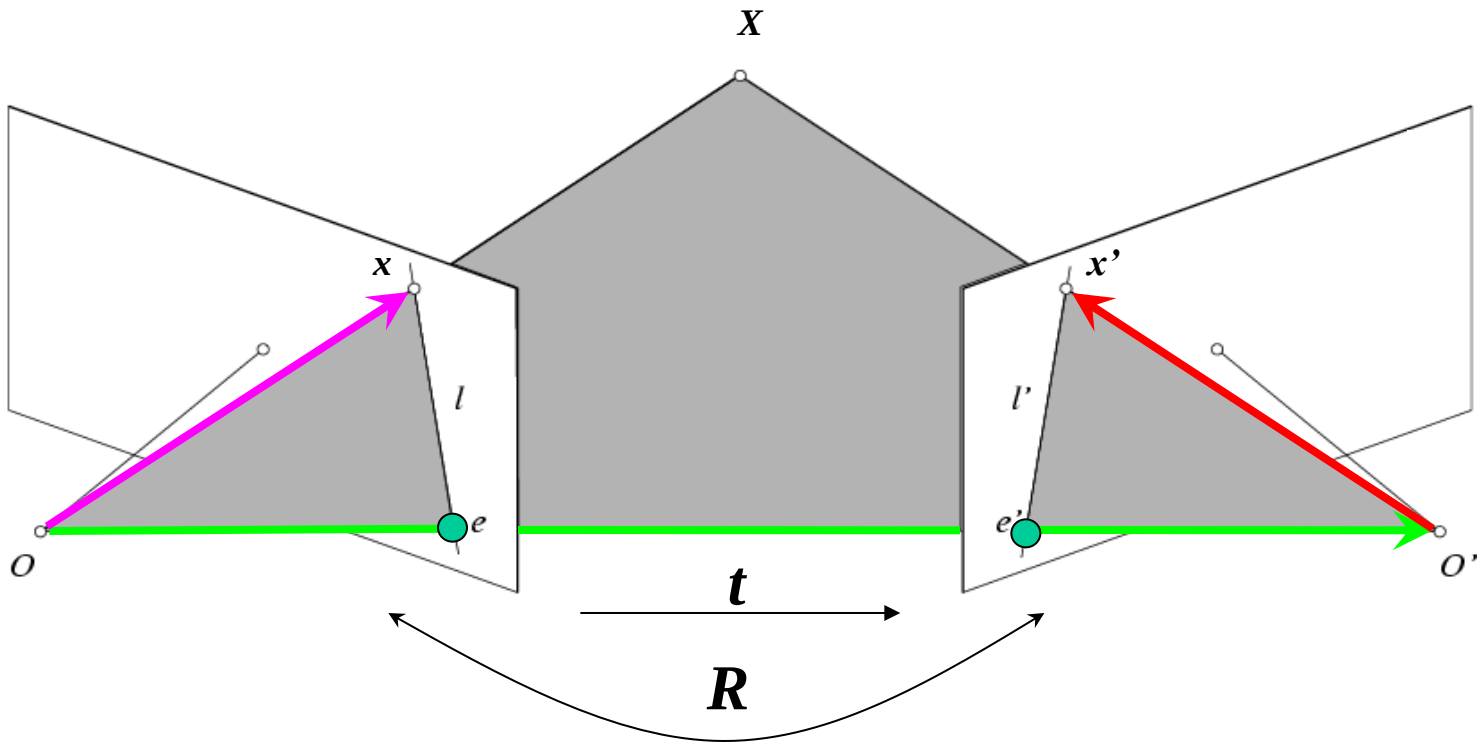
- This approach is the most accurate, but unlike the other two methods, it doesn't have a closed-form solution.
- Iterative algorithm
  - Initialize with linear estimate.
  - Optimize with Gauss-Newton or Levenberg-Marquardt (see F&P sec. 3.1.2 or H&Z Appendix 6).

Revisiting Epipolar Geometry



- Let's look again at the epipolar constraint
  - For the calibrated case (but in homogenous coordinates)
  - For the uncalibrated case

Epipolar Geometry: Calibrated Case



Camera matrix: [I|0]

$X = (u, v, w, 1)^T$

$x = (u, v, w)^T$

Camera matrix: $[R^T \mid -R^T t]$

Vector x' in second coord. system has coordinates Rx' in the first one.

The vectors x, t, and Rx' are coplanar

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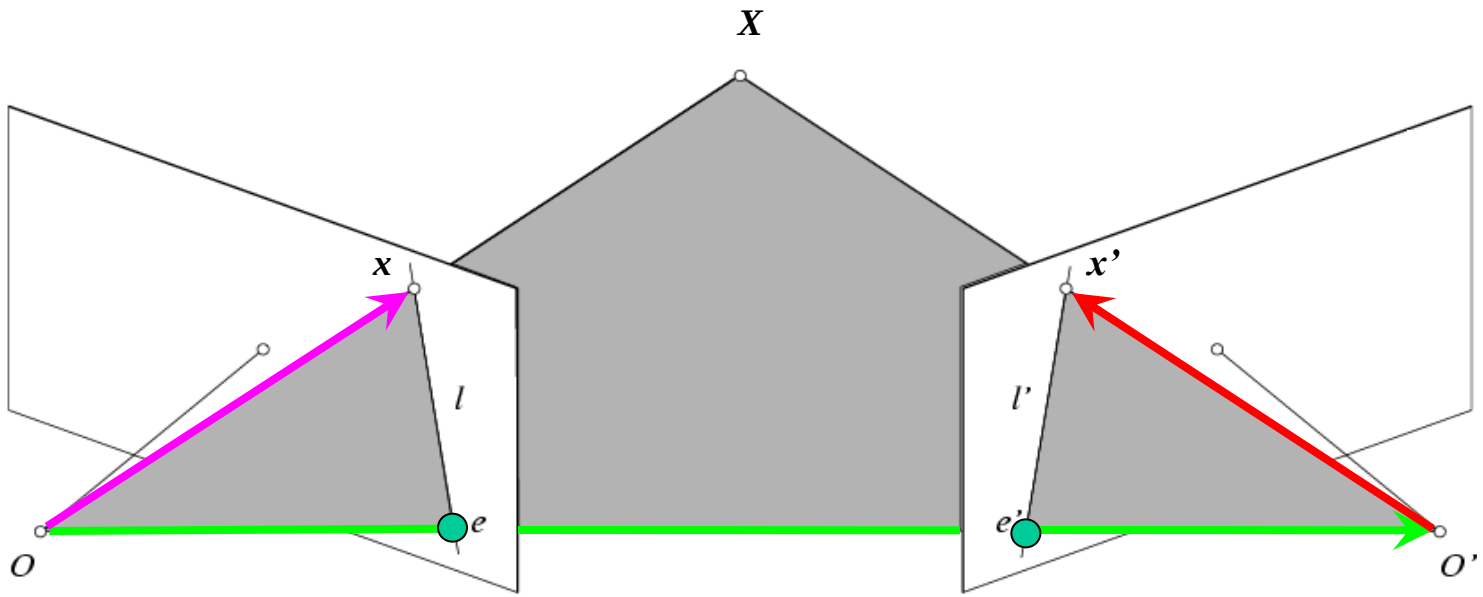
Epipolar Geometry: Calibrated Case



$$x \cdot [t \times (Rx')] = 0 \quad \Rightarrow \quad x^T E x' = 0 \quad \text{with} \quad E = [t_\times] R$$

Essential Matrix
(Longuet-Higgins, 1981)

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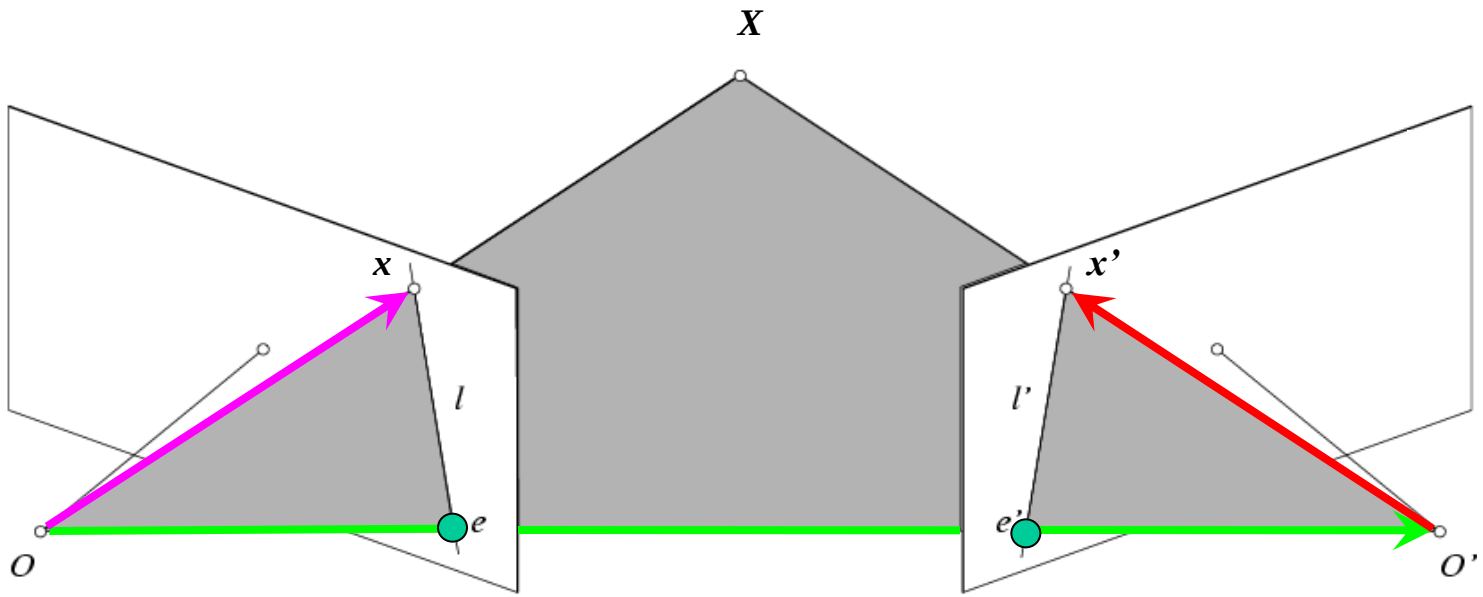
Epipolar Geometry: Calibrated Case



$$x \cdot [t \times (Rx')] = 0 \quad \Rightarrow \quad x^T E x' = 0 \quad \text{with} \quad E = [t_\times]R$$

- ***E x'*** **is the epipolar line associated with *x'* (*l* = *E x'*)**
- ***E*^T*x*** **is the epipolar line associated with *x* (*l'* = *E*^T*x*)**
- ***E e'*** **= 0 and *E*^T*e* = 0**
- ***E*** **is singular (rank two)**
- ***E*** **has five degrees of freedom (up to scale)**

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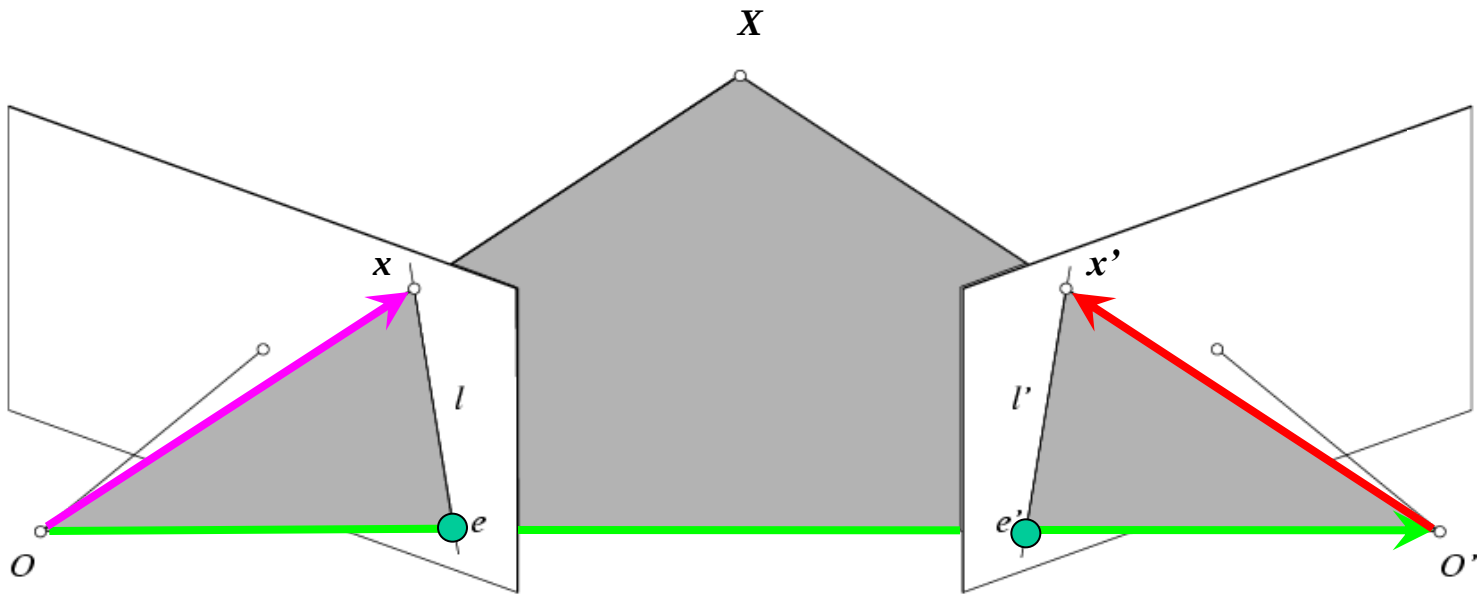
Epipolar Geometry: Uncalibrated Case



- The calibration matrices *K* and *K'* of the two cameras are unknown
- We can write the epipolar constraint in terms of *unknown* normalized coordinates:

$$\hat{x}^T E \hat{x}' = 0 \qquad x = K\hat{x}, \quad x' = K'\hat{x}'$$

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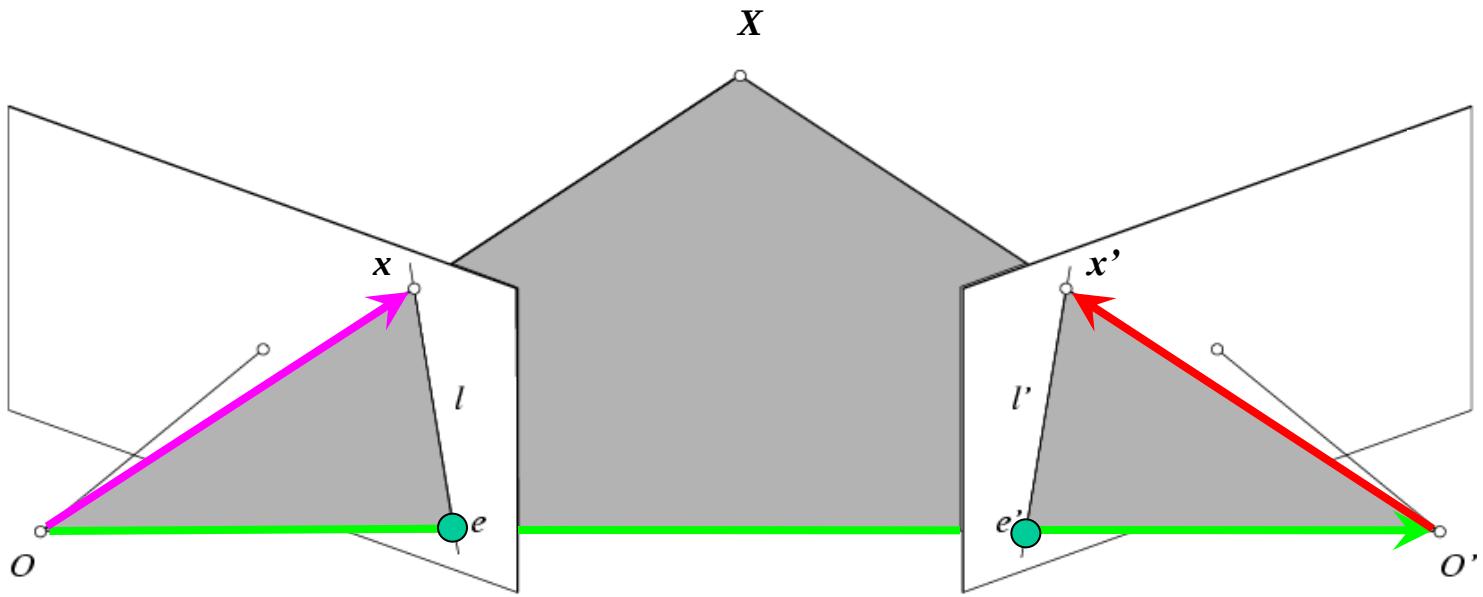
Epipolar Geometry: Uncalibrated Case



$$\hat{x}^T E \hat{x}' = 0 \quad \Rightarrow \quad x^T F x' = 0 \quad \text{with} \quad F = K^{-T} E K'^{-1}$$

$$x = K\hat{x}$$

$$x' = K'\hat{x}'$$

Fundamental Matrix
(Faugeras and Luong, 1992)

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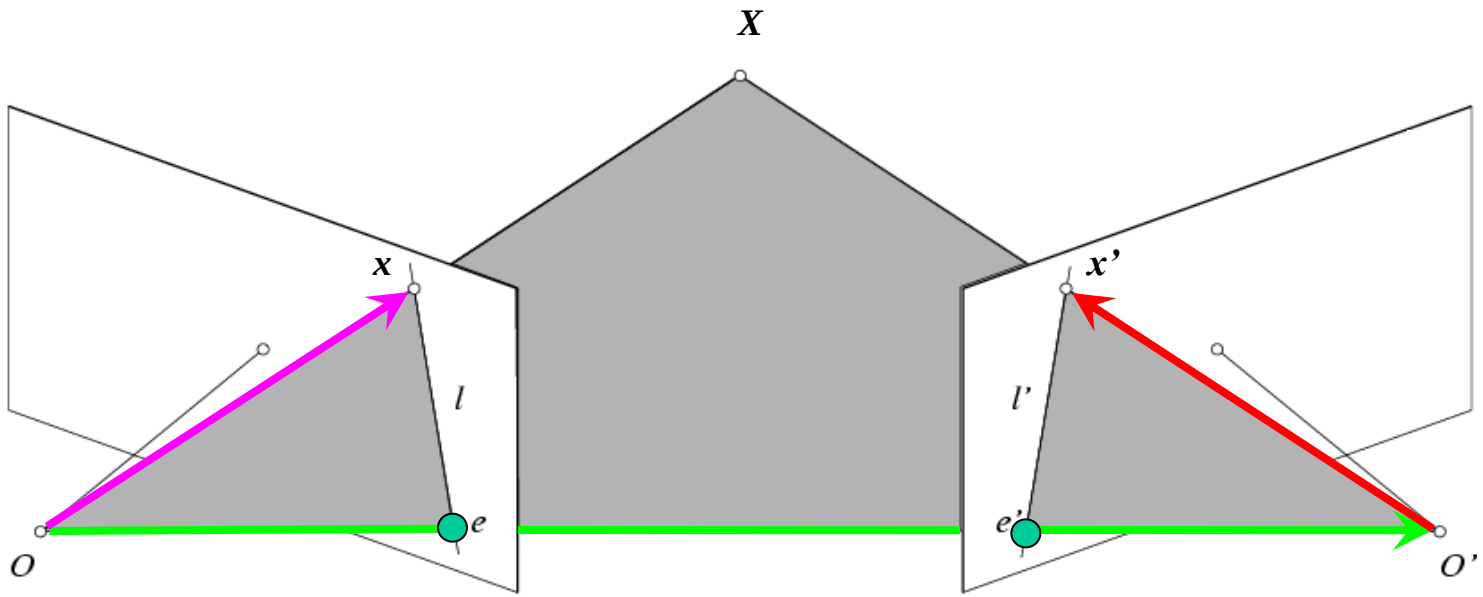
Epipolar Geometry: Uncalibrated Case



$$\hat{x}^T E \hat{x}' = 0 \quad \Rightarrow \quad x^T F x' = 0 \quad \text{with} \quad F = K^{-T} E K'^{-1}$$

- ***F x'*** **is the epipolar line associated with *x'* (*l* = *F x'*)**
- ***F*^T*x*** **is the epipolar line associated with *x*** **(*l'* = *F*^T*x*)**
- ***F e'*** **= 0 and** ***F*^T*e*** **= *0***
- ***F*** **is singular (rank two)**
- ***F*** **has seven degrees of freedom**

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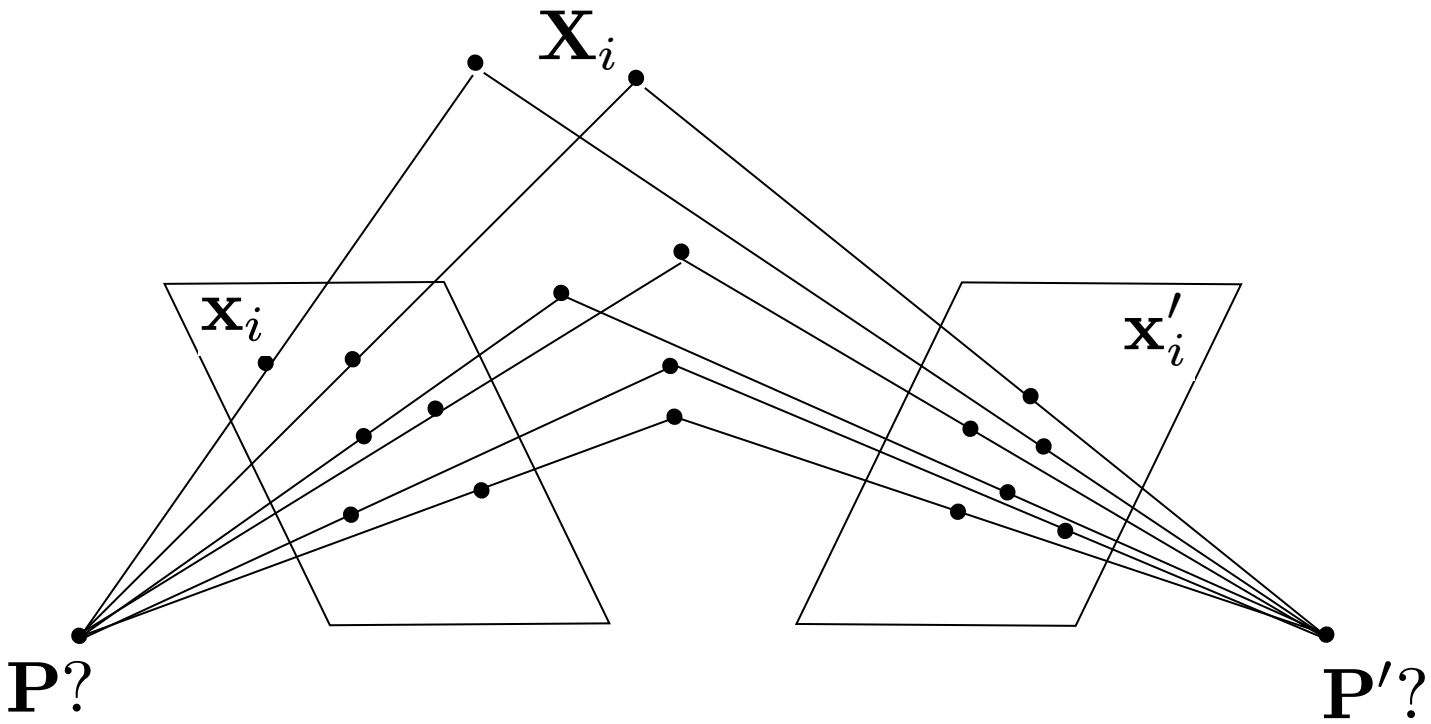
Estimating the Fundamental Matrix

- The Fundamental matrix defines the epipolar geometry between two uncalibrated cameras.
- How can we estimate F from an image pair?
  - We need correspondences…



The Eight-Point Algorithm

$x = (u, v, 1)^T, \quad x' = (u', v', 1)^T$

$$(u, v, 1)\begin{pmatrix} F_{11} & F_{12} & F_{13} \\ F_{21} & F_{22} & F_{23} \\ F_{31} & F_{32} & F_{33} \end{pmatrix}\begin{pmatrix} u' \\ v' \\ 1 \end{pmatrix} = 0$$

$$[u'u, u'v, u', uv', vv', v', u, v, 1]\begin{bmatrix} F_{11} \\ F_{12} \\ F_{13} \\ F_{21} \\ F_{22} \\ F_{23} \\ F_{31} \\ F_{32} \\ F_{33} \end{bmatrix} = 0$$

- **Taking 8 correspondences:**

$$\begin{bmatrix} u'_1u_1 & u'_1v_1 & u'_1 & u_1v'_1 & v_1v'_1 & v'_1 & u_1 & v_1 & 1 \\ u'_2u_2 & u'_2v_2 & u'_2 & u_2v'_2 & v_2v'_2 & v'_2 & u_2 & v_2 & 1 \\ u'_3u_3 & u'_3v_3 & u'_3 & u_3v'_3 & v_3v'_3 & v'_3 & u_3 & v_3 & 1 \\ u'_4u_4 & u'_4v_4 & u'_4 & u_4v'_4 & v_4v'_4 & v'_4 & u_4 & v_4 & 1 \\ u'_5u_5 & u'_5v_5 & u'_5 & u_5v'_5 & v_5v'_5 & v'_5 & u_5 & v_5 & 1 \\ u'_6u_6 & u'_6v_6 & u'_6 & u_6v'_6 & v_6v'_6 & v'_6 & u_6 & v_6 & 1 \\ u'_7u_7 & u'_7v_7 & u'_7 & u_7v'_7 & v_7v'_7 & v'_7 & u_7 & v_7 & 1 \\ u'_8u_8 & u'_8v_8 & u'_8 & u_8v'_8 & v_8v'_8 & v'_8 & u_8 & v_8 & 1 \end{bmatrix}\begin{bmatrix} F_{11} \\ F_{12} \\ F_{13} \\ F_{21} \\ F_{22} \\ F_{23} \\ F_{31} \\ F_{32} \\ F_{33} \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \end{bmatrix}$$

$$\mathbf{Af} = \mathbf{0}$$

Solve using… SVD!

This minimizes:

$$\sum_{i=1}^{N}(x_i^T F x_i')^2$$

Excursion: Properties of SVD

- **Frobenius norm**
  - Generalization of the Euclidean norm to matrices

$$\|A\|_F = \sqrt{\sum_{i=1}^{m}\sum_{j=1}^{n}|a_{ij}|^2} = \sqrt{\sum_{i=1}^{\min(m,n)}\sigma_i^2}$$

- **Partial reconstruction property of SVD**
  - Let σ_i $i=1,\ldots,N$ be the singular values of A.
  - Let $A_p = U_p D_p V_p^T$ be the reconstruction of A when we set $\sigma_{p+1}, \ldots, \sigma_N$ to zero.
  - Then $A_p = U_p D_p V_p^T$ is the best rank-p approximation of A in the sense of the Frobenius norm (i.e. the best least-squares approximation).

The Eight-Point Algorithm

- **Problem with noisy data**
  - The solution will usually not fulfill the constraint that F only has rank 2.

  ⇒ *There will be no epipoles through which all epipolar lines pass!*

- **Enforce the rank-2 constraint using SVD**

$$F \overset{\text{SVD}}{=} \boldsymbol{U}\boldsymbol{D}\boldsymbol{V}^T = \boldsymbol{U}\begin{bmatrix} d_{11} & & \\ & d_{22} & \\ & & d_{33} \end{bmatrix}\begin{bmatrix} v_{11} & \cdots & v_{13} \\ \vdots & \ddots & \vdots \\ v_{31} & \cdots & v_{33} \end{bmatrix}^T$$

Set d_{33} to zero and reconstruct F

- **As we have just seen, this provides the best least-squares approximation to the rank-2 solution.**

Problem with the Eight-Point Algorithm

- **In practice, this often looks as follows:**

$$\begin{bmatrix} u'_1u_1 & u'_1v_1 & u'_1 & u_1v'_1 & v_1v'_1 & v'_1 & u_1 & v_1 & 1 \\ u'_2u_2 & u'_2v_2 & u'_2 & u_2v'_2 & v_2v'_2 & v'_2 & u_2 & v_2 & 1 \\ u'_3u_3 & u'_3v_3 & u'_3 & u_3v'_3 & v_3v'_3 & v'_3 & u_3 & v_3 & 1 \\ u'_4u_4 & u'_4v_4 & u'_4 & u_4v'_4 & v_4v'_4 & v'_4 & u_4 & v_4 & 1 \\ u'_5u_5 & u'_5v_5 & u'_5 & u_5v'_5 & v_5v'_5 & v'_5 & u_5 & v_5 & 1 \\ u'_6u_6 & u'_6v_6 & u'_6 & u_6v'_6 & v_6v'_6 & v'_6 & u_6 & v_6 & 1 \\ u'_7u_7 & u'_7v_7 & u'_7 & u_7v'_7 & v_7v'_7 & v'_7 & u_7 & v_7 & 1 \\ u'_8u_8 & u'_8v_8 & u'_8 & u_8v'_8 & v_8v'_8 & v'_8 & u_8 & v_8 & 1 \end{bmatrix} \begin{bmatrix} F_{11} \\ F_{12} \\ F_{13} \\ F_{21} \\ F_{22} \\ F_{23} \\ F_{31} \\ F_{32} \\ F_{33} \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \end{bmatrix}$$

Problem with the Eight-Point Algorithm

- In practice, this often looks as follows:

$$\begin{bmatrix} 250906.36 & 183269.57 & 921.81 & 200931.10 & 146766.13 & 738.21 & 272.19 & 198.81 & 1 \\ 2692.28 & 131633.03 & 176.27 & 6196.73 & 302975.59 & 405.71 & 15.27 & 746.79 & 1 \\ 416374.23 & 871684.30 & 935.47 & 408110.89 & 854384.92 & 916.90 & 445.10 & 931.81 & 1 \\ 191183.60 & 171759.40 & 410.27 & 416435.62 & 374125.90 & 893.65 & 465.99 & 418.65 & 1 \\ 48988.86 & 30401.76 & 57.89 & 298604.57 & 185309.58 & 352.87 & 846.22 & 525.15 & 1 \\ 164786.04 & 546559.67 & 813.17 & 1998.37 & 6628.15 & 9.86 & 202.65 & 672.14 & 1 \\ 116407.01 & 2727.75 & 138.89 & 169941.27 & 3982.21 & 202.77 & 838.12 & 19.64 & 1 \\ 135384.58 & 75411.13 & 198.72 & 411350.03 & 229127.78 & 603.79 & 681.28 & 379.48 & 1 \end{bmatrix} \begin{bmatrix} F_{11} \\ F_{12} \\ F_{13} \\ F_{21} \\ F_{22} \\ F_{23} \\ F_{31} \\ F_{32} \\ F_{33} \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \end{bmatrix}$$

⇒ Poor numerical conditioning

⇒ Can be fixed by rescaling the data

The Normalized Eight-Point Algorithm

1. Center the image data at the origin, and scale it so the mean squared distance between the origin and the data points is 2 pixels.
2. Use the eight-point algorithm to compute *F* from the normalized points.
3. Enforce the rank-2 constraint using SVD.

$$F \overset{\text{SVD}}{=} \boldsymbol{U}\boldsymbol{D}\boldsymbol{V}^T = \boldsymbol{U}\begin{bmatrix} d_{11} & & \\ & d_{22} & \\ & & d_{33} \end{bmatrix}\begin{bmatrix} v_{11} & \cdots & v_{13} \\ \vdots & \ddots & \vdots \\ v_{31} & \cdots & v_{33} \end{bmatrix}^T$$

Set d_{33} to zero and reconstruct F

4. Transform fundamental matrix back to original units: if T and T' are the normalizing transformations in the two images, than the fundamental matrix in original coordinates is $T^T F\, T'$.

Slide credit: Svetlana Lazebnik

[Hartley, 1995]

The Eight-Point Algorithm

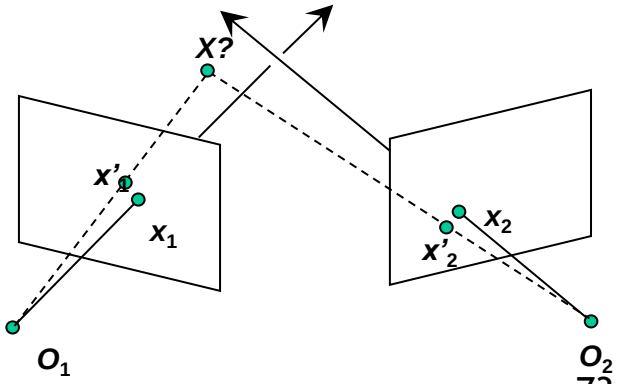
- **Meaning of error** $\sum_{i=1}^{N} (x_i^T F x_i')^2$ **:**

  Sum of Euclidean distances between points x_i and epipolar lines Fx'_i (or points x'_i and epipolar lines $F^T x_i$), multiplied by a scale factor

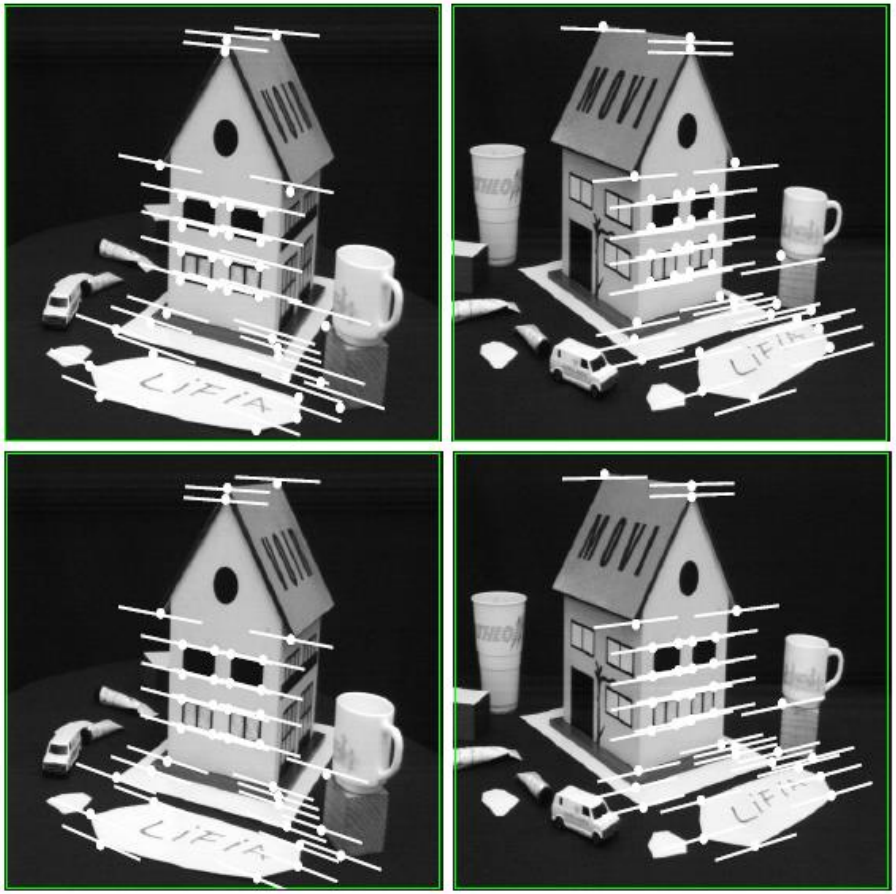
- **Nonlinear approach: minimize**

$$\sum_{i=1}^{N} \left[d^2(x_i, F x_i') + d^2(x_i', F^T x_i) \right]$$

  - **Similar to nonlinear minimization approach for triangulation.**
  - **Iterative approach (Gauss-Newton, Levenberg-Marquardt,…)**



Comparison of Estimation Algorithms



	8-point	Normalized 8-point	Nonlinear least squares
Av. Dist. 1	2.33 pixels	0.92 pixel	0.86 pixel
Av. Dist. 2	2.18 pixels	0.85 pixel	0.80 pixel

Slide credit: Svetlana Lazebnik

3D Reconstruction with Weak Calibration

- Want to estimate world geometry without requiring calibrated cameras.
- Many applications:
  - Archival videos
  - Photos from multiple unrelated users
  - Dynamic camera system
- Main idea:
  - Estimate epipolar geometry from a (redundant) set of point correspondences between two uncalibrated cameras.

Slide credit: Kristen Grauman

Stereo Pipeline with Weak Calibration

- So, where to start with uncalibrated cameras?
  - Need to find fundamental matrix F *and* the correspondences (pairs of points (u',v') ↔ (u,v)).

- Procedure
  1. Find interest points in both images
  2. Compute correspondences
  3. Compute epipolar geometry
  4. Refine

Stereo Pipeline with Weak Calibration

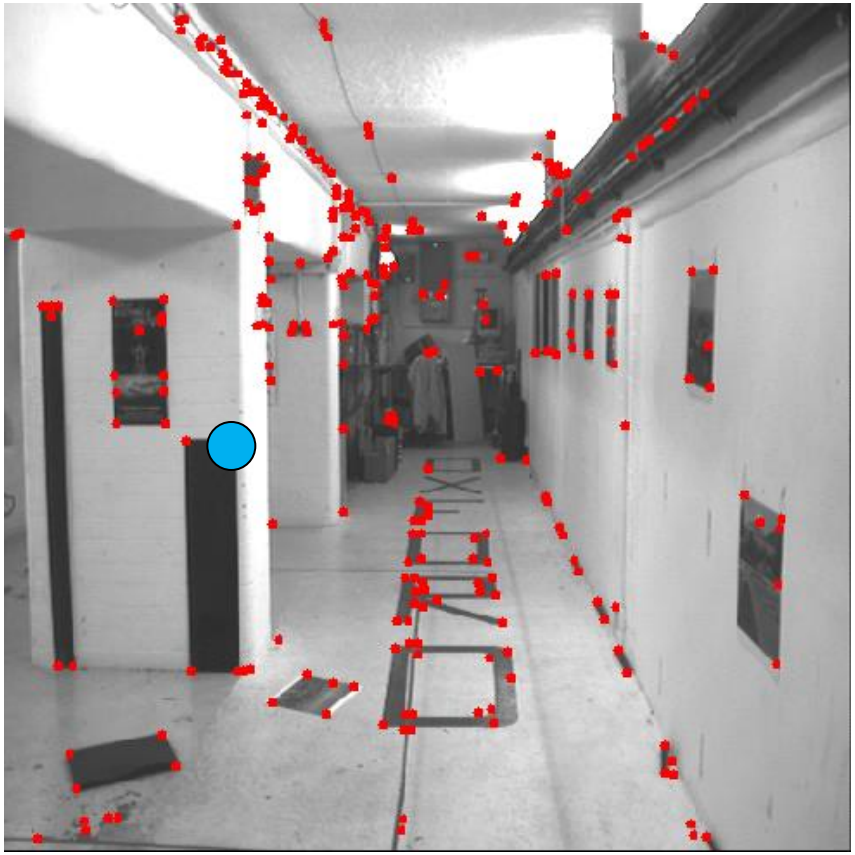
1. Find interest points (e.g. Harris corners)

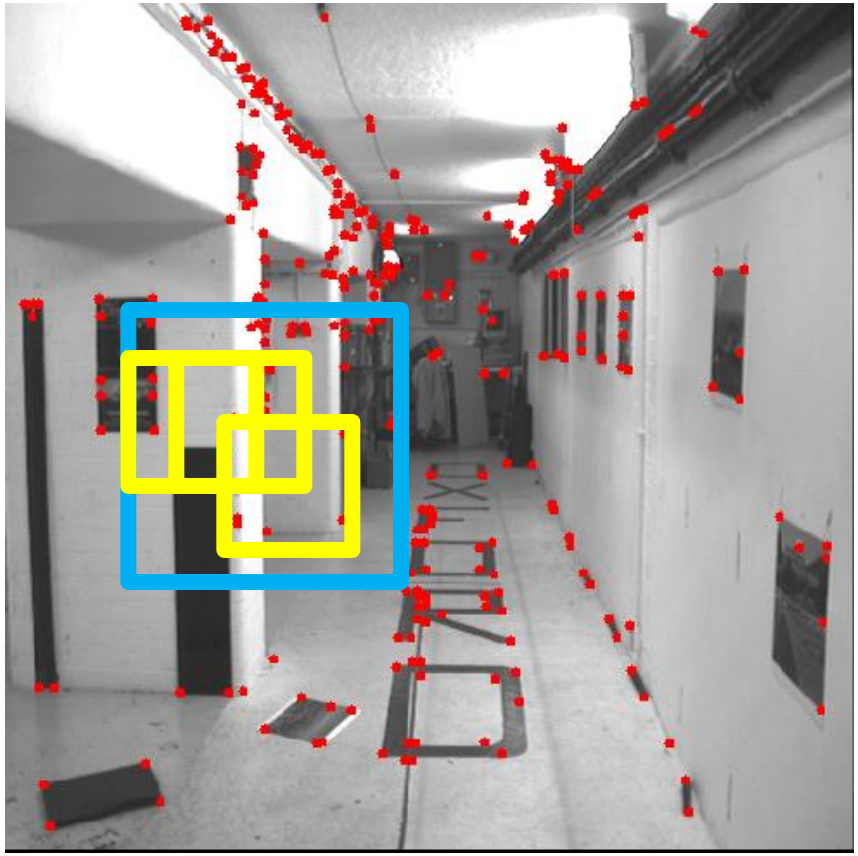
Slide credit: Kristen Grauman

Example from Andrew Zisserman

Stereo Pipeline with Weak Calibration

2. Match points using only proximity

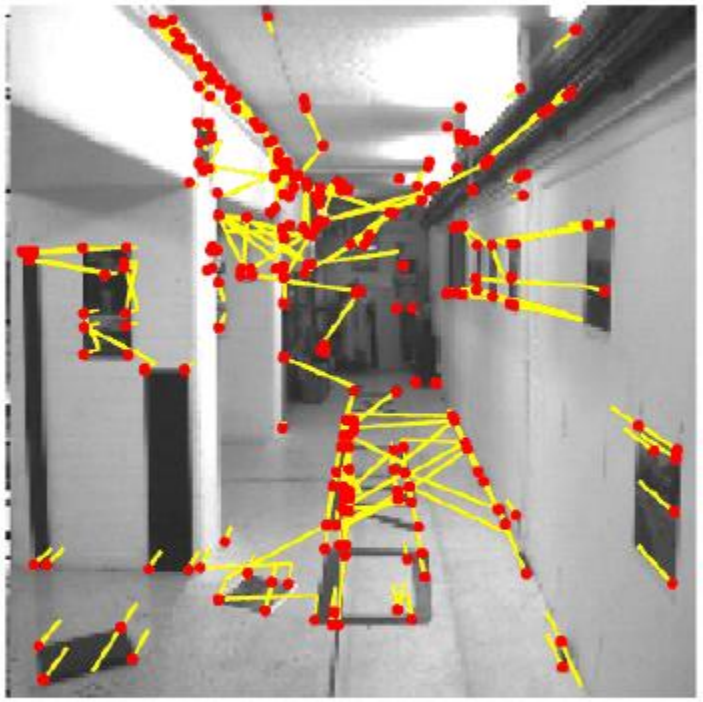




Slide credit: Kristen Grauman

Example from Andrew Zisserman

Putative Matches based on Correlation Search



- Many wrong matches (10-50%), but enough to compute F

Example from Andrew Zisserman

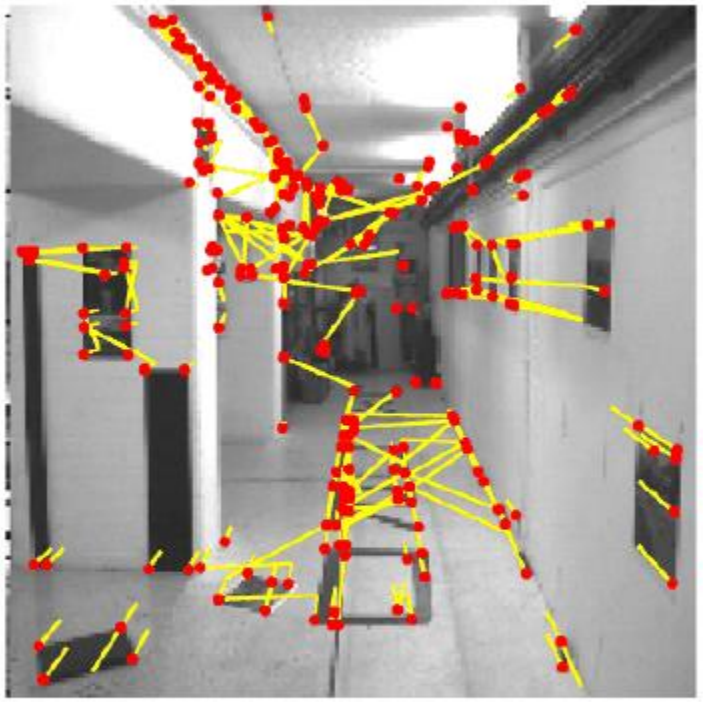
RANSAC for Robust Estimation of F

- Select random sample of correspondences
- Compute F using them
  - This determines epipolar constraint
- Evaluate amount of support – number of inliers within threshold distance of epipolar line

- Choose F with most support (#inliers)

Slide credit: Kristen Grauman

Putative Matches based on Correlation Search

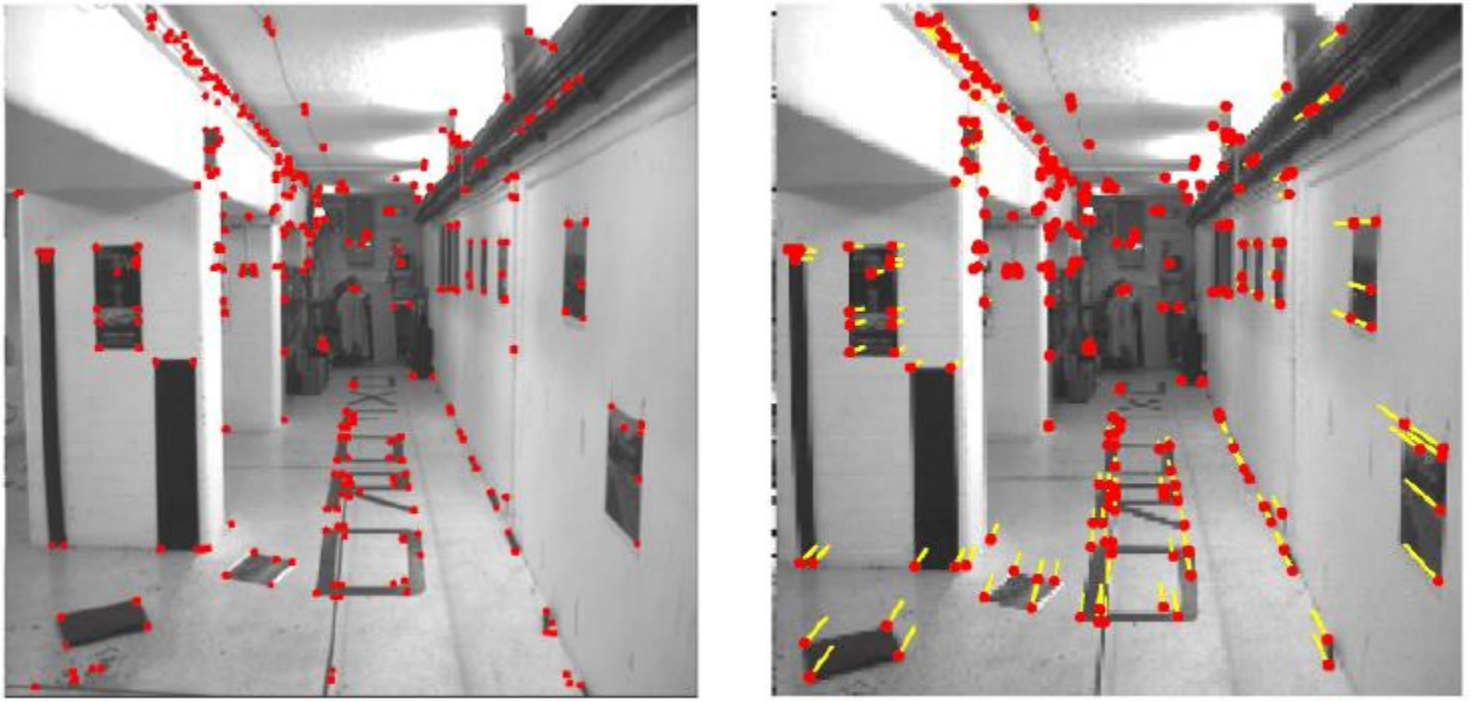


- Many wrong matches (10-50%), but enough to compute F

Example from Andrew Zisserman

Pruned Matches

- Correspondences consistent with epipolar geometry



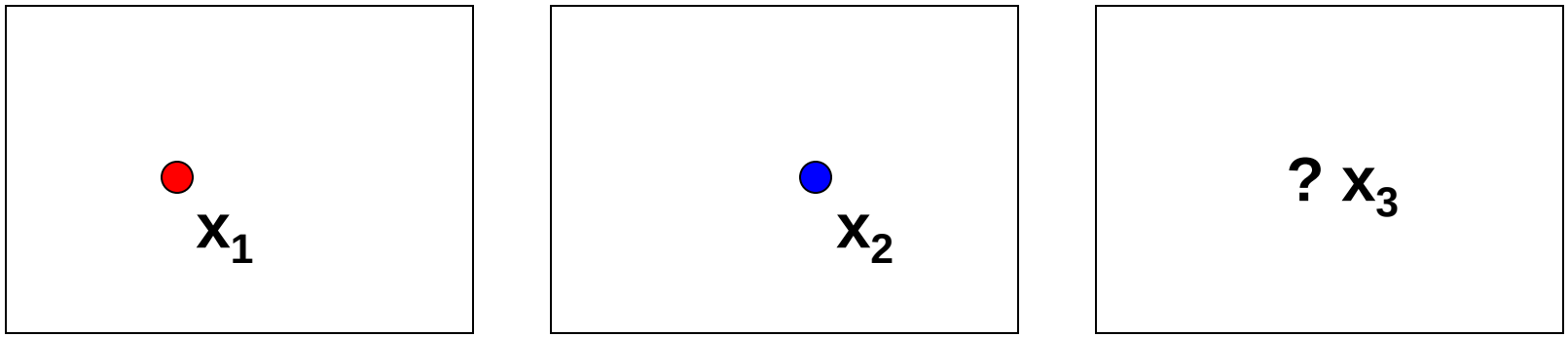
Example from Andrew Zisserman

Resulting Epipolar Geometry

Example from Andrew Zisserman

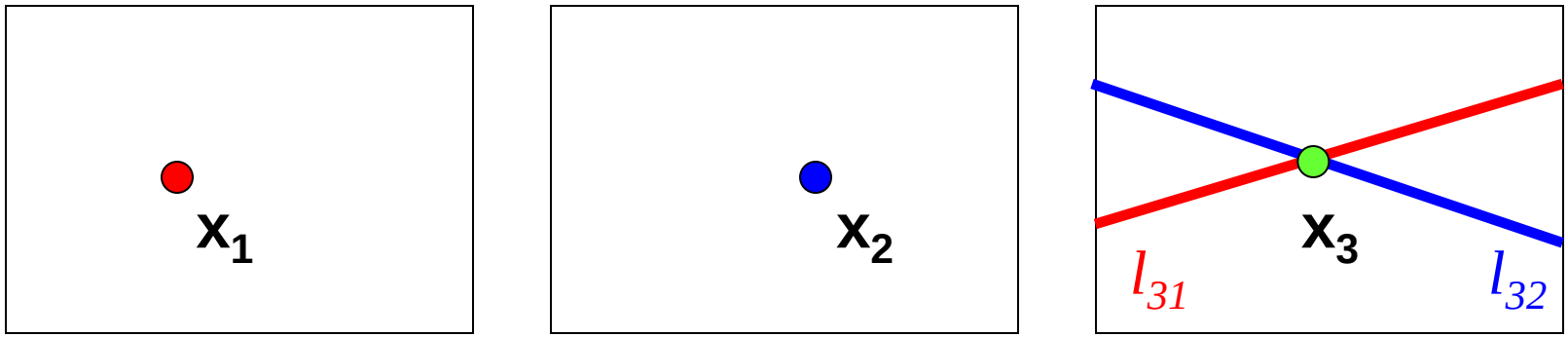
Epipolar Transfer

- Assume the epipolar geometry is known
- Given projections of the same point in two images, how can we compute the projection of that point in a third image?



Extension: Epipolar Transfer

- Assume the epipolar geometry is known
- Given projections of the same point in two images, how can we compute the projection of that point in a third image?



$$l_{31} = F^T{}_{13}\, x_1$$

$$l_{32} = F^T{}_{23}\, x_2$$

When does epipolar transfer fail?

Topics of This Lecture

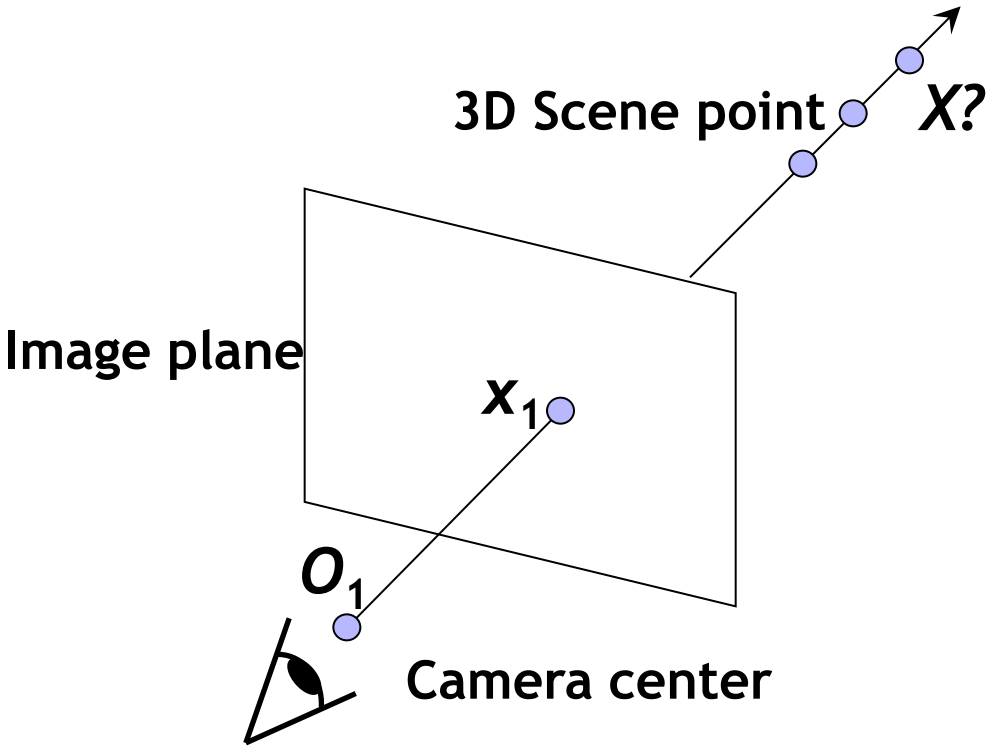
- Camera Calibration
  - Camera parameters
  - Calibration procedure
- Revisiting Epipolar Geometry
  - Triangulation
  - Calibrated case: Essential matrix
  - Uncalibrated case: Fundamental matrix
  - Weak calibration
  - Epipolar Transfer
- **Active Stereo**
  - **Laser scanning**
  - **Kinect sensor**

Microsoft Kinect – How Does It Work?

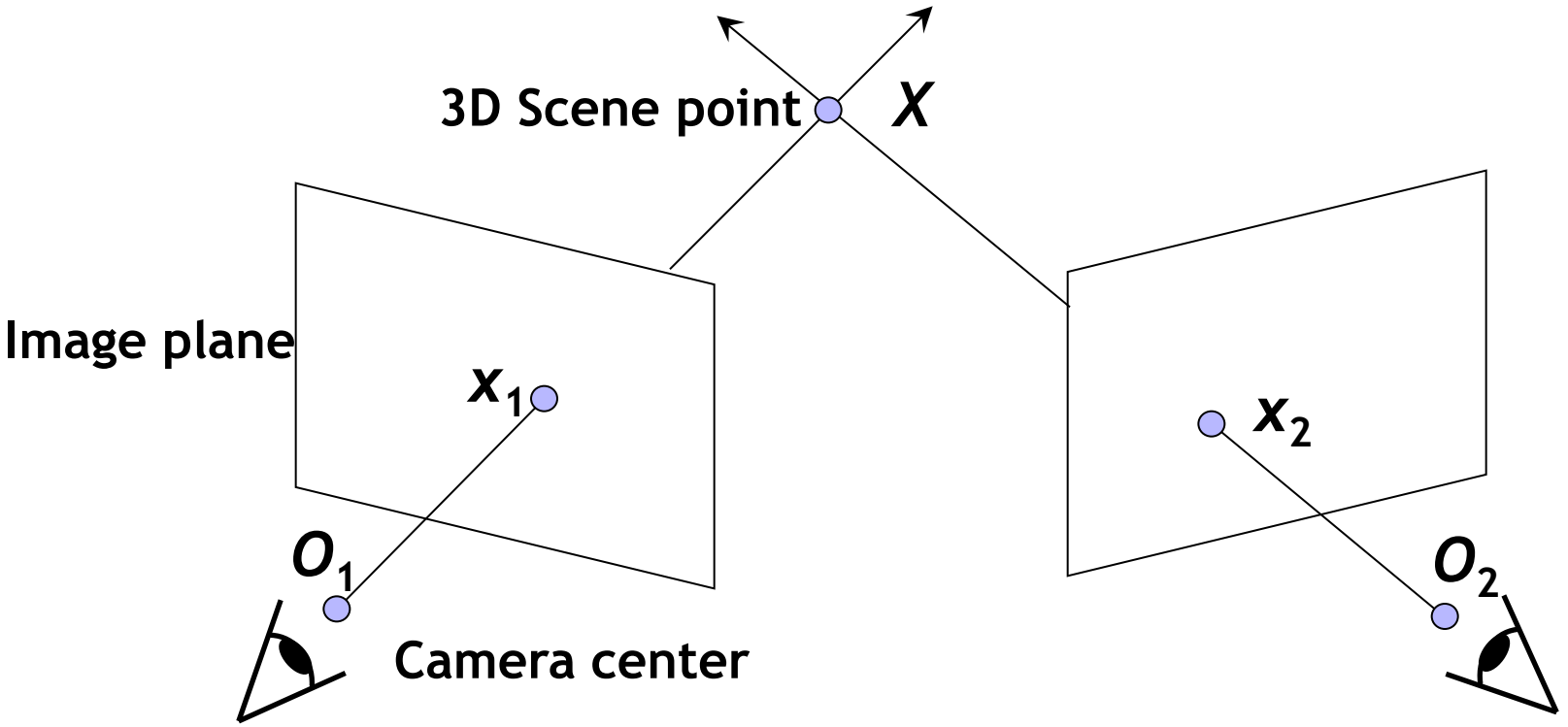
KINECT for XBOX 360.

- Built-in IR projector
- IR camera for depth
- Regular camera for color

Recall: Optical Triangulation

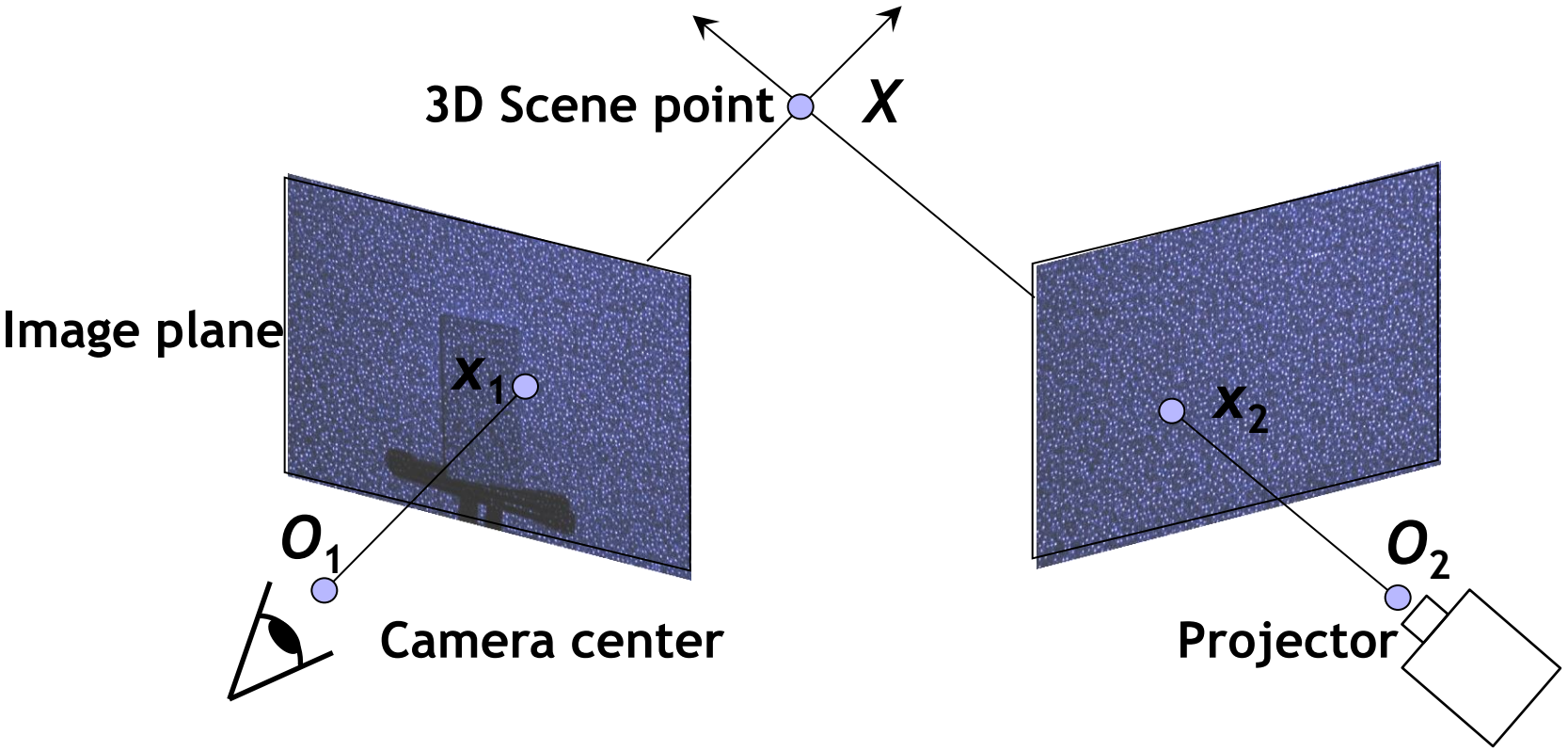


Recall: Optical Triangulation



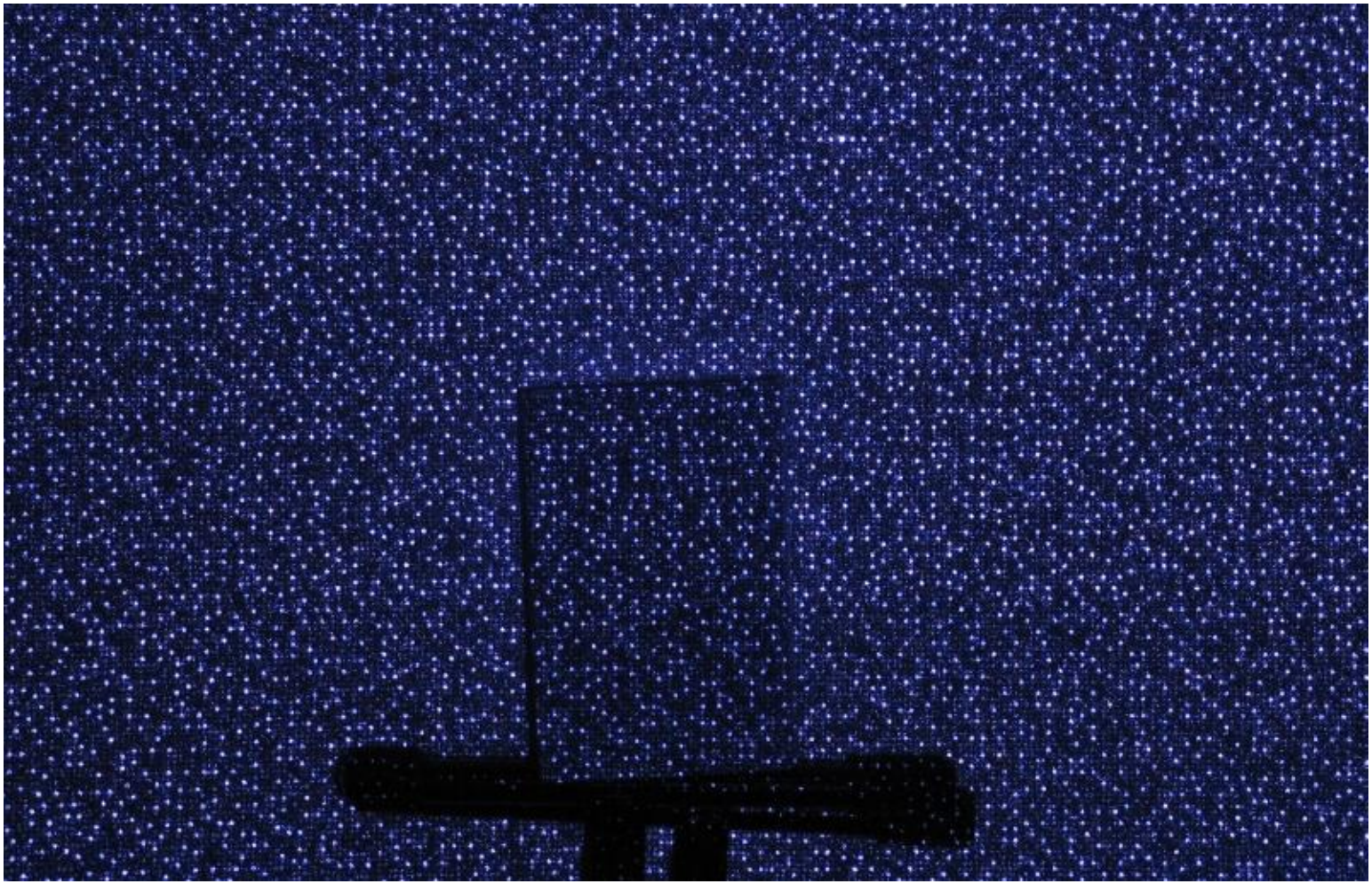
- Principle: 3D point given by intersection of two rays.
  - Crucial information: point correspondence
  - Most expensive and error-prone step in the pipeline…

Active Stereo with Structured Light



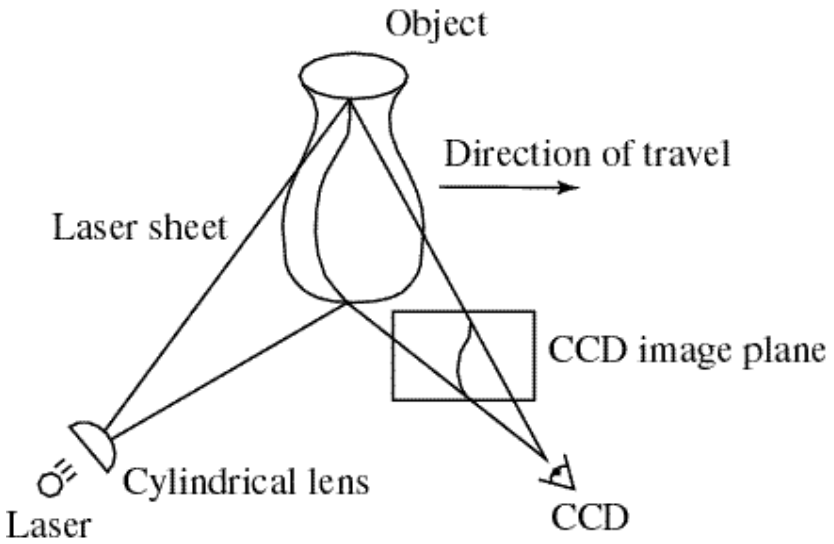
- Idea: Replace one camera by a projector.
  - Project "structured" light patterns onto the object
  - Simplifies the correspondence problem

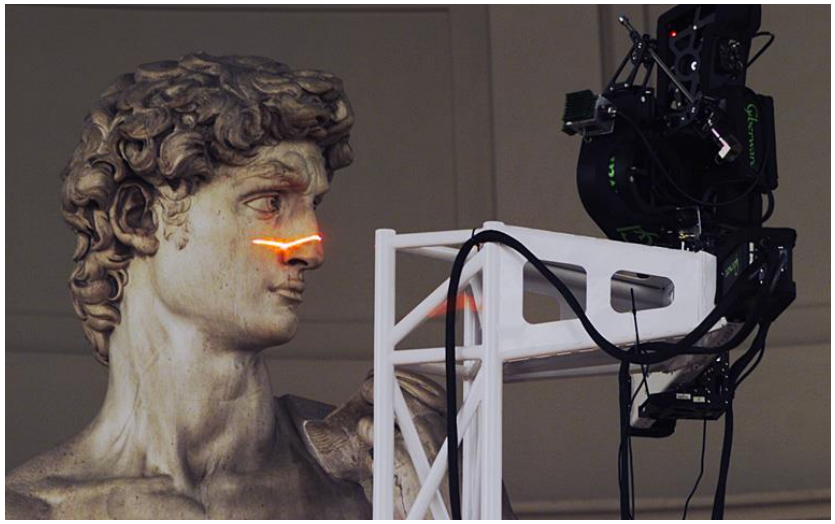
What the Kinect Sees…



3D Reconstruction with the Kinect

Laser Scanning





Digital Michelangelo Project
http://graphics.stanford.edu/projects/mich/

- **Optical triangulation**
  - Project a single stripe of laser light
  - Scan it across the surface of the object
  - This is a very precise version of structured light scanning

Slide credit: Steve Seitz

Laser Scanned Models

***The Digital Michelangelo Project*, Levoy et al.**

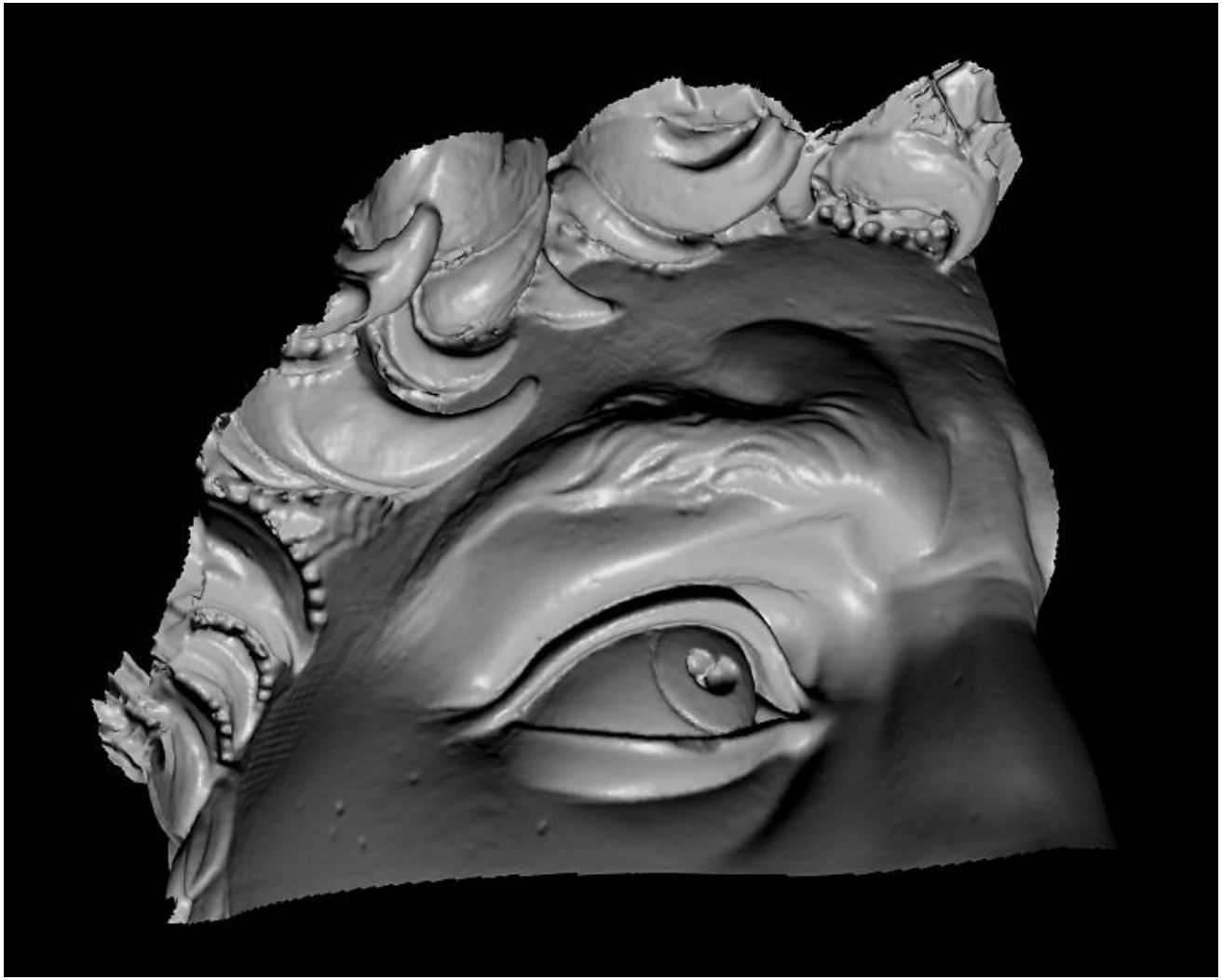
Slide credit: Steve Seitz

Laser Scanned Models

***The Digital Michelangelo Project*, Levoy et al.**

Slide credit: Steve Seitz

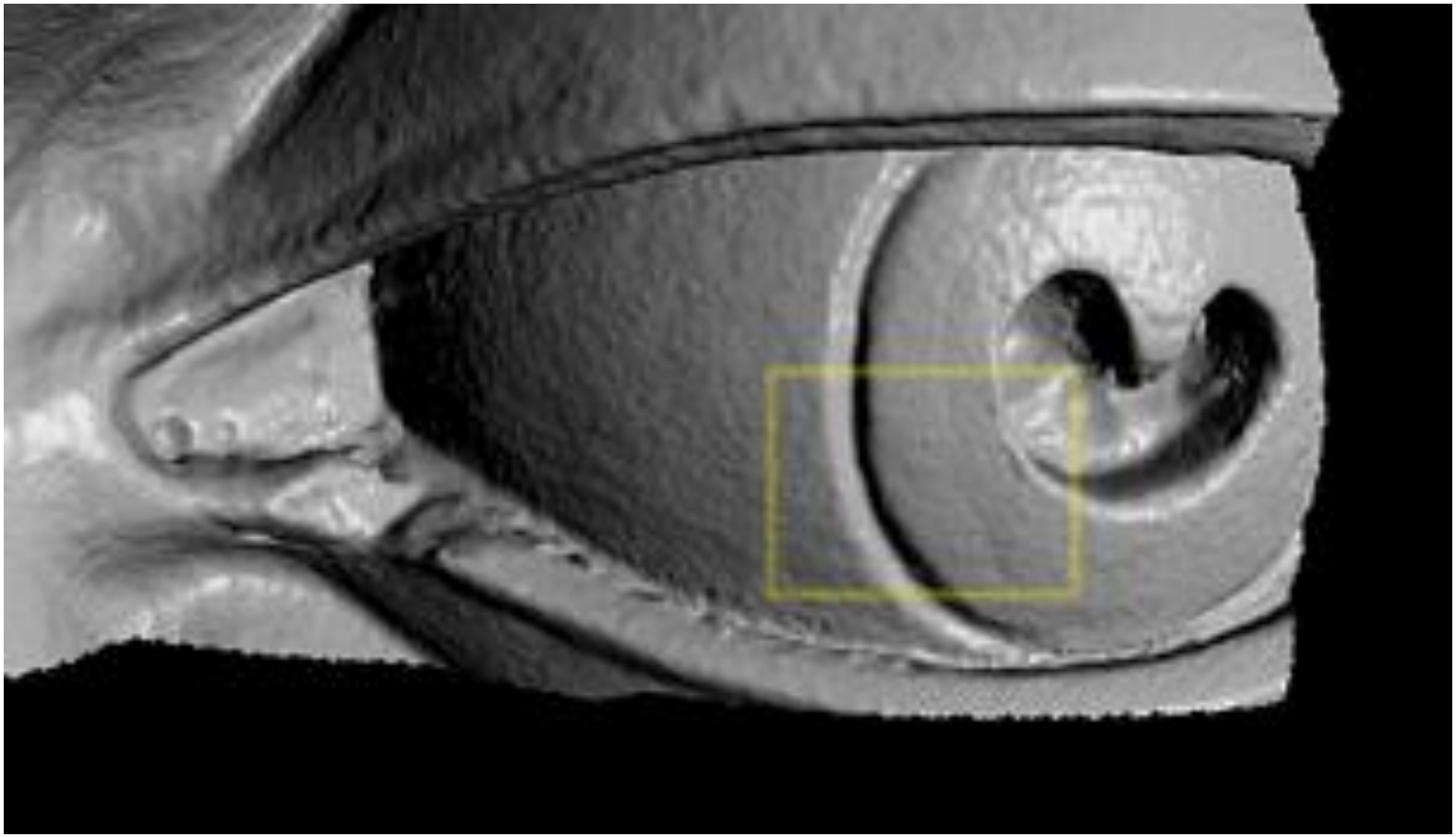
Laser Scanned Models



***The Digital Michelangelo Project*, Levoy et al.**

Slide credit: Steve Seitz

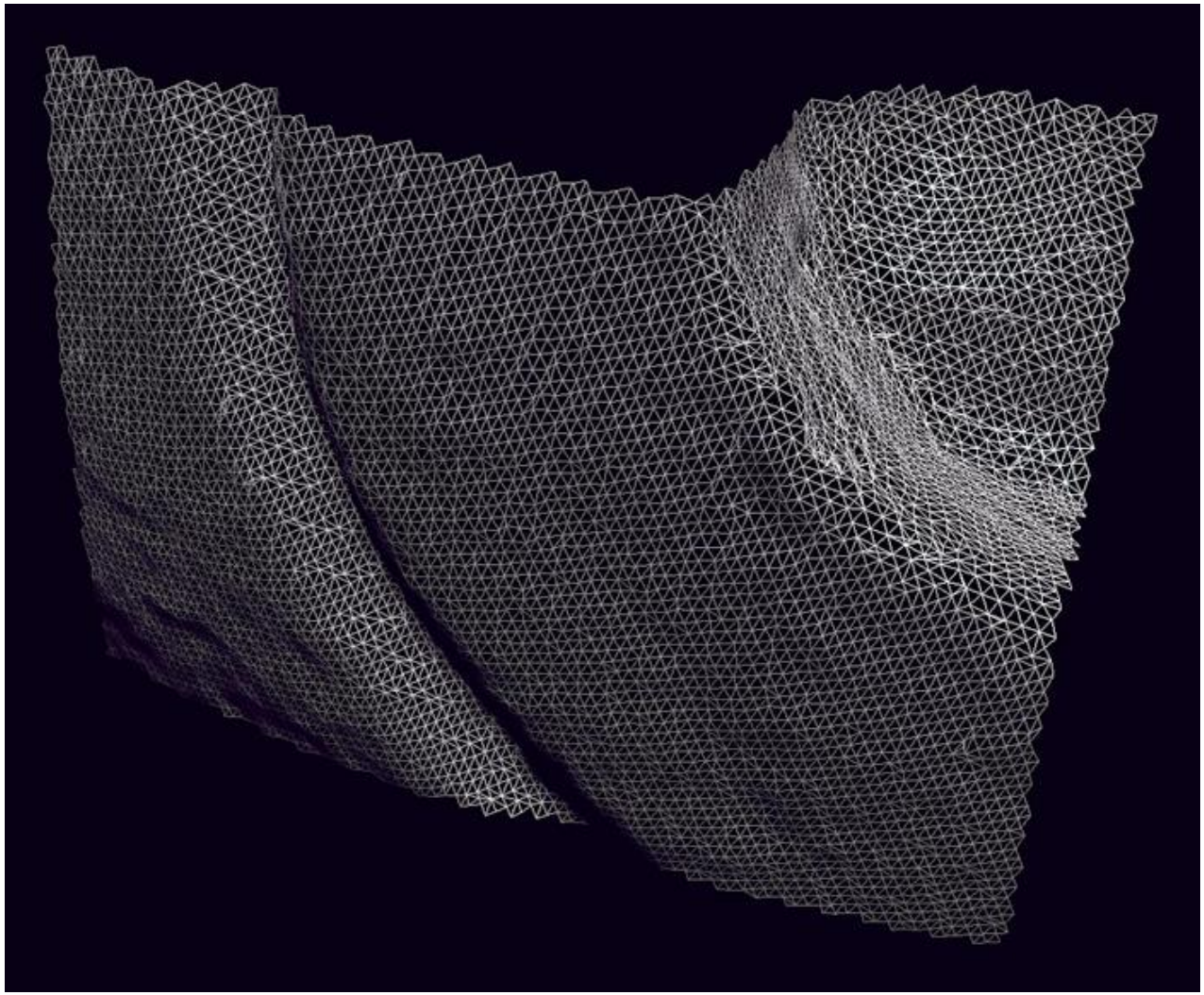
Laser Scanned Models



The Digital Michelangelo Project, Levoy et al.

Slide credit: Steve Seitz

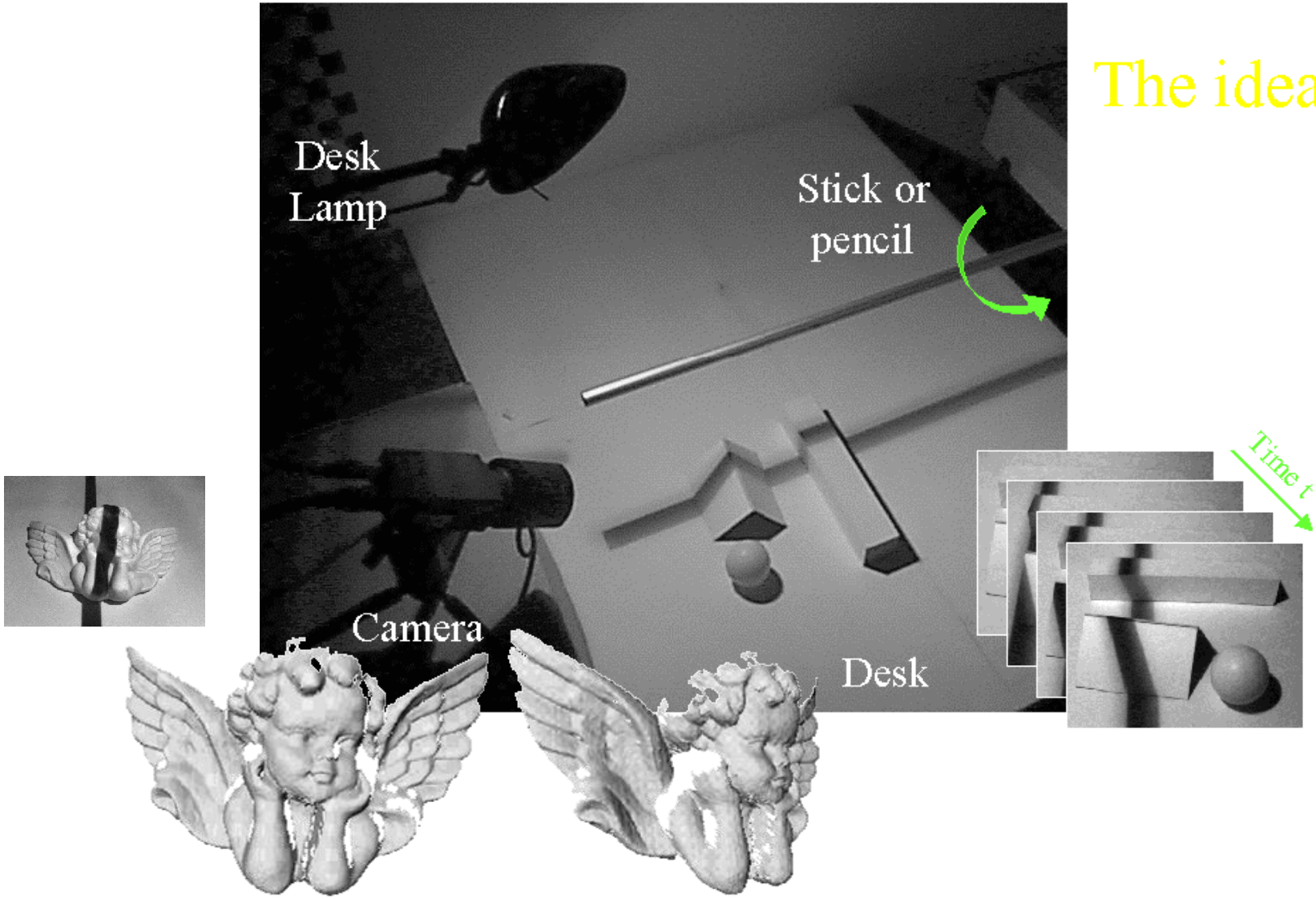
Laser Scanned Models



***The Digital Michelangelo Project*, Levoy et al.**

Slide credit: Steve Seitz

Poor Man's Scanner



Bouget and Perona, ICCV'98

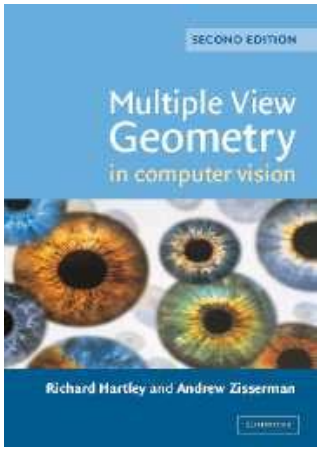
Slightly More Elaborate (But Still Cheap)

Software freely available from Robotics Institute TU Braunschweig
http://www.david-laserscanner.com/

References and Further Reading

- Background information on camera models and calibration algorithms can be found in Chapters 6 and 7 of

R. Hartley, A. Zisserman
Multiple View Geometry in Computer Vision
2nd Ed., Cambridge Univ. Press, 2004



- Also recommended: Chapter 9 of the same book on Epipolar geometry and the Fundamental Matrix and Chapter 11.1-11.6 on automatic computation of F.